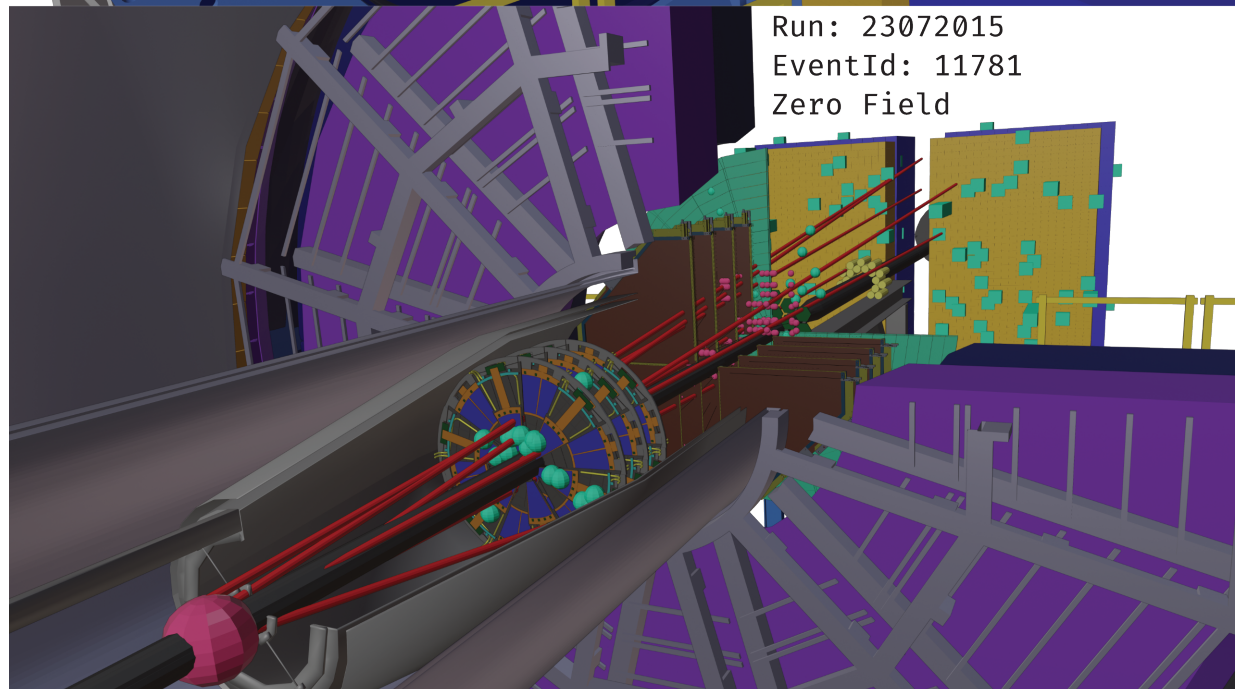
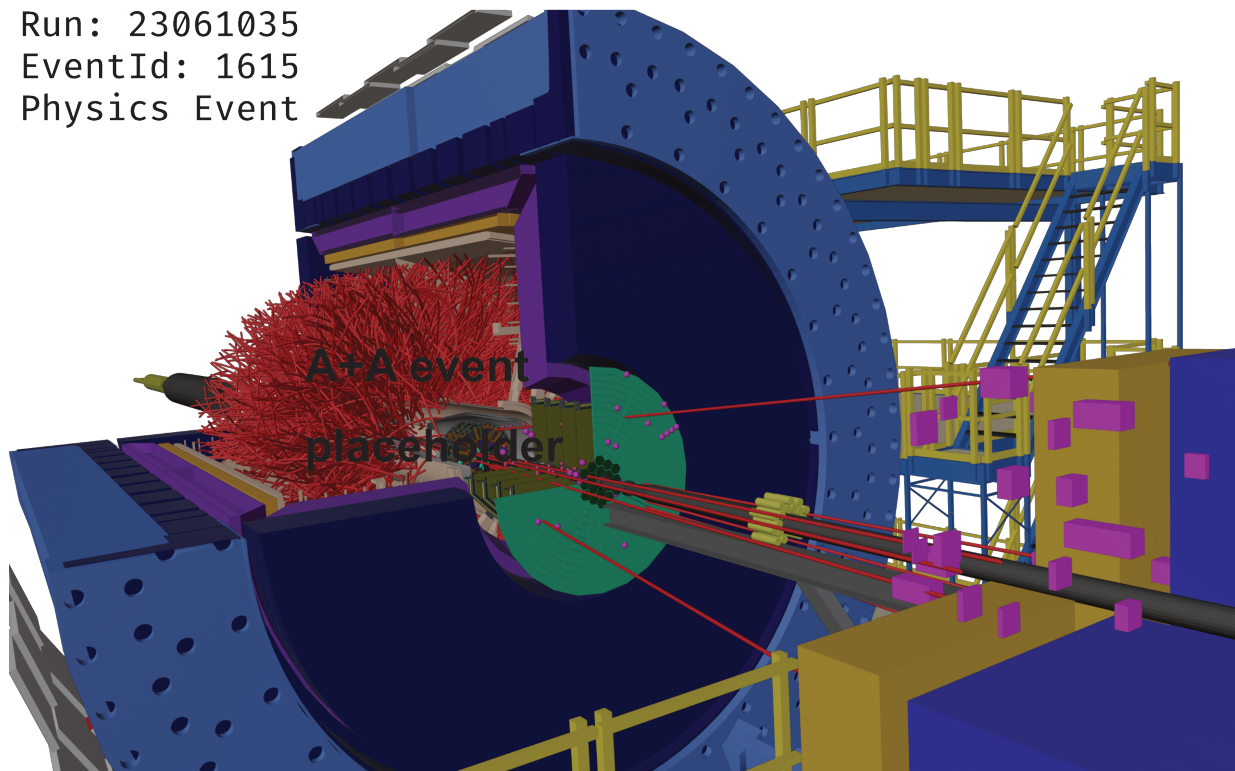


The STAR Beam Use Request for Run-23-25

The STAR Collaboration

Run: 23061035
EventId: 1615
Physics Event



1 Executive Summary

2 This Beam Use Request outlines the compelling physics programs proposed by STAR col-
3 laboration for data taking in 2023-25.

4 STAR’s **highest scientific priority** is to record a combination of high statistics soft
5 and hard probes data from Au+Au, p +Au, and p + p data at $\sqrt{s_{NN}} = 200$ GeV, during 2023-
6 25 as outlined in Table 1. When fully collected, these datasets will enable the successful
7 completion of RHIC’s scientific mission via examination of the microstructure of the Quark
8 Gluon Plasma (QGP) and a continuation of our unique forward physics program.

Table 1: Proposed Run-23 - Run-25 assuming 28 cryo-weeks of running every year, and 6 weeks set-up time to switch species in 2024. For p + p and p +Au sampled luminosities assume a “take all” trigger. For Au+Au we provide the requested event count for our minimum bias trigger, and the requested sampled luminosity from our a high- p_T trigger that covers all v_z .

$\sqrt{s_{NN}}$ (GeV)	Species	Number Events/ Sampled Luminosity	Year
200	Au+Au	20B / 40 nb ⁻¹	2023+2025
200	p + p	235 pb ⁻¹	2024
200	p +Au	1.3 pb ⁻¹	2024

9 STAR’s scientific program is enabled by the combination of the detector upgrades for
10 Beam Energy Scan phase II (BES-II) and the Forward Upgrades. In combination they gen-
11 erate STAR’s unique capabilities in particle identification (PID) over an extended rapidity
12 acceptance and down to very low transverse momentum (p_T), while maintaining a low mate-
13 rial budget. All these new detectors are now fully commissioned and operated exceptionally
14 well during Run-22.

15 Significantly increased luminosities, the extended acceptance at mid-rapidity due to the
16 iTPC, improved event plane and triggering capabilities via the EPD, and the ability to probe
17 the previously inaccessible forward region are all exploited in our Hot QCD program, that
18 informs on the microstructure of the QGP, and our Cold QCD program that will utilize
19 transverse polarization setting the stage for related future measurements at the Electron-Ion
20 Collider (EIC).

21 Combined Au+Au datasets collected in Run-23 and Run-25 will allow STAR to address
22 important questions about the inner workings of the QGP, including the temperature de-
23 pendence of the shear and bulk viscosities, the 3-D nature of the initial state, how global
24 vorticity is transferred to the spin angular momentum of particles on such short time scales
25 and the chiral properties of the medium.

26 STAR considers it critical that we collect approximately equal nucleon-nucleon luminosi-
27 ties for p + p and p +Au at 200 GeV during Run-24. This optimizes the statistical precision of
28 several critical observables that require comparisons between results in both p + p and p +Au.
29 We request transversely polarized protons for both datasets. Assuming 28 cryo-weeks in

30 Run-24 we expect to record samples that represent a factor 4.5 times the luminosity that
 31 STAR sampled during the last transversely polarized $p+p$ collisions in Run-15, and 3 times
 32 the luminosity sampled during Run-15’s transversely polarized $p+Au$ collision period.

33 As requested, we also considered the scenario that each run is reduced to 24 cryo-weeks.
 34 Under this scenario the STAR collaboration continues to request Au+Au, $p+Au$, and $p+p$
 35 running as outlined in Table 2. In this way we will take the best possible advantage of our
 36 recent upgrades. However, this scenario would result in a significant increase in both the
 37 statistical and systematic uncertainties of all the data, impacting the excellent precision we
 38 aim for with the measurements described in this BUR.

39 We estimate that 24 as opposed to 28 cryo-weeks will decrease STAR’s Au+Au data
 40 sample by at least 16%. Measurements of hard probes (jets and quarkonia), thermal di-
 41 lepton and photon-induced processes (di-lepton and J/ψ) will be most impacted since they
 42 are the most statistically demanding Hot QCD measurements proposed.

43 There is a much more significant effect on $p+p$ and $p+Au$ running due to both the
 44 6 weeks needed to change beam species, the ramp-up times, and the fact that no low-
 45 luminosity running is requested. We estimate at least a 22-25% loss in sampled $p+p$ and
 46 $p+Au$ luminosity. There will be an even larger impact on the nuclear PDFs, fragmentation
 47 functions, and gluon saturation measurements since these require comparisons of the same
 48 observables measured in both $p+p$ and $p+Au$ collisions.

Table 2: Proposed Run-23 - Run-25 assuming 24 cryo-weeks of running every year, and 6 weeks set-up time to switch species in 2024. For $p+p$ and $p+Au$ sampled luminosities assume a “take all” trigger. For Au+Au we provide the requested event count for our minimum bias trigger, and the requested sampled luminosity from our a high- p_T trigger that covers all v_z .

$\sqrt{s_{NN}}$ (GeV)	Species	Number Events/ Sampled Luminosity
200	Au+Au	17B / 34 nb ⁻¹
200	$p+p$	176 pb ⁻¹
200	$p+Au$	0.98 pb ⁻¹

49 Finally in Section 5 we propose the collection of two datasets if the opportunity arises after
 50 collection of our higher priority datasets outlined above. One proposal enables the imaging of
 51 the shape and radial profile of atomic nuclei via collective flow measurements. Such studies
 52 are important to improve our understanding of the complex initial conditions and subsequent
 53 hydrodynamical response of the medium. Information on these deformation and nuclear
 54 skin parameters are also of significant interest to the nuclear structure physics community.
 55 Heavy ion collision data have different sensitivities to nuclear structure experiments and are
 56 therefore promising complementary tools to probe different aspects of the nucleus’ shape
 57 and substructure. The other proposal expands our fixed-target program to include other
 58 light beam and target combinations. These data will help clarify the role and mechanisms
 59 of nucleon stopping. In addition, light nucleus cross sections in the target/projectile regions

60 using beams of 3-50 GeV/n are of great interest to the NASA Space Radiation Protection
61 community.

62 Contents

63	1 Highlights from the STAR Program	1
64	1.1 Highlights from the Heavy Ion Program	1
65	1.1.1 Search for the Chiral Magnetic Effect	1
66	1.1.2 Bulk Correlations	4
67	1.1.3 LFSUPC Highlights	10
68	1.1.4 Heavy-flavor Measurements	15
69	1.1.5 Jet Measurements	16
70	1.1.6 BES-II Results	21
71	1.2 Highlights from the Spin and Cold QCD Program	30
72	1.3 Run-22 Performance	36
73	1.3.1 Forward Upgrade	39
74	2 Run-23 and Run-25 Requests for Au+Au Collisions at 200 GeV	44
75	2.1 Explore the Microstructure of the QGP	44
76	2.2 Ultra-Peripheral Collisions	72
77	3 Run-24 Request for Polarized pp and $p+A$ Collisions at 200 GeV	89
78	3.1 Spin Physics with Polarized pp and $p+A$ Collisions at 200 GeV	90
79	3.2 Physics Opportunities with Unpolarized proton-Nucleus Collisions	101
80	4 Computing Resources	116
81	5 Future Opportunities	118
82	5.1 Imaging shape and radial profile of atomic nuclei via collective flow measure-	
83	ments	118
84	5.2 Fixed-target Measurements Using Light Beam and Target Combinations	125
85	6 Charge for the 2022 NPP PAC	129

1 Highlights from the STAR Program

1.1 Highlights from the Heavy Ion Program

1.1.1 Search for the Chiral Magnetic Effect

Results from the isobar blind analyses A decisive experimental test of the Chiral Magnetic Effect (CME) has become one of the major scientific goals of the heavy-ion physics program at RHIC. Isobars were collided to utilize the fact that the collisions of ruthenium produce larger magnetic fields than those of zirconium by 5–9%, hence a 10–18% larger CME correlation signal because of its B^2 dependence. Therefore, the CME would cause the ratio of CME-sensitive observables in Ru+Ru over Zr+Zr to be greater than one, assuming that backgrounds are the same in the two systems. The isobar run was specially designed to reduce the systematics in this ratio. In order to minimize unconscious and pre-determined biases a blind analysis was performed with pre-defined criteria on what would constitute observation of a CME signal. For example, the double ratio of the primary CME-sensitive correlator $\Delta\gamma$ scaled by ellipticity v_2 in ruthenium over zirconium is expected to be greater than one if there is a non-zero CME fraction.

The measurements of the double ratio of $\Delta\gamma/v_2$ with various kinematic cuts from the isobar blind analysis are shown in Fig. 1. A precision in our measurement down to 0.4% was observed in the measurement of the $\Delta\gamma/v_2$ ratio. However, no predefined signature of CME was observed. The observation that the double ratio of $\Delta\gamma/v_2$ is significantly below unity can be attributed to the multiplicity difference between Ru+Ru and Zr+Zr as shown by the ratio of the inverse of uncorrected tracks $1/N_{\text{trk}}^{\text{offline}}$ measured within the acceptance of $|\eta| < 0.5$. This ratio being less than one is explainable not by larger charge separation in Zr+Zr compared to Ru+Ru, but rather by larger multiplicity dilution ($\propto 1/N_{\text{trk}}^{\text{offline}}$) in Ru+Ru. This argument is further demonstrated by the ratio of a similar quantity $r(m_{\text{inv}})$, which measures the relative pair multiplicity difference opposite-sign and same-sign pion pairs; in a model in which the background for $\Delta\gamma$ is solely due to flowing clusters, $\Delta\gamma/v_2$ would scale simply as r .

A number of other CME sensitive observables were also measured, such as the factorization coefficients κ_{112}, k_2 , the inverse width of the R -variable as shown in Fig. 1. The ratios of these observables in Ru+Ru over Zr+Zr are also found to be less than unity, again not consistent with pre-defined CME signatures. In addition, CME-insensitive charge separation measures using third harmonic event planes such as $\Delta\gamma_{123}/v_3$ and k_3 were also measured to provide data-driven baselines. The utility of these baselines are not affected by multiplicity dilution although their constraining powers are limited by their larger uncertainties as compared to the equivalent observables involving second harmonics.

Non-flow effects on the isobar baseline The overall conclusion from the blind analysis is that no predefined CME signature has been observed in the isobar data. However, to extract a quantitative result utilizing the full sensitivity of the isobar run, careful consideration must be given to the baseline; the baseline of unity is expected to be affected by the

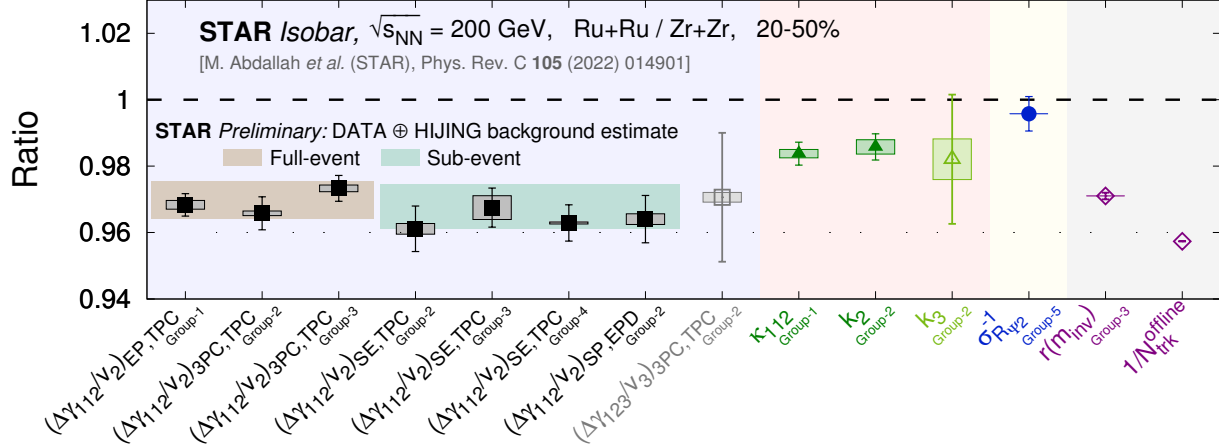


Figure 1: Compilation of results from the blind analysis. Results are shown in terms of the ratio of measures in Ru+Ru collisions divided by Zr+Zr collisions. Solid dark symbols denote CME-sensitive measures whereas open light symbols show counterpart measures that are designed to be insensitive to CME. The vertical lines indicate statistical uncertainties whereas boxes indicate systematic uncertainties. The colors in the background are intended to separate different types of measures. The two data points (open markers) have been added on the right to indicate the ratio of inverse multiplicities ($N_{\text{trk}}^{\text{offline}}$) and the ratio of relative pair multiplicity difference (r) as explained in the text. The two bands show estimates for background calculated using isobar data and the HIJING model incorporating the multiplicity difference between the two isobars and non-flow effects.

125 multiplicity difference between the two isobars. At the last quark matter conference (QM
 126 2022), the STAR collaboration presented important progress toward quantifying possible
 127 remaining CME signals by incorporating the multiplicity difference between the two isobars
 128 and non-flow effects which are also different between the Zr and Ru. As a first step, the
 129 estimates are made for the background contribution to the double ratio of the $\Delta\gamma/v_2$ by
 130 incorporating: 1) the difference in the multiplicity dilution ($\propto 1/N_{\text{trk}}^{\text{offline}}$) between the two
 131 isobars, 2) data-driven estimates of various sources of two-particle non-flow correlations and,
 132 3) sources of three-particle non-flow correlations estimated using a HIJING simulation. The
 133 background estimates for two difference kinematic regions involving full TPC acceptance
 134 (Full-event) and TPC acceptance with two sub-events (Sub-event) are shown by bands with
 135 different colors in Fig. 1. The conclusion is that the measurements of $\Delta\gamma/v_2$ from isobars
 136 are consistent with our preliminary estimate for background expectations.

137 **CME measurement in Au+Au collisions** The most recent measurement of charge
 138 separation in Au+Au collisions was performed with the spectator plane (SP) and participant
 139 plane (PP) using a recently developed method. [1, 2] The idea is straightforward: the CME
 140 signal is sensitive to the magnetic field which is primarily generated by spectator protons,
 141 so the signal is the strongest in the measurement made with respect to SP; on the other
 142 hand, flow is strongest along the direction of PP, so is the flow-induced background for

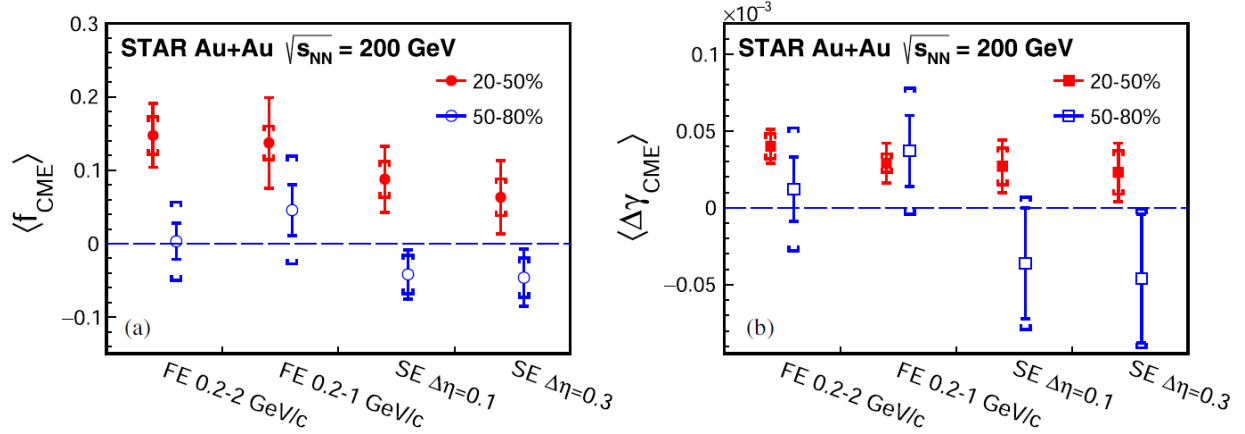


Figure 2: The flow-background removed f_{CME} (a) and $\Delta\gamma_{CME}$ (b) signal in 50%–80% (open markers) and 20%–50% (solid markers) centrality Au+Au collisions at $\sqrt{s_{NN}}=200$ GeV, extracted by various analysis methods [full-event (FE), sub-event (SE)] and kinematic cuts. Error bars show statistical uncertainties; the caps indicate the systematic uncertainties.

143 the CME. From the charge correlation measurements with respect to SP and PP, one can
 144 resolve the CME signal and the flow-induced background. Figure 2 shows the CME signal
 145 fraction (f_{CME}) in the inclusive $\Delta\gamma$ measurement via this SP/PP method. [3] An indication
 146 of a positive CME signal is seen in mid-central 20–50% central Au+Au collisions, while the
 147 signal is consistent with zero in more peripheral collisions. The significance of the CME
 148 signal is on the order of 2σ .

149 Since the v_2 measurement and the 3-particle correlator measurement with respect to PP
 150 are contaminated by non-flow effects, the measured f_{CME} is still affected by non-flow. [3,
 151 4] Unlike isobar collisions where non-flow affects both measurements, non-flow in Au+Au
 152 collisions affects only the PP measurements, thus is relatively easier to estimate. Model
 153 studies together with non-flow data measurements [5] suggest that non-flow effects on f_{CME}
 154 may be small or even negative. [4] This makes the measured positive f_{CME} , although with
 155 large uncertainties, intriguing. It is noteworthy that the non-observation of the CME in
 156 isobar collisions (~ 4 billion MB events) and a hint of a positive CME signal in Au+Au
 157 collisions (~ 2.4 billion MB events) are not contradictory. It was recently realized, based
 158 on mundane physics, that the CME signal to background ratio in isobar collisions can be a
 159 factor of 3 smaller than in Au+Au collisions. [6]

160 **CME measurements with the BES-II data** One important question regarding the
 161 CME is: What happens at lower collision energies? In this context a new idea has emerged.
 162 The newly installed event-plane detector (EPD) upgrade provides a new capability at STAR
 163 towards the CME search at lower collision energy and for the BES-II program. [7]

164 The first idea is simple, at lower energies the EPD acceptance ($2.1 < |\eta| < 5.1$) falls
 165 in the region of beam rapidity (Y_{beam}) and can measure the plane of strong directed flow
 166 (Ψ_1) of spectator protons, beam fragments and stopped protons, which is therefore strongly

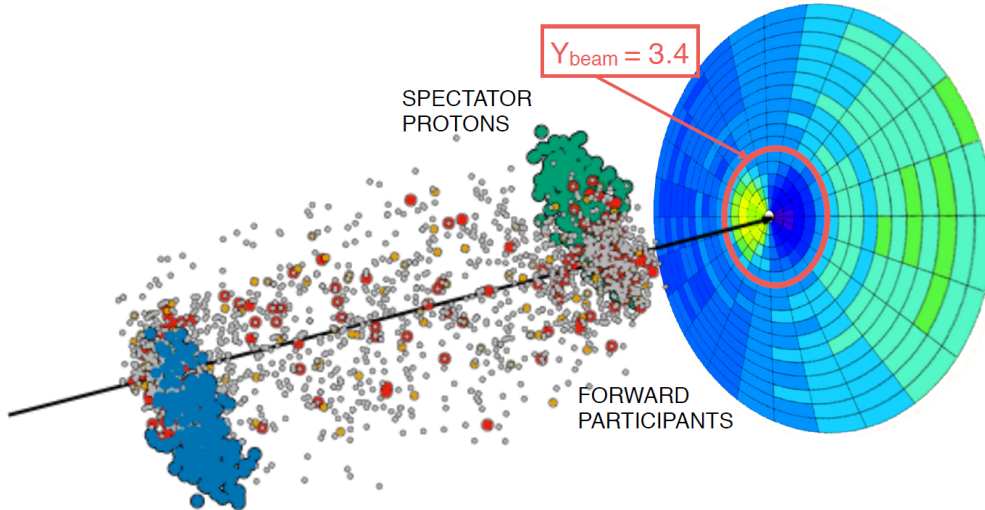


Figure 3: Prospect for the CME search with the BES-II data. (Left) Single simulated UrQMD event and EPD detector acceptance that covers beam rapidity and detects both forward participants and spectators in 27 GeV Au+Au collisions that have large directed flow which changes sign at $\eta = Y_{\text{beam}} = 3.4$. (Right) γ -correlators scaled by v_2 across different event-planes and double ratio of spectator/participant event plane results which would be above unity for finite CME scenario.

167 correlated to the B-field direction (see Fig. 3). The next step is to measure $\Delta\gamma$ with respect
 168 to Ψ_1 and compare it with the measurement of $\Delta\gamma$ along the Ψ_2 planes determined from the
 169 outer regions of EPD and the TPC at mid-rapidity that are relatively more weakly correlated
 170 to the B-field direction. A test of the CME scenario will be to see if a large difference is
 171 observed in the measurements. First preliminary measurements from STAR as shown in
 172 Fig. 3 are dominated by uncertainty, but seem to show good prospects for the CME search
 173 at lower energies. With the higher statistics data from the BES-II collider data (7.7-19.6
 174 GeV) and fixed target program more precise measurements are possible.

175 1.1.2 Bulk Correlations

176 Over the past years, the STAR collaboration has performed a series of correlation measure-
 177 ments directed towards a comprehensive understanding of the QCD phase diagram and the
 178 transport properties of the QGP phase. Here we highlight the most recent STAR results on
 179 bulk correlations.

180 **Global spin polarization and alignment** Non-central heavy ion collisions can generate
 181 a large orbital angular momentum (OAM) in the system. Part of OAM is transferred to
 182 the system in the form of preferential alignment of the intrinsic angular momentum (spin)
 183 of particles along the OAM direction through spin-orbit couplings, a phenomenon called
 184 global polarization. [8, 9] The global polarization of quarks influences vector mesons such as
 185 $\phi(1020)$ and $K^{*0}(892)$. The spin state of a vector meson can be described by a 3×3 spin

186 density matrix with unit trace. [10] The diagonal elements of this matrix, namely, ρ_{11}, ρ_{00}
 187 and ρ_{-1-1} , are probabilities for the spin component along a quantization axis to take the
 188 values of 1, 0, and -1 respectively. The quantization axis is a chosen axis onto which the
 189 projection of angular momentum has well-defined quantum numbers. When the three spin
 190 states have equal probability to be occupied, all three elements are $1/3$ and there is no spin
 191 alignment. If $\rho_{00} \neq 1/3$, the probabilities of the three spin states along the quantization axis
 are different and there is a spin alignment.

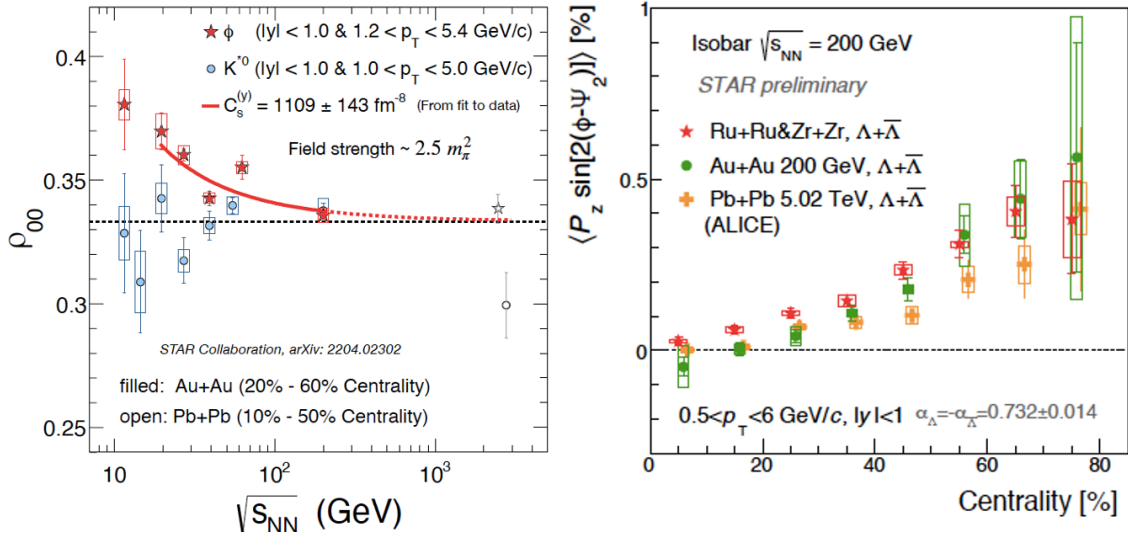


Figure 4: Left panel: Global spin alignment measurement of ϕ and K^{*0} vector mesons in Au+Au (Pb+Pb) collisions. The measured matrix element ρ_{00} is plotted as a function of beam energy for the ϕ and K^{*0} vector mesons within the indicated windows of centrality, transverse momentum p_T , and rapidity y . The two points on the right (Pb+Pb collisions at 2.76 TeV) are integrated over the ALICE collaboration results [?], with a p_T range of 1.0–5.0 GeV/c for ϕ and K^{*0} . The red solid curve is a fit to data in the range of $\sqrt{s_{NN}} = 19.6$ to 200 GeV, based on a theoretical calculation with a ϕ -meson field. [11] Right panel: Λ spin polarization along beam direction in isobaric collisions at 200 GeV, Au+Au at 200 GeV, and Pb+Pb at 5.02 TeV.

192

193 Hyperons are natural candidates to explore global spin polarization since in the parity
 194 violating weak decays of hyperons the momentum vector of the decay baryon is highly
 195 correlated with the hyperon spin. The first observation of positive polarization of Λ hyperons
 196 in the Beam Energy Scan-I provided evidence for the creation of the most vortical fluid
 197 ever observed. [12] In non-central collisions strong anisotropic flow can generate a non-zero
 198 vorticity along the beam axis. The vorticity and the corresponding polarization exhibits
 199 a quadrupole structure in the transverse plane. This polarization is characterized by the
 200 second harmonic sine component in the Fourier decomposition of the polarization along the
 201 beam axis (P_z). The P_z for Λ hyperons was measured by STAR and was found to have
 202 opposite sign compared to the hydrodynamic and transport model calculations, known as
 203 "spin puzzle". The introduction of shear induced polarization can reproduce the sign of P_z

204 indicating that it is sensitive to the hydrodynamic gradients as well as the dynamics of the
 205 spin degrees of freedom.

206 Figure 4 presents the ρ_{00} for ϕ and K^{*0} vector mesons in Au+Au collisions at beam
 207 energies between $\sqrt{s_{\text{NN}}} = 11.5$ and 200 GeV. The ϕ -meson results are presented for transverse
 208 momentum $1.2 < p_T < 5.4$ GeV/ c ; ρ_{00} for this species is significantly above 1/3 for collision
 209 energies of 62 GeV and below, indicating finite global spin alignment. The ρ_{00} for ϕ mesons,
 210 integrated over beam energies of 62 GeV and below, is 0.3541 ± 0.0017 (stat.) ± 0.0018
 211 (sys.); this is a significance of 8.4 σ for the ϕ -meson ρ_{00} to be above 1/3. Figure 4 also
 212 presents the beam-energy dependence of ρ_{00} for K^{*0} within $1.0 < p_T < 5.0$ GeV/ c . We
 213 observe that ρ_{00} for K^{*0} is largely consistent with 1/3, in marked contrast to the case for ϕ .
 214 The surprisingly large positive deviation for ϕ meson in mid central collisions is consistent
 215 with a model which introduce polarization by a strong force field of vector meson.

216 Figure 4 right panel shows the Λ polarization along beam direction (P_z) as a function of
 217 centrality in isobaric collisions at 200 GeV, Au+Au at 200 GeV, and Pb+Pb at 5.02 TeV.
 218 The amplitude of the sine modulation tends to increase from central to peripheral collisions.
 219 The results hint at a colliding system size dependence rather than beam energy dependence.

220 Measurements sensitive to the initial state

221 Beam-energy dependence of anisotropic flow fluctuations and correlations

222 The multi-particle flow harmonics $v_n\{k\}$, for $k=2, 4,$ and 6, obtained via multi-particle
 223 correlation methods [13,14] can give direct access to the event-by-event flow fluctuations. [15,
 224 16] Also the flow-plane decorrelations (measured by $r_n(\eta)$) that are driven by the eccentricity
 225 decorrelations [17,18] are expected to be caused by (i) the effect of the initial state torque [19,
 226 20], and (ii) hydrodynamic fluctuations [21] and expected to give information to the event-by-
 227 event flow fluctuations. In addition, correlations between the average transverse radial flow
 228 ($[p_T]$) and the v_n coefficients ($\rho(v_n^2, [p_T])$) could encode crucial information on the correlation
 229 between the size and the eccentricities in the initial state, and on the correlations of the
 230 strength of the hydrodynamic response with the flow coefficients. The ($\rho(v_n^2, [p_T])$) is given
 231 by [22–27]

$$\rho(v_n^2, [p_T]) = \frac{\text{cov}(v_n^2, [p_T])}{\sqrt{\text{Var}(v_n^2)}\sqrt{\text{Var}([p_T])}}. \quad (1)$$

232 Consequently, extensive measurements of $v_n\{k\}$ and $\rho(v_n^2, [p_T])$ for different beam energies
 233 could help to disentangle the fluctuation and correlation contributions from their respec-
 234 tive sources, as well as establish whether flow fluctuations and correlations depend on the
 235 temperature, T , baryon chemical potential, μ_B , or both. It could also provide unique sup-
 236 plemental constraints to distinguish between different initial-state models and reduce the
 237 fluctuations-related uncertainties associated with the extraction of $\eta/s(T, \mu_B)$.

238 Figure 5 provides a summary of the centrality dependence of $v_2\{2\}$ (a), $v_2\{4\}$ (b),
 239 $v_2\{6\}$ (c) and the ratio $v_2\{4\}/v_2\{2\}$ (d) for the respective beam energies as indicated. The
 240 $v_2\{4\}/v_2\{2\}$ ratios shown in Fig. 5 (d) suggest that within the given uncertainties, the flow

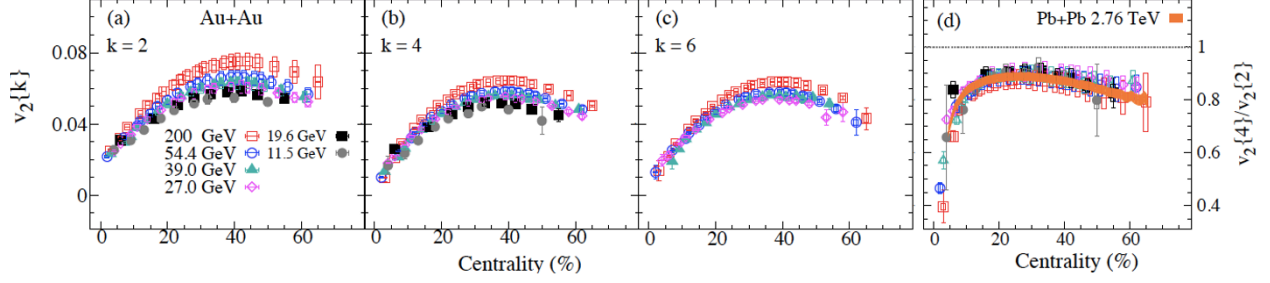


Figure 5: Comparison of the centrality dependence of the charged hadrons $v_2\{2\}$ (a), $v_2\{4\}$ (b), $v_2\{6\}$ (c), and the ratio $v_2\{4\}/v_2\{2\}$ (d), in the p_T range $0.2 - 4.0$ GeV/ c for Au+Au collisions at $\sqrt{s_{NN}} = 11.5-200$ GeV. The vertical lines and the open boxes indicate the respective statistical and systematic uncertainties. The shaded band in (d) indicates the ratios obtained from the LHC measurements for the p_T range $0.2 - 3.0$ GeV/ c for Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV [28].

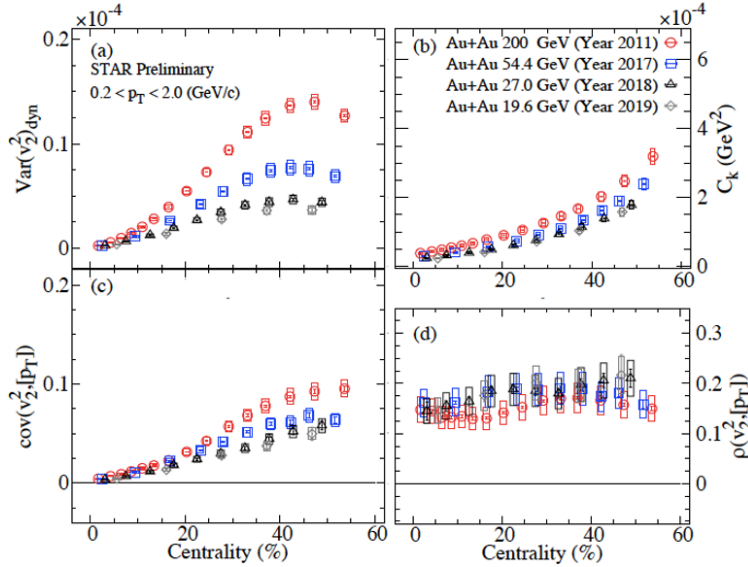


Figure 6: Comparison of the centrality dependence of the values for $\text{Var}(v_n^2)_{dyn}$ (a), c_k (b), $\text{cov}(v_n^2, [p_T])$ (c), and $\rho(v_n^2, [p_T])$ (d), measured for Au+Au collisions at $\sqrt{s_{NN}} = 200, 54.4, 27$ and 19.6 GeV.

241 fluctuations are weakly dependent on the beam energy, if at all, irrespective of the collision
 242 centrality. The magnitude and trend of these ratios are also comparable to those for the
 243 LHC measurements for Pb+Pb collisions at $\sqrt{s_{NN}} = 2.76$ TeV [28] and to the $\varepsilon_2\{4\}/\varepsilon_2\{2\}$
 244 ratios, in central to mid-central collisions, shown in Fig. 5 (b). These results suggest that
 245 the flow fluctuations associated with the expansion dynamics do not change substantially
 246 over the beam energy range $\sqrt{s_{NN}} = 11.5-2760$ GeV.

247 Figure 6 shows the beam-energy dependence of $\text{Var}(v_n^2)$ (a), c_k (b), $\text{cov}(v_n^2, [p_T])$ (c), and
 248 $\rho(v_n^2, [p_T])$ (d). They indicate patterns which depend on beam energy. These results suggest
 249 that the beam energy dependence of $\text{Var}(v_n^2)$ and $\text{cov}(v_n^2, [p_T])$ could provide important con-

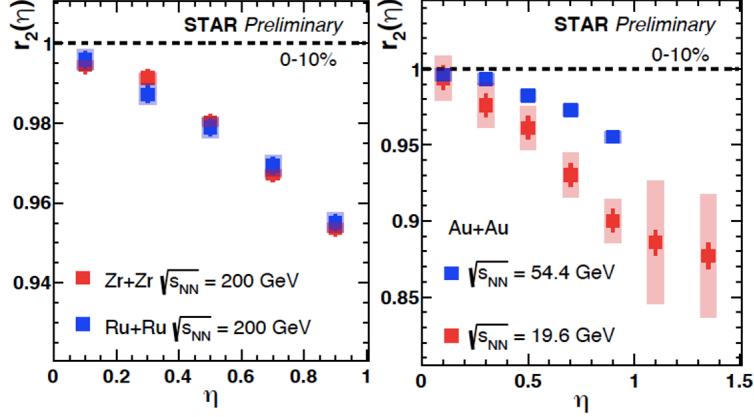


Figure 7: Comparison of the η dependence of the values for $r_n(\eta)$ for Au+Au collisions at $\sqrt{s_{\text{NN}}} = 54.4$ and 19.6 GeV.

250 constraints for η/s while the measurements for $\rho(v_n^2, [p_T])$ provide complimentary constraints
 251 for the initial-state eccentricity and its fluctuations.

252 Figure 7 shows the beam-energy dependence of $r_n(\eta)$. They indicate patterns and values
 253 which depend on beam energy.

254 Nuclear deformation and neutron skin thickness measurements

255 Nuclear deformation and neutron skin thickness are fundamental properties of atomic
 256 nuclei that reflect the correlated nature of the dynamics of nucleons within a quantum many-
 257 body system. The majority of atomic nuclei possess an intrinsic deformation, most of which
 258 are an axial quadrupole, or ellipsoidal, deformation. Prior relativistic heavy-ion collision
 259 measurements from STAR reported strong signatures of nuclear deformation using detailed
 260 comparisons between Au+Au collisions and U+U collisions. [29] These measurements suggest
 261 that U+U collisions are much more deformed in their ground state. Consequently, we can
 262 say that detailed comparisons between different nuclei enabled us to examine the geometry
 263 of the colliding ions. [30–32]

264 Recently we analyzed the Ru+Ru and Zr+Zr collision data and found that they could be
 265 used to study the nuclear deformation [33, 34] as well as the neutron skin thickness. [35–37]
 266 Figure 8 shows the $N_{\text{trk}}^{\text{offline}}$, v_2 , v_3 , and $\langle \delta p_T^2 \rangle / \langle p_T \rangle^2$ ratios between the isobar systems. All
 267 of them show non-monotonic centrality dependencies similar in shape to the theoretical
 268 prediction [33] that include effects of neutron skin as well as deformation parameters β_2 and
 269 β_3 . Figure 9 shows the centrality dependence of the Ru+Ru/Zr+Zr ratio of $\langle p_T \rangle$ compared
 270 to the theoretical expectations. [37] It is shown that this ratio increases with the symmetry
 271 energy slope parameter $L(\rho_c)$ because the neutron skin effect, larger in Zr than in Ru,
 272 increases with $L(\rho_c)$. Such an effect is non-trivial and can reach as much as 0.5%. The data
 273 model comparison should help constrain the symmetry energy slope parameter and the β_2
 274 and β_3 deformation parameters.

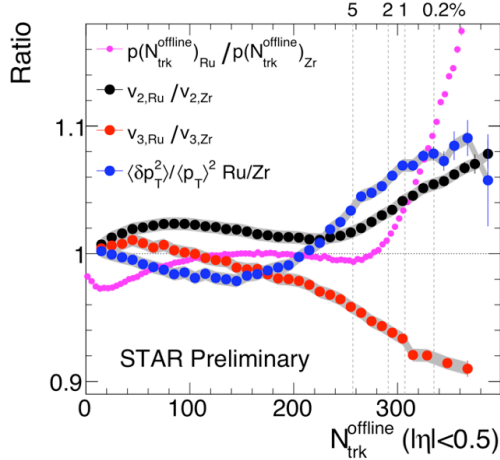


Figure 8: The N_{ch} dependence of the Ru+Ru/Zr+Zr ratio of $N_{trk}^{offline}$, v_2 , v_3 , and $\langle \delta p_T^2 \rangle / \langle p_T \rangle^2$.

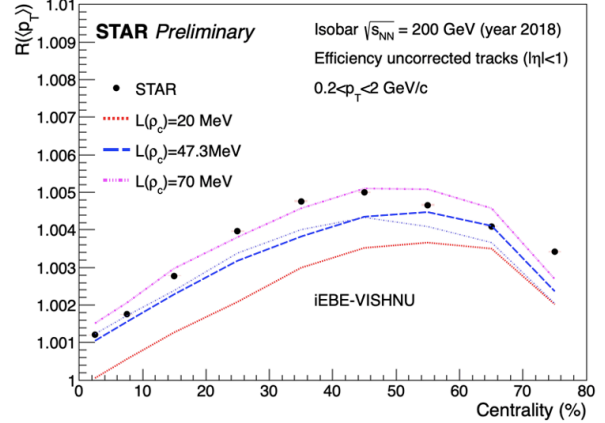


Figure 9: The centrality dependence of the Ru+Ru/Zr+Zr ratio of $\langle p_T \rangle$. The lines represent the theoretical predictions. [37]

Azimuthal anisotropy measurements of identified hadrons

Stronger constraints on transport and hydrodynamic model simulations can be achieved via investigating the azimuthal anisotropy of identified particles as a function of transverse momentum and collision centrality. Also, one can understand the initial conditions in heavy-ion collisions via varying the collision system size.

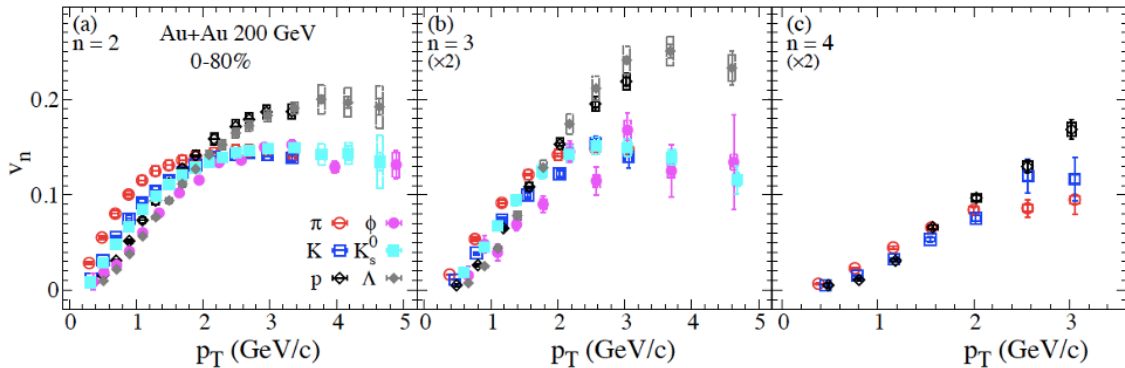


Figure 10: The transverse momentum dependence of the identified particle v_2 (a), v_3 (b), and v_4 (c) for 0–80% central Au+Au collisions at 200 GeV.

Recently we reported the results on flow coefficients of v_2 (a) and v_3 (b) of π , K , p , Λ , φ and K_s^0 and v_4 (c) of π , K and p for 0–80% central Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The measurements indicate similar increasing then flattening trends as a function of p_T in $v_{n=2,3,4}(p_T)$ for all particles shown. Also mass ordering at low p_T is observed for v_2 , v_3 , and v_4 . The shapes of the flow harmonics for light and strange mesons are comparable, which suggests similar flow strength for u, d, and s quarks.

286 Charge dependent directed flow

287 In non-central heavy-ion collisions, the charged particles in the approaching nuclei can
288 generate a substantial magnetic field. Theoretical calculations predicted that the magnetic
289 field is large ($B \sim 10^{18}$ Gauss) but short lived. As noted above the presence of such a strong
290 magnetic field can lead to novel QCD phenomena such as CME and CMW. To understand
291 the Chiral phenomena, it is of utmost important to understand the initial magnetic field
292 that could drives the charge separation. It was first proposed in [38] that the initial B-field
293 can induce a measurable effect in the form of a charge-odd contribution to the directed flow
294 coefficient (v_1). Experimental attempts have been made by STAR and ALICE by measuring
295 charge dependent Δv_1 for D^0 , \bar{D}^0 and inclusive charged hadron species, but the statistical
296 significance of those measurements are marginal.

297 Recently, STAR reported a striking centrality dependence of the v_1 slope difference
298 ($\Delta dv_1/dy$) of protons and anti-protons. The left panel of Fig. 11 presents centrality de-
299 pendence of $\Delta dv_1/dy$ between proton and anti-proton in 200 GeV Au+Au and isobar colli-
300 sions. It is observed that the $\Delta dv_1/dy$ changes sign from positive to negative from central
301 to peripheral collisions. While the positive $\Delta dv_1/dy$ is consistent with expectation from
302 transported quarks, the negative sign (with a significance of $\sim 5\sigma$) is qualitatively consistent
303 with expectation from electromagnetic field induced effects, and can be explained by the
304 dominance of the Faraday/Coulomb effect [38].

305 STAR also followed another novel approach to probe the electromagnetic fields by utiliz-
306 ing the hadrons with constituent quarks (K^- , \bar{p} , $\bar{\Lambda}$, ϕ , Ξ and Ω) that are produced in collisions,
307 which avoids contributions from transported quarks. Under the assumptions of quark co-
308 alescence, $\Delta dv_1/dy$ is studied for various pairs of particle combinations corresponding to
309 varying electric charge difference (Δq) and strangeness difference (ΔS). It is observed that
310 the $\Delta dv_1/dy$ increases with Δq and ΔS and the increase is stronger for 27 GeV than for 200
311 GeV Au+Au collisions. The right panel of Fig. 11 presents $\Delta dv_1/dy$ as function of ΔS for 10-
312 40% Au+Au collisions at 27 and 200 GeV. It is found that the PHSD calculations including
313 electromagnetic fields can describe the charge-dependent splitting within uncertainties.

314 1.1.3 LFSUPC Highlights

315 The Light-flavor Spectra and Ultra-peripheral Collisions (LFSUPC) Physics Working Group
316 (PWG) divides its efforts along six different lines of analysis: Light-charged particle (π , K,
317 p) spectra identified through dE/dx and time-of-flight (TOF) information, strange-hadron
318 spectra identified through the secondary vertex decay topology, light-nuclei spectra identified
319 through dE/dx and TOF, hypernuclei identified through decay topology, di-lepton produc-
320 tion, and ultra-peripheral collisions. Analysis efforts on the first five topics have focused on
321 newly reconstructed/processed BES-II/FXT datasets (including two articles submitted for
322 publication [], and five talks at Quark Matter 2022 []) and the submitted results be reviewed
323 in section! 1.1.6.

324 A linearly polarized photon can be quantized from the Lorentz-boosted electromagnetic
325 field of a nucleus traveling at ultra-relativistic speed. By utilizing this source of polarized
326 photons, STAR is experimentally investigating the Breit-Wheeler process through the mea-

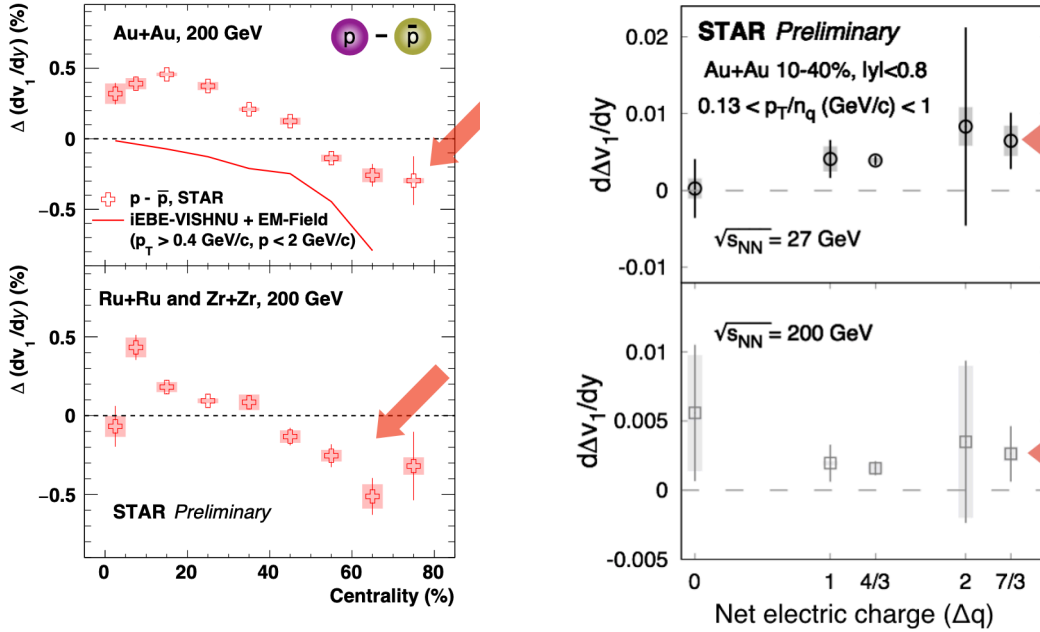


Figure 11: Left: $\Delta dv_1/dy$ as a function of centrality between proton and anti-protons in 200 GeV Au+Au and isobar (Ru+Ru and Zr+Zr) collisions. Right: $\Delta dv_1/dy$ as a function of electric charge difference (Δq) in 10–40% Au+Au collisions at 27 and 200 GeV.

327 surement of electron-positron pairs in ultraperipheral Au+Au collisions at $\sqrt{s_{NN}} = 200$
 328 GeV. [39] The measurements reveal a large fourth-order angular modulation (a $\cos 4\Delta\phi$, as
 329 seen in fig. 12) in the angle (ϕ) between the transverse momentum of the pair and the
 330 transverse momentum of one of its daughters. The differential cross section as a function of
 331 e^+e^- pair transverse momentum P_\perp peaks at low values (~ 30 MeV/c) and displays a sig-
 332 nificant centrality dependence. These features are consistent with QED calculations for the
 333 collision of linearly polarized photons quantized from the extremely strong electromagnetic
 334 fields generated by the highly charged Au nuclei at ultrarelativistic speed. The experimental
 335 results have implications for vacuum birefringence and for mapping the magnetic field which
 336 is important for emergent QCD phenomena.

337 When two relativistic heavy nuclei pass one another at a distance of a few nuclear radii,
 338 the photon from one nucleus may interact through a virtual quark-antiquark pair with glu-
 339 ons from the other nucleus forming a short-lived vector meson (e.g. ρ^0). STAR has studied
 340 diffractive photoproduction in Au+Au and U+U ultraperipheral collisions. [40] The polar-
 341 ization was utilized to observe a unique spin interference pattern in the angular distribution
 342 of $\rho \rightarrow \pi^+\pi^-$ decays as seen in fig. 13. The observed interference is a result of an overlap
 343 of two wave functions at a distance an order of magnitude larger than the ρ^0 travel distance
 344 within its lifetime. The strong-interaction nuclear radii were extracted from these diffractive
 345 interactions (fig. 13 right panel), and found to be 6.53 ± 0.06 fm (^{197}Au) and 7.29 ± 0.08 fm

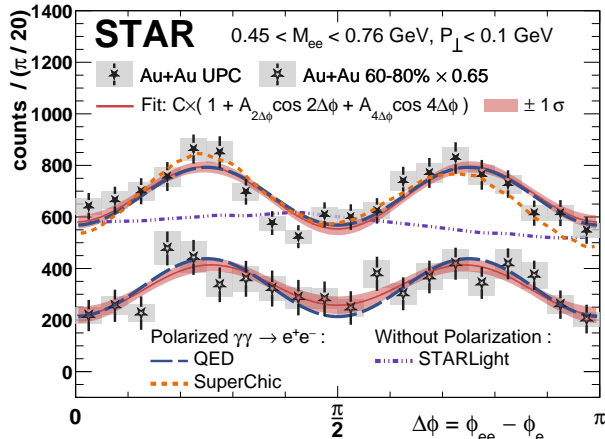


Figure 12: The $\Delta\phi = \phi_{ee} - \phi_e$ distribution from UPCs and 60-80% central collisions for $M_{ee} > 0.45$ GeV with calculations from QED, STARLight, and SuperChic3.

(^{238}U), larger than the nuclear charge radii. The observable is demonstrated to be sensitive to the nuclear geometry and quantum interference of non-identical particles.

Understanding gluon density distributions and their modifications in nuclei are among the most important goals of nuclear physics. Diffractive vector meson production measured in UPCs at heavy-ion colliders has provided a new tool for probing the gluon density. STAR has measured J/ψ photoproduction off the deuteron in UPCs at $\sqrt{s_{NN}} = 200$ GeV in d+Au collisions. [41] The differential cross section as a function of momentum transfer $-t$ is shown in fig. 14. In addition, cross section data with a neutron tagged in the deuteron-going Zero-Degree Calorimeter is found to be consistent with the expectation of incoherent diffractive scattering at low momentum transfer. Theoretical predictions based on the Color Glass Condensate (CGC) saturation model and the Leading Twist Approximation (LTA) nuclear shadowing model are compared with the data quantitatively. A better agreement with the saturation model has been observed. With the current measurement, the results are found to be directly sensitive to the gluon density distribution of the deuteron and the deuteron breakup process, which provides insights into the nuclear gluonic structure.

Copious amounts of dielectrons can also be produced by heavy-ion collisions that interact with enough energy to produce a quark-gluon plasma. As this super-heated phase of QCD matter cools, the QGP radiates e^+e^- pairs. Since leptons may travel away from the medium unimpeded by the dense environment of strongly interacting matter, they provide a pristine probe of the temperature of the emitting thermal source. Further, since the dielectrons are not effected by the collective motion of the rapidly expanding fireball, their spectrum is not blue-shifted but instead reveal the true temperature of the medium.

Various phases of the cooling QCD matter may be individually probed by analyzing dielectrons with various invariant masses, with higher invariant mass pairs corresponding to early times, and lower invariant masses corresponding to later times. In the low mass region ($0.2 < M_{ee} < 1.2$ GeV/ c^2), thermal dielectrons are predicted to originate predominately from radiation of the in-medium ρ^0 meson in the hadronic phase. In this region, the temperature

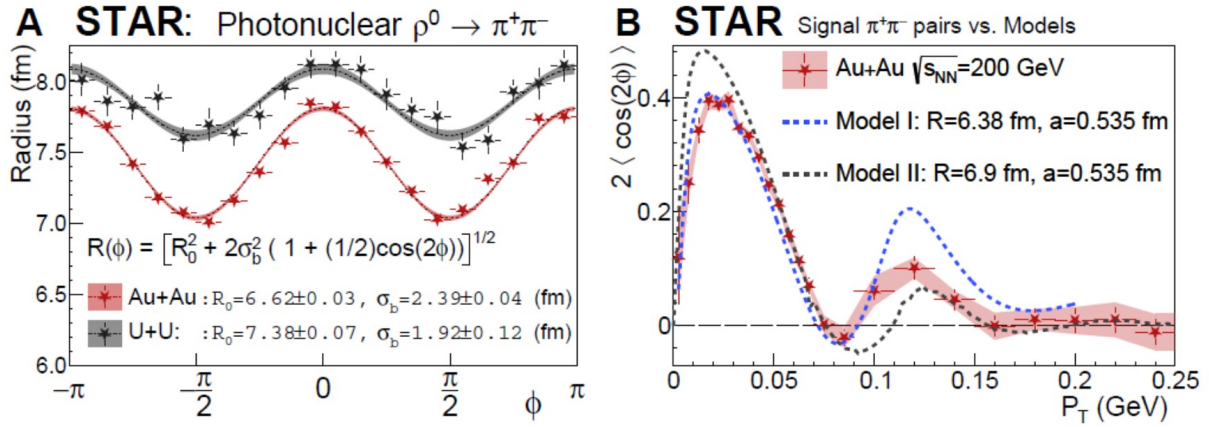


Figure 13: (Left) Radial parameter as a function of the ϕ angle for Au+Au and U+U with an empirical second order modulation fit. (Right) Comparison between the fully corrected Au+Au distribution and theoretical calculations that include the photon's linear polarization and two source interference effects.

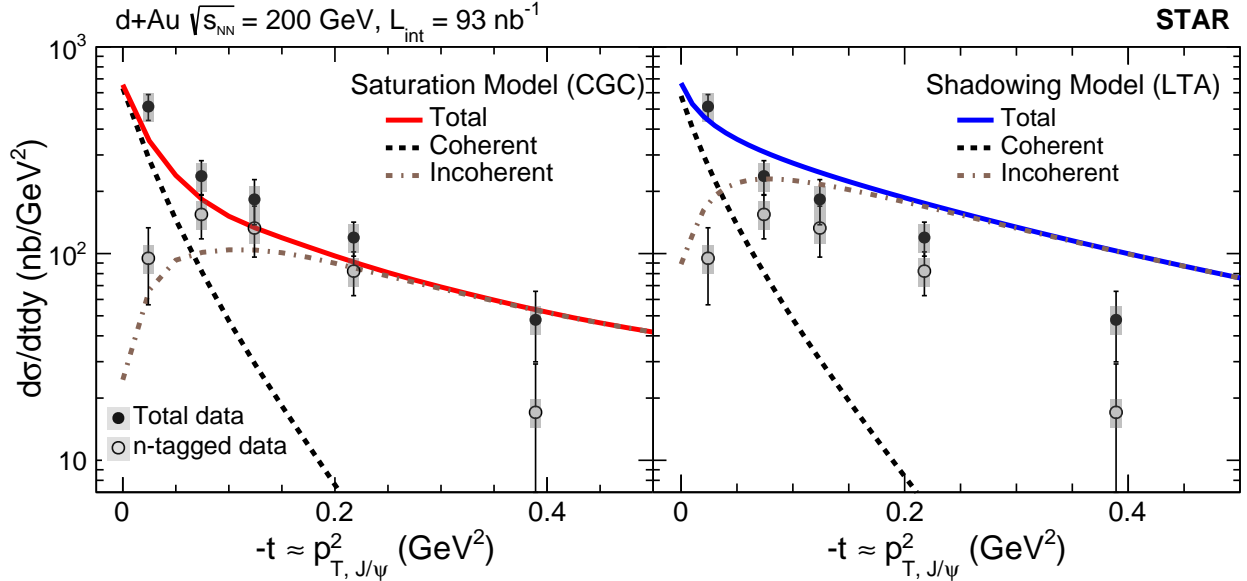


Figure 14: Theoretical predictions of the CGC saturation model (left) and the LTA nuclear shadowing model(right). Coherent and incoherent contributions from the two models are presented separately by the dashed lines.

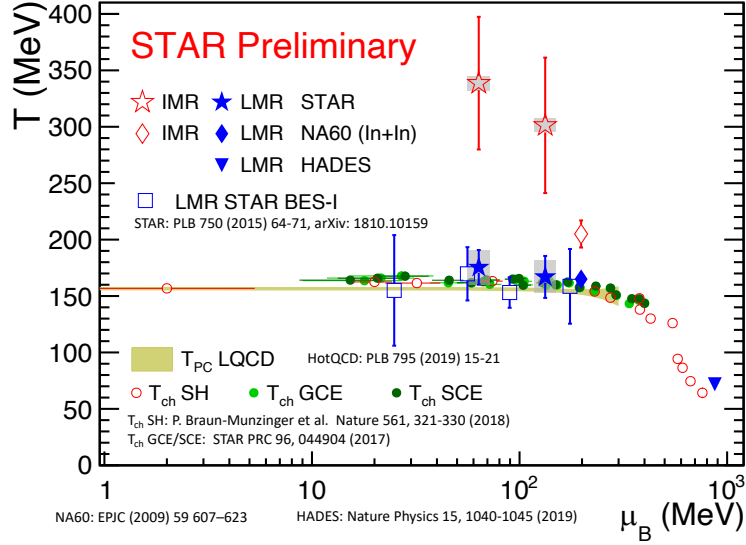


Figure 15: Extracted temperatures vs. baryonic chemical potential (μ_B). Temperatures extracted from the in-medium ρ^0 (blue squares) dominant region and the QGP (open red square) dominant region from STAR data are compared to the temperatures extracted from NA60 data (blue rhombus and open red rhombus) and HADES (blue triangle). Temperatures extracted from the statistical models (GCE, SCE, SH) are shown as solid dots and open circles. The QCD pseudo-critical temperature T_C vs. μ_B at small chemical baryon density predicted by lattice QCD calculations are shown at dotted blue lines.

373 are extracted to be 167 ± 20 MeV and 174 ± 15 MeV, in $\sqrt{s_{NN}} = 27$ GeV and 54.4 GeV
 374 Au+Au collisions, respectively. These measured temperatures are surprisingly consistent
 375 with the temperature (165 ± 4 MeV) extracted from the NA60 data measured in $\sqrt{s_{NN}}$
 376 $= 17.3$ GeV In+In collisions – a much lower collision energy and a significantly smaller
 377 collision system. These temperature measurements provides the first strong evidence that
 378 the in-medium ρ^0 mesons are dominantly produced around a constant temperature close to
 379 the phase transition boundary temperature (156 ± 1.5 MeV) as predicted by lattice QCD
 380 calculations. On the other hand, in the higher mass region ($1.0 < M < 2.9$ GeV/ c^2), the
 381 temperatures is extracted to be 301 ± 60 MeV and 338 ± 59 MeV, in $\sqrt{s_{NN}} = 27$ GeV
 382 and 54.4 GeV Au+Au collisions, respectively. These temperature values, which are well
 383 above the phase transition temperature, indicate that these thermal dielectrons originate
 384 predominantly from radiation of the ultra-hot phase of deconfined QCD matter, the quark-
 385 gluon plasma.

386 1.1.4 Heavy-flavor Measurements

387 Heavy-flavor (HF) quarks are produced predominately via initial hard scatterings of partons
 388 in p(A)+p(A) collisions. Kinematic distributions and hadronization probabilities of HF
 389 quarks in \bar{A} collisions can be different than those in $p+p$ collisions due to interactions of
 390 HF quarks with the QGP medium. Understanding these differences allows us to determine
 391 properties of the QGP.

392 STAR has recently published two papers on heavy flavor production: 1) the measurement
 393 of cold nuclear matter effects for inclusive J/ψ in $p+Au$ collisions at $\sqrt{s_{NN}} = 200$ GeV [42]
 394 and 2) measurement of inclusive electrons from open heavy-flavor hadron decays in $p+p$
 395 collisions at $\sqrt{s} = 200$ GeV [43].

396 J/ψ production has been found to be suppressed in Au+Au collisions at RHIC top en-
 397 ergies [44, 45]. Such a suppression can be caused by the color screening of the $c\bar{c}$ potential
 398 by the QGP medium, and by cold nuclear matter (CNM) effects from e.g., nuclear parton
 399 distribution functions, energy loss or absorption in the nucleus, and interaction with co-
 400 moving hadrons. Therefore, in order to precisely determine the suppression due to the color
 401 screening effect alone, it is important to quantify the CNM effects. The former paper reports
 402 the nuclear modification factor R_{pA} for inclusive J/ψ at mid-rapidity through the dimuon
 403 decay channel. At low $p_T < 2$ GeV/ c , a suppression of approximately 30% is observed in-
 404 dicating that the CNM effects contribute significantly to the J/ψ suppression in heavy-ion
 405 collisions in this p_T range. On the other hand, higher p_T J/ψ (> 3 GeV/ c) are observed
 406 to be minimally affected by the CNM effects. This provides evidence that the strong J/ψ
 407 suppression in Au+Au collisions at higher p_T is due to the presence of the QGP. The mea-
 408 surement provides also further constrains on model calculations of the CNM effects for J/ψ .
 409 The latter paper provides a high precision reference for measurements of R_{AA} for inclusive
 410 electrons from open-charm and -bottom hadron decays in heavy-ion collisions. Compare
 411 to the previous measurements, the precision was significantly improved for $p_T > 6$ GeV/ c ,
 412 which provides also additional constrains on theoretical pQCD calculations.

413 In heavy-ion collisions, in addition to the color screening effect that suppresses the J/ψ

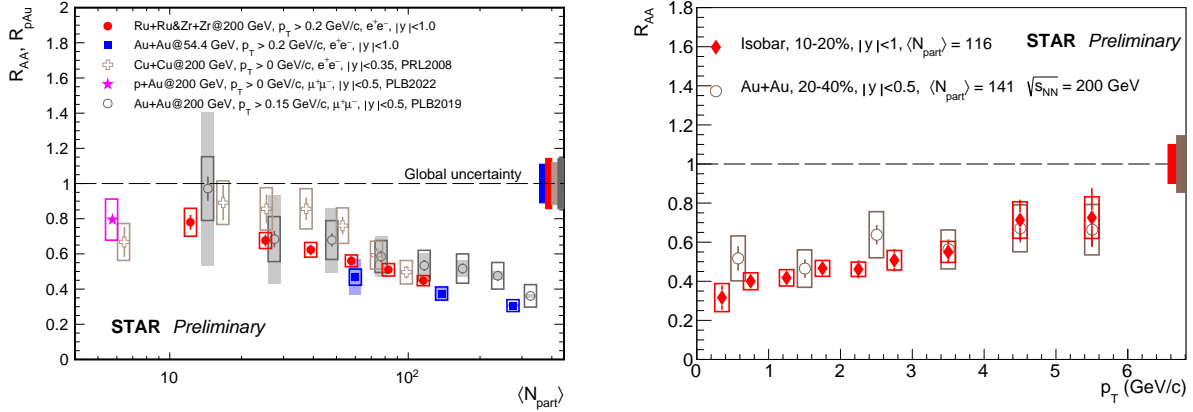


Figure 16: Left: R_{AA} vs. N_{part} . for inclusive J/ψ . Red circles: Ru+Ru and Zr+Zr collisions at $\sqrt{s_{NN}} = 200$ GeV (this analysis), blue squares: Au+Au collisions at $\sqrt{s_{NN}} = 54.4$ GeV, open circles: Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV [45], open crosses: Cu+Cu collisions at $\sqrt{s_{NN}} = 200$ GeV [46], magenta star: p +Au collisions at $\sqrt{s_{NN}} = 200$ GeV [42]. Right: R_{AA} vs. p_T for inclusive J/ψ at $\sqrt{s_{NN}} = 200$ GeV. Red diamonds: Ru+Ru and Zr+Zr for 10–20% centrality, open circles: Au+Au for 20–40% centrality. [45]

414 production, J/ψ can be produced from recombination of uncorrelated c and \bar{c} in the QGP.
 415 STAR has recently reported preliminary result on the nuclear modification factor R_{AA} of
 416 inclusive J/ψ in Ru+Ru and Zr+Zr collisions. The result is extracted in the dielectron
 417 channel from isobar data at $\sqrt{s_{NN}} = 200$ GeV collected in 2018. Isobar collisions being
 418 smaller (larger) collision systems compare to Au+Au (Cu+Cu) allow us to study the depen-
 419 dence of the hot nuclear matter effects - color screening vs recombination - on the medium
 420 size and geometry at the same collisions energy. As can be seen in Fig. 16(left) R_{AA} decreases
 421 with N_{part} and no significant species dependence is observed. The result is also consistent
 422 with the preliminary R_{AA} in Au+Au collisions at $\sqrt{s_{NN}} = 54.4$ GeV, confirming the previous
 423 observation of no significant energy dependence of the J/ψ suppression at RHIC that sug-
 424 gests a partial cancellation of the J/ψ suppression due to the color screening effect by J/ψ
 425 produced from recombination. R_{AA} as a function of p_T in isobar collisions shows increasing
 426 trend in central and mid-central collisions. When compared to the Au+Au measurement at
 427 similar $\langle N_{part} \rangle$ the two results are in agreement, see Fig. 16(right).

428 1.1.5 Jet Measurements

429 Jet is a useful tool to study the properties of QGP. With the help of newly developed tech-
 430 niques and significantly increased statistics in recent RHIC runs, STAR has explored various
 431 aspects of jet properties in heavy ion and pp collisions. In this section, we first briefly discuss
 432 recent publications of jet measurements in $p+p$ and Au+Au collisions and then highlight new
 433 studies with tagged jets, system size dependence of jet quenching and a new data driven way
 434 of estimating jet formation time with a study of jet splittings.

435

436 **Recent published results:**

437 As jets are composite objects built from parton showers and fragmentation, they contain
 438 rich substructure information that can be exploited via jet finding algorithms [47]. These
 439 algorithms typically employ an iterative clustering procedure that generates a tree-like struc-
 440 ture, which upon an inversion, gives access to a jet’s substructure at different steps along
 441 the cluster tree. The most common toolkit for such measurements is SoftDrop grooming [48]
 442 which employs a Cambridge/Aachen (C/A) re-clustering of a jet’s constituents and imposes a
 443 criterion at each step as we walk backwards in the de-clustered tree. The SoftDrop kinematic
 444 variables are,

$$z_g = \frac{\min(p_{T,1}, p_{T,2})}{p_{T,1} + p_{T,2}} > z_{\text{cut}} \left(\frac{R_g}{R_{\text{jet}}} \right)^\beta ; R_g = \Delta R(1, 2). \quad (2)$$

445 Where $z_{\text{cut}} = 0.1$ is a momentum fraction threshold and β is the angular exponent which in
 446 our analysis is set to zero [48]. The subscripts 1 and 2 represent the constituent jet pairs
 447 in re-clustered tree with C/A algorithm. These parameters make the SoftDrop observable
 448 comparable to theoretical calculations, and at the infinite momentum limit they converge
 449 to the DGLAP splitting functions. A recent STAR publication highlighted in PRC presents
 450 the differential measurements of jet substructure and partonic energy loss in Au+Au and
 451 $p+p$ collisions through substructure observables of SoftDrop z_g , R_g , and subjet momentum
 452 fraction ($z_{S,J}$) and opening angle ($\Theta_{S,J}$) [49]. In these studies, no significant modifications of
 453 the subjet observables are found in Au+Au collisions compared to $p+p$ collisions, implying
 454 vacuum- like splittings, with a possible interpretation that energy loss in this population of
 455 high momentum di-jet pairs is due to soft medium-induced gluon radiation from a single
 456 color-charge as it traverses the medium.

457 STAR also published the groomed and ungroomed jet mass in $p+p$ collisions at $\sqrt{s} = 200$
 458 GeV, together with comparisons to leading-order Monte Carlo event generators predic-
 459 tions [50]. In this study, while STAR-tuned PYTHIA-6 reproduced the data, LHC tunes
 460 of PYTHIA-8 and HERWIG-7 failed to do so. The agreement with STAR-tuned PYTHIA-6
 461 and disagreement with LHC tunes were also previously observed in z_g and R_g measurements
 462 of jets with a varying resolution parameters of $R=0.2-0.6$ for a wide transverse momentum
 463 range of $15 < p_{T,\text{jet}} < 60 \text{ GeV}/c$ in $p+p$ collisions at $\sqrt{s} = 200$ GeV [51]. These measurements
 464 establish a baseline for future jet mass measurements in heavy-ion collisions at RHIC and
 465 compliment LHC measurements at lower kinematic region to provide further tuning inputs
 466 to further constrain Monte Carlo simulations.

467 Preliminary STAR results on system size dependence of inclusive hadron suppression, jet
 468 formation time in $p+p$, flavor dependence of jet shape modification, intra-jet broadening and
 469 $\gamma_{\text{dir+jet}}$ (h+jet) acoplanarity measurements are discussed in the following paragraphs.

471 **System size dependence of inclusive charged hadrons suppression:** During the
 472 recent runs, RHIC facility provided us an opportunity to study system size dependence of
 473 jet quenching.

474 The left panel of Fig. 17 shows the inclusive charged hadron suppression (R_{AA}) as a

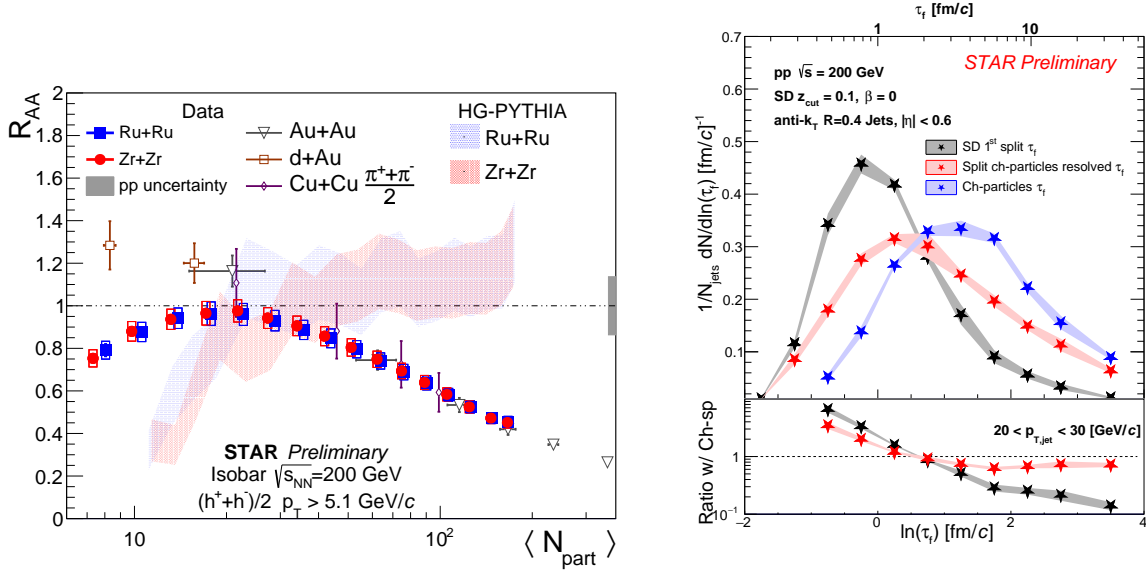


Figure 17: Left: Inclusive charged hadron R_{AA} in different collision systems. Right: Formation time distributions in $p+p$ collisions.

476 function of N_{part} for Ru+Ru, Zr+Zr, d+Au, Cu+Cu and Au+Au collisions at $\sqrt{s_{NN}} = 200$
 477 GeV. For Isobar (Ru+Ru, Zr+Zr) collisions, the charged hadrons are selected with $p_T > 5.1$
 478 GeV/c. It is observed that R_{AA} is independent of collision system for $N_{part} > 20$ with a
 479 decreasing trend. For $N_{part} < 20$. The HG-Pythia, that can describe the centrality bias
 480 observed at the LHC [52, 53], overpredicts the suppression observed in peripheral Ru+Ru
 481 and Zr+Zr collisions. Further studies including high- p_T hadron selection bias and differential
 482 measurement on path length dependence are ongoing.

483
 484 **Jet formation time and jet substructure:** STAR has recently explored the multi-
 485 scale nature of jet evolution in $p+p$ collisions. Utilizing the SoftDrop splitting momentum
 486 fraction (z) and opening angle (θ), it is possible to define a formation time at a given split
 487 as

$$\tau = \frac{1}{z(1-z)\theta^2 E}, \quad (3)$$

488 where E is the combined energy of the two objects used to calculate the z and θ . The
 489 black markers in the right panel of Fig. 17 are the formation times at the first SoftDrop splits
 490 for $R = 0.4$ jets with $20 < p_T < 30$ GeV/c. These splits correspond to mostly early times
 491 with the most probable value of the distribution being smaller than 1 fm/c. These splittings
 492 are expected to be predominantly perturbative in nature, which is supported by the fact
 493 that the substructure observables, such as z and θ , are well described by perturbative cal-
 494 culations. The blue markers in the same figure are the formation times calculated using the
 495 leading and sub-leading charged particles within the jet. This formation time is independent
 496 of the jet clustering history. As seen in the figure, charged-particle formation time shifts
 497 significantly towards later times as compared to the first SoftDrop splits. Via the red mark-

498 ers we introduce the resolved splittings which correspond to the formation time calculated
 499 from the jet clustering tree wherein the two leading charged-particles are first separated into
 500 two individual prongs. The bottom right panel of Fig. 17 shows the ratios of the clustering
 501 formation time distributions with respect to that of the charged particles. Comparison of
 502 the different splits highlights the transition from pQCD to npQCD. Resolved splits show
 503 a similar shape as the charged particle splits at large formation time occurring in the pre-
 504 dominantly non-perturbative region. These observables are presented in $p+p$ collisions as an
 505 outline for measurements in Au+Au collisions, leading towards a first ever space-time study
 506 of jet quenching phenomena.

507

508 **Flavour dependence of jet shape modification:** Jets with heavy quarks are expected
 509 to probe the full evolution of the QGP as they are produced early in the collision via hard
 510 partonic scatterings. To characterize the jet-medium interactions and distinguish between
 511 competing energy loss mechanisms, mass dependence of the energy loss needs to be also
 512 studied. Heavy-flavor mesons within a jet is expected to be sensitive to the production
 513 mechanism of mesons, energy loss and diffusion of heavy flavor quarks in the QGP.

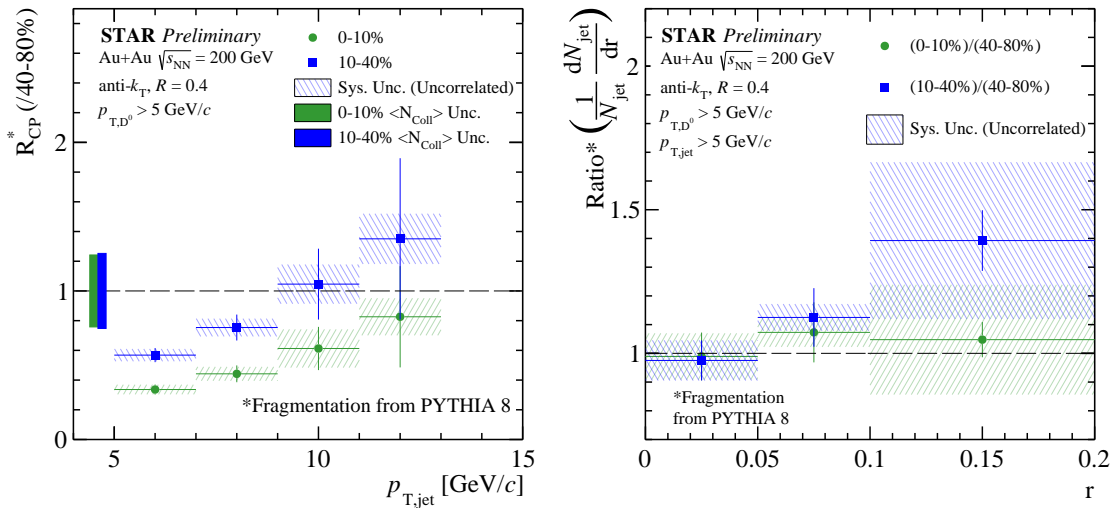


Figure 18: Left: Nuclear modification factor of D^0 -tagged jet. Right: Ratio of radial distribution of D^0 -tagged jets in central and mid-central collisions to that in peripheral Au+Au collisions.

514

515 The nuclear modification factor for jets that include a D^0 meson with $p_T > 5$ GeV/ c is
 516 shown in the left panel of Fig. 18. As can be seen in this figure, jets that are formed in the
 517 most central collisions appear to be more suppressed than those in mid-central collisions,
 518 especially for the lower p_T ranges of $5 < p_T < 10$ GeV/ c . The radial profile, i.e., the distri-
 519 bution of D^0 meson the distance from the jet axis (r), is also studied. As shown in the right
 520 panel of Fig. 18, the ratio of the radial distributions in most central collisions to that in most
 521 peripheral ones is consistent with unity within uncertainties. Theoretical calculations [54]

522 **Jet R dependence of suppression and intra-jet broadening:** In STAR, the $\gamma_{\text{dir}}/\pi^0$
 523 discrimination method using BEMC and BSMD detectors as well as uncorrelated background
 524 jet mitigation procedure using Mixed Event techniques are well calibrated to measure both
 525 $\gamma_{\text{dir}}+\text{jet}$ and $\pi^0+\text{jet}$ in $p+p$ and Au+Au collisions.

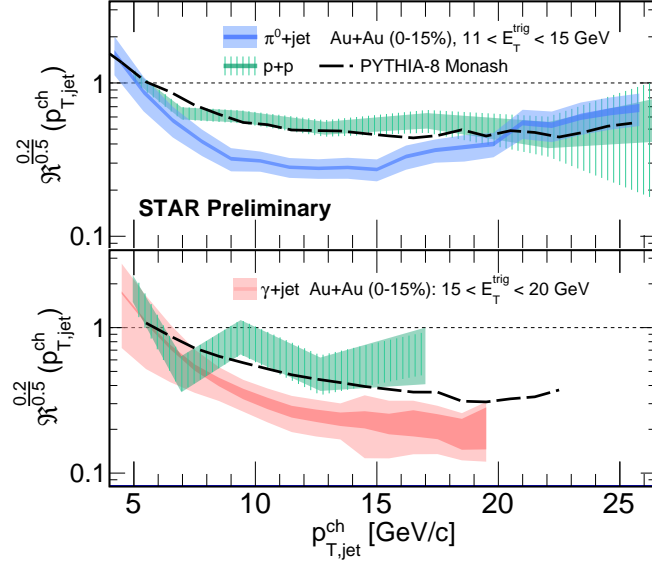


Figure 19: Yield ratio of recoil jets with $R = 0.2$ over $R = 0.5$ as a function of jet $p_{T,\text{jet}}^{\text{ch}}$. Upper and lower panels are for $\pi^0+\text{jet}$ and $\gamma_{\text{dir}}+\text{jet}$, respectively. Green bands are for $p+p$ collisions.

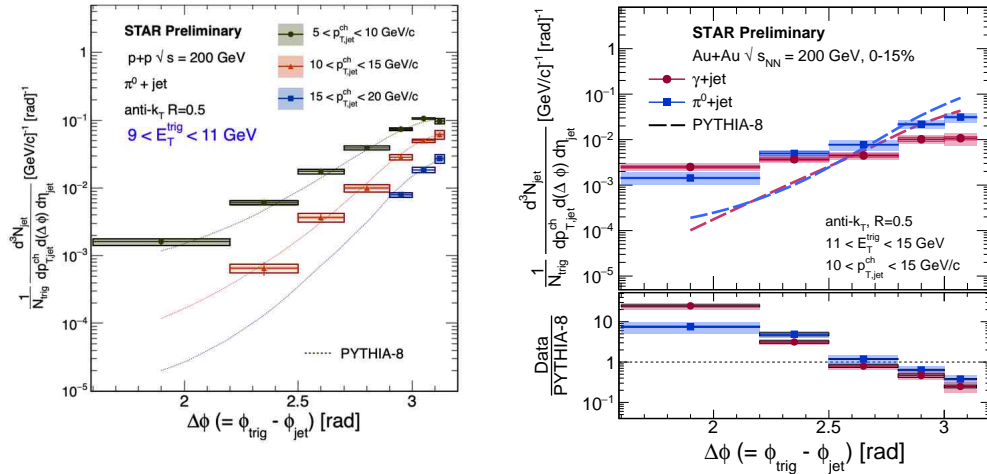


Figure 20: The $\gamma_{\text{dir}}+\text{jet}$ (red) and $\pi^0+\text{jet}$ (blue) acoplanarity measurements in $p+p$ (left) and central Au+Au (right) collisions $\sqrt{s_{\text{NN}}} = 200$ GeV. Dashed lines represent PYTHIA-8 predictions.

526 To investigate the resolution parameter dependence of the suppression of recoil jets, the
 527 jet yield ratios of jets that are reconstructed with $R = 0.2$ to those that are reconstructed

528 with $R = 0.5$ as a function of $p_{T,\text{jet}}^{\text{ch}} (R^{\frac{\text{small}-R}{\text{large}-R}})$ for π^0 +jet (upper panel) and γ_{dir} +jet (bottom
529 panel) shown in Fig. 19. The differences of $R^{\frac{\text{small}-R}{\text{large}-R}}$ in Au+Au to those in $p+p$ implies
530 intra-jet broadening in heavy-ion collisions due to jet quenching.

531 **γ_{dir} +jet and π^0 +jet acoplanarity in $p+p$ and Au+Au collisions:** At Born level,
532 dijet or γ_{dir} +jet productions in $p+p$ collisions are back-to-back in azimuth. However, soft-
533 gluon radiation and NLO effects introduce acoplanarity (decorrelation) between dijet or
534 γ_{dir} +jet even in vacuum. The acoplanarity measurement in $p+p$ is important studying QCD
535 effects. This also provides a baseline for similar measurement in heavy-ion collisions. Semi-
536 inclusive π^0 +jet (alike dijet) $\Delta\phi$ distributions in $p+p$ collisions are reported in the left panel
537 Fig. 20. Here $\Delta\phi$ represents the difference between trigger ϕ^{trig} and recoil jet ϕ^{jet} . The π^0
538 triggers are selected between $9 < E_T^{\text{trig}} < 11$ GeV. The $\Delta\phi$ distributions of three different
539 recoil jet $p_{T,\text{jet}}^{\text{ch}}$ ranges ($5 < p_{T,\text{jet}}^{\text{ch}} < 10$ GeV/c, $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/c, and $15 < p_{T,\text{jet}}^{\text{ch}} < 20$
540 GeV/c) are compared with the PYTHIA-8, and a good agreement is seen. Due to limited
541 statistics, this measurement in $p+p$ collisions for γ_{dir} +jet is not feasible.

542 In heavy-ion collisions, jet deflection is considered one of the consequences of the jet
543 quenching phenomenon. We report both γ_{dir} +jet and π^0 +jet $\Delta\phi$ measurements with 11
544 $< E_T^{\text{trig}} < 15$ GeV and $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/c for $R = 0.5$ in the right Fig.20. Striking
545 differences in the acoplanarity distributions between PYTHIA-8 and Au+Au collisions are
546 seen. A similar observation is made by ALICE for h+jet measurement in higher kinematic
547 range. Such measurements with extended E_T^{trig} and recoil jet $p_{T,\text{jet}}$ ranges are important
548 understanding the nature of the acoplanarity of jets produced in $p+p$ and Au+Au collisions.

549 Aforementioned semi-inclusive jet (like γ_{dir} +jet and h+jet) measurements and sub-structure
550 observables with extended kinematic coverage need high statistics data for precision and in-
551 cisive conclusions to understand the inner-working of QGP. Upcoming Run23-25 $p+p$ and
552 heavy-ion collision data taking will be crucial in achieving this goal and a detailed discussion
553 with projections can be found in Section 2.1.

554 1.1.6 BES-II Results

555 Data taking for the BES-II/FXT program has completed, with all data acquisition targets
556 being achieved or exceeded. Figure 21 shows a bar chart of the BES-II/FXT data sets
557 recorded and compares the new datasets to the older BES-I data. Also shown in the figure
558 are the energies for which we have overlapping coverage from both the collider and fixed-
559 target programs. The bars are plotted as a function of μ_B , which illustrates the range of μ_B
560 and the step size. For clarity, the collision energies ($\sqrt{s_{\text{NN}}}$) are indexed along the top edge
561 of the plot.

562 Data acquisition is only the first step in the process of data analysis. The calibrations
563 team must carefully perform run-by-run calibrations for all the detector systems prior to
564 ‘production’, which turns all of the raw information into tracks, time-of-flight, or energy
565 signals (depending on the detector sub-system) which can be used by the analyzers. Following
566 production, run-by-run QA is carried out to exclude runs for which the detector was not
567 performing optimally. It was expected that roughly 5% of the acquired data volume would

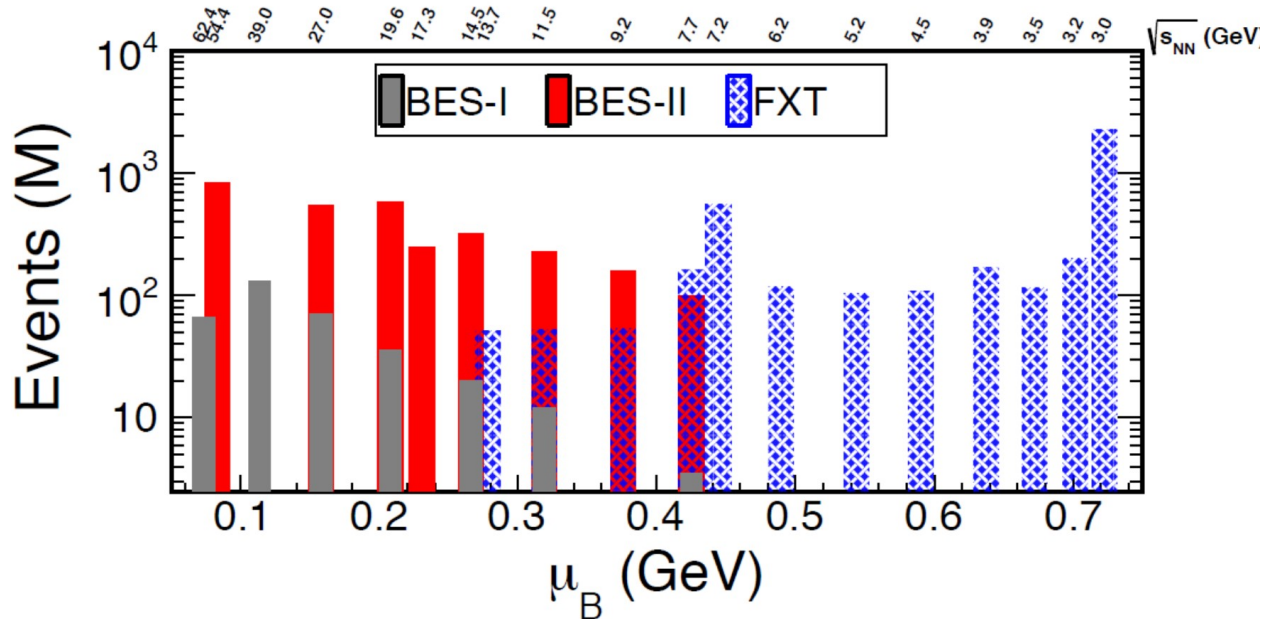


Figure 21: A summary of the good events acquired for the various collision energies (translated to μ_B). The BES-II collider data sets are shown in red bars. The FXT data sets are shown in hashed blue bars. For comparison the BES-I data sets are shown in grey bars. Note that the top FXT energy ($\sqrt{s_{NN}}=13.7$ GeV) does not quite overlap with the 14.6 GeV collider system; that FXT energy is a single beam energy of 100 GeV, which is the top energy to which RHIC can accelerate Au ions. Also note that the 54.4 GeV “BES-II” does not quite overlap with the 62.4 GeV BES-I system; the 54.4 GeV data were taken in 2017 parasitically with the first year of operation of the CeC program. This system is informally considered to be a part of the BES-II program. Likewise the data for the 7.2 GeV FXT system were parasitically acquired during single beam operations of CeC in 2018-2021.

568 be rejected in run-by-run QA. For the collider data sets, for which run-by-run QA has been
569 completed, we are indeed finding roughly 5% of the runs to be rejected. The fixed-target
570 data sets from 2019 and 2020 are passing run-by-run QA at a much higher rate, most likely
571 because they were all very short runs, and therefore the chance that a key detector component
572 fails during the run is much smaller. Following run-by-run QA, the centrality team defines
573 the basic event-by-event selection cuts (mostly to eliminate pile-up events) and defines the
574 centrality selections correcting for vertex position and luminosity. Figure 22 shows a table
575 of the energies acquired and status of each data set. This status is indicated with respect to
576 where it stands in the sequence of pre-analysis steps. For the data sets which are available to
577 the analysis teams, those listed as final are data sets for which papers have been published
578 or submitted. Those listed as preliminary are data sets for which preliminary results have
579 been shown at conferences.

580 The PAC recommended that STAR pay particular attention to analyses which are sensi-
581 tive to critical behavior. It was recommended "that the STAR collaboration does everything
582 possible to ensure that the analysis of critical observables in the Beam Energy Scan, such

2018	Start	Stop	Good	Target	Status
27 GeV	May 10 th	June 17 th	555 M	700 M	Final
3.0 FXT	May 30 th	June 4 th	258 M	100 M	Final
7.2 FXT	June 11 th	June 12 th	155 M	none	Final
2019	Start	Stop	Good	Target	
19.6 GeV	Feb 25 th	April 3 rd	478 M	400 M	Preliminary
14.6 GeV	April 4 th	June 3 rd	324 M	310 M	Post-prod QA
3.9 FXT	June 18 th	June 18 th	52.7 M	50 M	Produced
3.2 FXT	June 28 th	July 2 nd	200.6 M	200 M	Post-prod QA
7.7 FXT	July 8 th	July 9 th	50.6 M	50 M	Produced
200 GeV	July 11 th	July 12 th	138 M	140 M	Produced
2020	Start	Stop	Good	Target	Status
11.5 GeV	Dec 10 th	Feb 24 th	235 M	230 M	<i>Summer</i>
7.7 FXT	Jan 28 th	Jan 29 th	112.5 M	100 M	Produced
4.5 FXT	Jan 29 th	Feb 1 st	108 M	100 M	Produced
6.2 FXT	Feb 1 st	Feb 2 nd	118 M	100 M	Produced
5.2 FXT	Feb 2 nd	Feb 3 rd	103 M	100 M	Produced
3.9 FXT	Feb 4 th	Feb 5 th	117 M	100 M	Produced
3.5 FXT	Feb 13 th	Feb 14 th	115.6 M	100 M	Produced
9.2 GeV	Feb 24 th	Sep 1 st	161.8 M	160 M	<i>Summer</i>
7.2 FXT	Sep 12 th	Sep 14 th	317 M	None	<i>Fall</i>
2021	Start	Stop	Good	Target	Status
7.7 GeV	Jan 31 st	May 1 st	100.9 M	100 M	<i>May</i>
3.0 FXT	May 1 st	June 28 th	2103 M	2.0 B	<i>Fall</i>
9.2 FXT	May 6 th	May 6 th	53.9 M	50 M	<i>Fall</i>
11.5 FXT	May 7 th	May 7 th	51.7 M	50 M	<i>Fall</i>
13.7 FXT	May 8 th	May 8 th	50.7 M	50 M	<i>Fall</i>
17.3 GeV	May 25 th	June 7 th	256.1 M	250 M	<i>Fall</i>
7.2 FXT	June 3 rd	July 3 rd	88.6 M	None	<i>Fall</i>

Figure 22: A summary of the BES-II collider and FXT data sets taken from 2018-2021. The Start and Stop columns indicate the periods during which each data set was acquired. The Good and Target columns indicate the number of good events taken and requested. The Status column indicates where a given data set is in the analysis sequence.

583 as proton number cumulants, are carried out by at least two independent groups within
584 STAR". In addition to independent analyses, STAR has also decided not to release pre-
585 liminary results from such analyses, similar to the recommendation for the Chiral Magnetic
586 Effect analysis of the Isobar data. STAR has so far identified two lines of analysis that are
587 understood to address critical behavior: the net-proton cumulants, which are sensitive to
588 proton fluctuations, and the light nuclei ratios which are sensitive to neutron fluctuations.
589 Good progress has been made in both of these analysis efforts, although only the net-proton
590 fluctuations observed in the 3 GeV fixed-target data have matured to the point of journal
591 submission. For that analysis, a seminar at BNL was scheduled to coincide with submis-
592 sion of the results to PRL (December 2, 2021). The first presentation of these results at a
593 conference was at the recent QM2022.

594 The BES-II collider and FXT proposals identified a series of key physics analyses which
595 would have sensitivity to: formation of the QGP, the first order phase transition, the critical
596 point, and chirality. For all of these analyses, the collaboration determined the required event
597 count needed to make a definitive measurement and those events counts were used to set the
598 required number of events at each energy (see the "Target" numbers in Fig. 22). Analysis
599 teams have been identified to address all of these topics. Figure 23 shows the status of all
600 of these various analysis efforts. Significant progress has been made on all topics, with the
601 exception of the Chiral Magnetic Effect (CME). In the case of the expected CME analyses,
602 the teams with the requisite expertise have been fully committed to the analysis of the
603 isobar data and have not yet had a chance to turn their efforts to the new BES-II datasets
604 which are available for physics analyses. For all other expected lines of analysis first results
605 have been either published, submitted for publication, presented at QM2022, or are under
606 review within their respective PWGs. To date, publications have come from the 3 GeV FXT
607 data. Although some preliminary results have been shown for the 27, 19.6, and 14.6 GeV
608 collider datasets, it is expected that publications will wait until all of the collider energies
609 are available for physics analysis, which is expected to be in the Fall (see Fig. 22). The 3 GeV
610 FXT data set was unique enough to justify stand-alone papers. The next wave of papers
611 showing FXT results will cover the energy scan range from 3.0 to 7.7 GeV as those energies
612 are all now available. The three high energy FXT runs and the high statistics 3.0 GeV FXT
613 datasets from 2021 will be the last produced. Those data sets are for specialized analyses
614 which will likely result in another set of papers.

615 In this highlights section, we will focus on the published or submitted results from the
616 $\sqrt{s_{NN}} = 3$ GeV FXT system. Taken as a group, these results all show a marked change from
617 the behavior seen at collider energies of 7.7 GeV and above. It is not unexpected to see such
618 a significant change as the purpose of the FXT scan was to extend to reach of the energy
619 scan to regions for which QGP formation was likely not to be expected.

620 The first and second-order azimuthal anisotropy parameters v_1 and v_2 of light nuclei (p ,
621 d , t , He^3 , and He^4) were studied for 3 GeV Au+Au collisions. [55] The mid-rapidity slopes of the
622 directed flow (v_1) were found to scale with atomic mass number as shown in Fig. 24. The
623 elliptic flow (v_2) behavior is found to be unlike that at higher collision energies. The v_2 values
624 at mid-rapidity for all light nuclei are negative and no scaling is observed with the atomic

Physics Analysis	Status of Analyses
R_{CP} up to $p_T = 5$ GeV/c	Physics Working Group
Elliptic Flow	Published March 10, 2022
Chiral magnetic Effect	
Directed Flow	Published February 1, 2022
Azimuthal Femtoscopy	Physics Working Group
Net-proton Kurtosis	Submitted December 2, 2021
Di-leptons	QM2022 talk
Lambda Polarization	Published December 21, 2021
Multi-strange Baryons	Submitted December 23, 2021
Hyper-nuclei	Submitted October 18, 2021
Rapidity Dependent Spectra	QM2022 talks (2)

Figure 23: A summary of the key physics analyses which were listed in the BES-II collider and FXT proposals and the status of the efforts on each line of analysis.

625 mass number. Calculations using the Jet AA Microscopic Transport Model (JAM), with
626 baryonic mean-field plus nucleon coalescence, are in good agreement with the observations,
627 implying baryonic interactions dominate the collective dynamics in 3 GeV Au+Au collisions
628 at RHIC.

629 The partonic scaling of the elliptic flow (v_2) seen for various mesons and baryons at 200
630 GeV was seen as a signature of QGP formation and an indication that collective flow was
631 established during the partonic phase of the collisions. It is expected that at lower energies
632 this scaling should break down when one is below that produce a QGP phase. The v_2 results
633 for hadrons are shown in Fig. 25 for $\sqrt{s_{NN}} = 3, 27,$ and 54.5 GeV Au+Au collisions. While at
634 the two higher energy mid-central collisions the number-of-constituent-quark (NCQ) scaling
635 holds, at 3 GeV the v_2 at mid-rapidity is negative for all hadrons and NCQ scaling is absent.
636 It is not unexpected, or necessarily conclusive, that the v_2 is negative at the 3 GeV energy
637 as this had been seen in previous measurements and has been described as “squeeze-out”.
638 What is more telling, and had not been previously measured, is that the scaled v_2 of pions
639 is so different from that of protons and kaons. JAM and UrQMD model calculations with
640 baryonic mean-field potential reproduce the observed negative values of v_2 for protons at 3
641 GeV. This indicates that partonic interactions no longer dominate and baryonic scatterings
642 take over. This observation is clear evidence that predominantly hadronic matter is created
643 in such low energy collisions.

644 As a function of collision energy, a rise and then fall of the net-proton C_4/C_2 (or $\kappa|\sigma^2$)
645 has been predicted to indicate the critical behavior expected near the critical point in the
646 QCD phase diagram. Results from BES-I had shown an enhancement at 7.7 GeV and

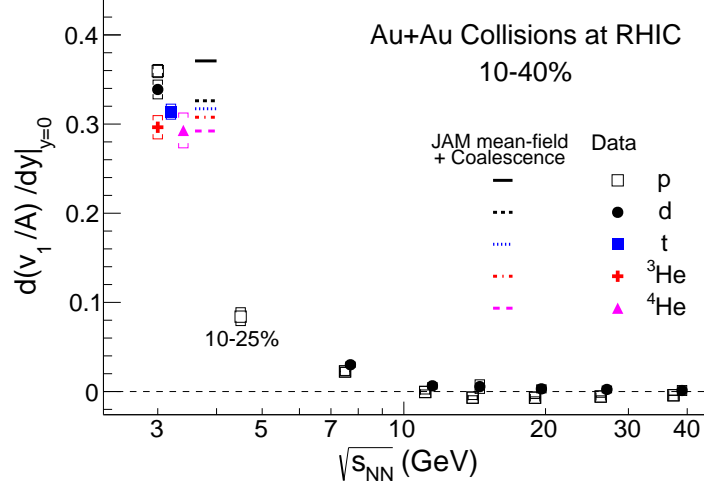


Figure 24: Light nucleus scaled v_1 slopes as a function of collision energy in 10-40% mid-central Au+Au collisions.

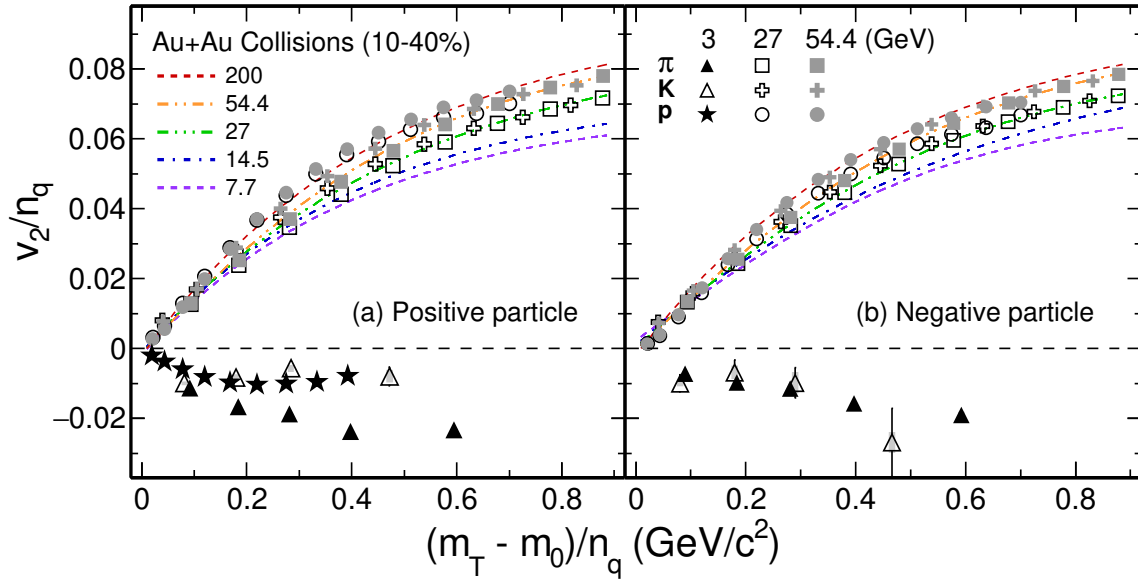


Figure 25: v_2 scaled by the number of constituent quarks, v_2/n_q , as a function of scaled transverse kinetic energy $((m_T - m_0)/n_q)$ for pions, kaons and protons from Au+Au collisions in 10-40% centrality at $\sqrt{s_{NN}} = 3, 27,$ and 54.4 GeV for positively charged particles (left panel) and negatively charged particles (right panel).

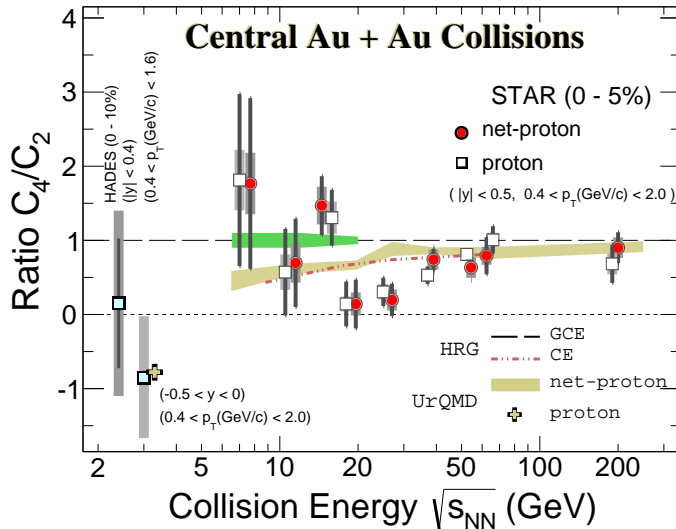


Figure 26: Collision energy dependence of the ratios of cumulants, C_4/C_2 , for proton (squares) and net-proton (red circles) from top 0–5% Au+Au collisions at RHIC. The points for protons are shifted horizontally for clarity. The new result for proton from $\sqrt{s_{\text{NN}}} = 3.0$ GeV collisions is shown as a filled square. HADES data of $\sqrt{s_{\text{NN}}} = 2.4$ GeV 0–10% collisions is also shown. Results from the HRG model and transport model UrQMD are shown.

647 a subsequent fall around 20 GeV. In order to determine if the value observed above the
 648 Poisson baseline is a peak it is necessary both to remeasure that point with high precision
 649 and also to carefully measure points at both higher and lower energies. At very low energies,
 650 where QGP formation is not expected, the C_4/C_2 signal should be consistent with baseline
 651 expectations. HADES has completed a measurement at $\sqrt{s_{\text{NN}}} = 2.4$ GeV, and their final
 652 result is below the Poisson baseline, albeit with large uncertainty, as shown in Fig. 26. Also
 653 shown in this figure is the new STAR result at $\sqrt{s_{\text{NN}}} = 3.0$ GeV. [56] The STAR result is well
 654 below the Poisson baseline and even negative. By comparing the STAR result to a UrQMD
 655 model, which has no phase transition, but does include baryon conservation, we conclude
 656 that this energy regime is dominated by hadronic interactions.

657 Global hyperon polarization, P_H , in Au+Au collisions over a large range of collision
 658 energy, $\sqrt{s_{\text{NN}}}$ was recently measured and successfully reproduced by hydrodynamic and
 659 transport models with intense fluid vorticity of the QGP. While a naive extrapolation of
 660 data trends suggests a increasing P_H as the collision energy is reduced, the behavior of
 661 P_H at very low energy is unknown. STAR has recently measured the polarization of Λ
 662 hyperons along the direction of global angular momentum in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 3$
 663 GeV as shown in Fig. 27. [57] The observation of substantial polarization in these collisions
 664 may require a reexamination of the viscosity of any fluid created in the collision, of the
 665 thermalization timescale of rotational modes, and of hadronic mechanisms to produce global
 666 polarization.

667 Strange hadron yields as well as the ratios in Au+Au collisions at $\sqrt{s_{\text{NN}}} \approx 3$ GeV were mea-

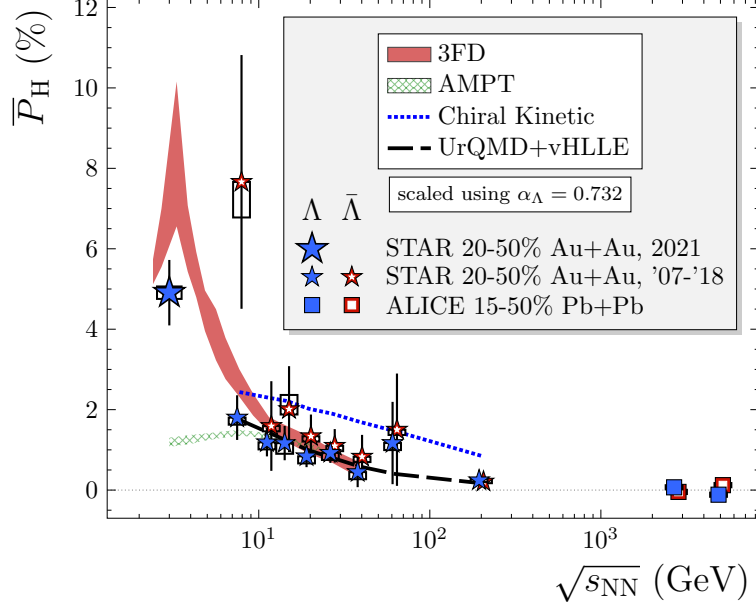


Figure 27: Global Hyperon (Λ) polarization as a function of $\sqrt{s_{\text{NN}}}$ in mid-central heavy-ion collisions. The trend of increasing P_H with decreasing $\sqrt{s_{\text{NN}}}$ is maintained at the low energy for $\sqrt{s_{\text{NN}}} = 3$ GeV.

668 sured. [58] The 4π yields and ratios are compared to thermal model and hadronic transport
 669 model predictions. At this collision energy, as shown in Fig. 28, the thermal model with
 670 grand canonical ensemble (GCE) under-predicts the ϕ/K^- and ϕ/Ξ^- ratios while the result
 671 of canonical ensemble (CE) calculations reproduce the ratios with correlation lengths r_c of
 672 3-4 fm. Thermal calculations with GCE work well for strangeness production in high energy
 673 collisions. The change to CE at 3 GeV implies a different medium property at high baryon
 674 density.

675 In relativistic heavy-ion collisions, hypernuclei form when hyperons (mostly Λ s coalesce
 676 with neutrons and protons to form nuclei. The study of such exotic nuclei allows one to better
 677 understand the hyperon-nucleon interaction and to determine if the lifetime of the hyperon
 678 is affected as it is bound into a nucleus. Thermal models have predicted that the maximum
 679 yield of hypernuclei should occur in the collision energy range covered by the STAR FXT
 680 program. Precision measurements of hypernuclei ${}^3_{\Lambda}H$ ${}^4_{\Lambda}H$ were obtained from Au+Au collisions
 681 at $\sqrt{s_{\text{NN}}} = 3.0$ GeV. [57] Their lifetimes are measured to be $221 \pm 15(\text{stat.}) \pm 19(\text{syst.})$
 682 ps for ${}^3_{\Lambda}H$ and $218 \pm 6(\text{stat.}) \pm 13(\text{syst.})$ ps for ${}^4_{\Lambda}H$. Figure 29 shows the p_T -integrated yields
 683 compared to model calculations. The thermal model, using the canonical ensemble for
 684 strangeness, describes the ${}^3_{\Lambda}H$ yield well, while underestimating the ${}^4_{\Lambda}H$ yield. Transport
 685 models, combining baryonic mean-field and coalescence (JAM) or utilizing dynamical cluster
 686 formation via baryonic interactions (PHQMD) for light nuclei and hypernuclei production,
 687 approximately describe the measured yields. These new measurements provide means to pre-
 688 cisely assess the understanding of the fundamental baryonic interactions with strange quarks,
 689 which can impact our understanding of more complicated systems involving hyperons, such

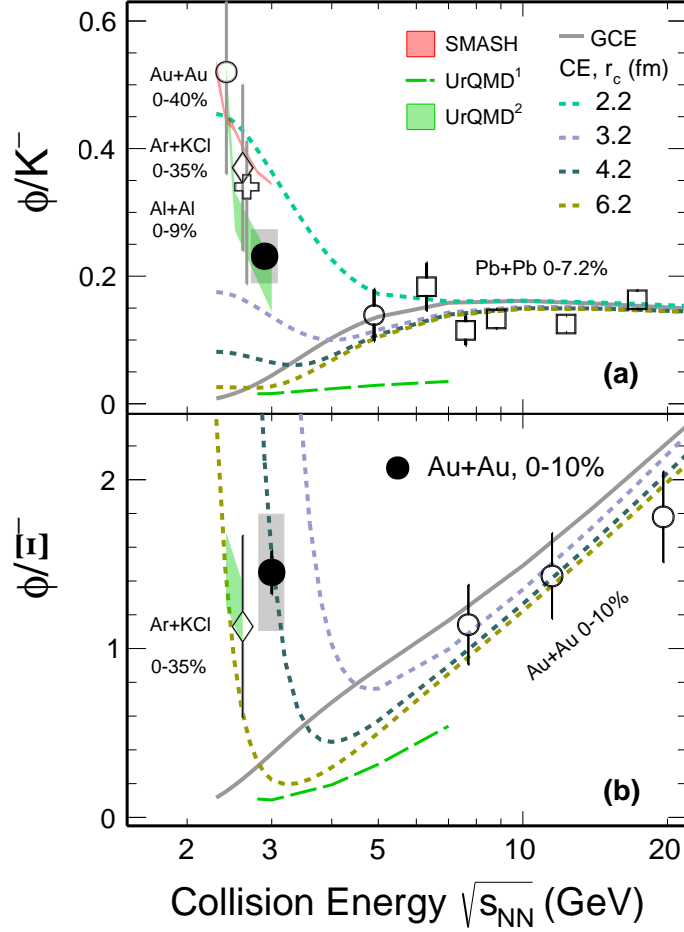


Figure 28: ϕ/K^- (top) and ϕ/Ξ^- (bottom) ratio as a function of collision energy, $\sqrt{s_{NN}}$. The solid black circles show the new STAR measurements. Data from other energies and/or collision systems are shown with open. The grey solid line represents a THERMUS calculation based on the Grand Canonical Ensemble (GCE) while the dotted lines depict calculations based on the Canonical Ensemble (CE) with different values of the strangeness correlation radius (r_c). The green dashed line, green shaded band and the solid red line show transport model calculations from UrQMD1, UrQMD2, and SMASH.

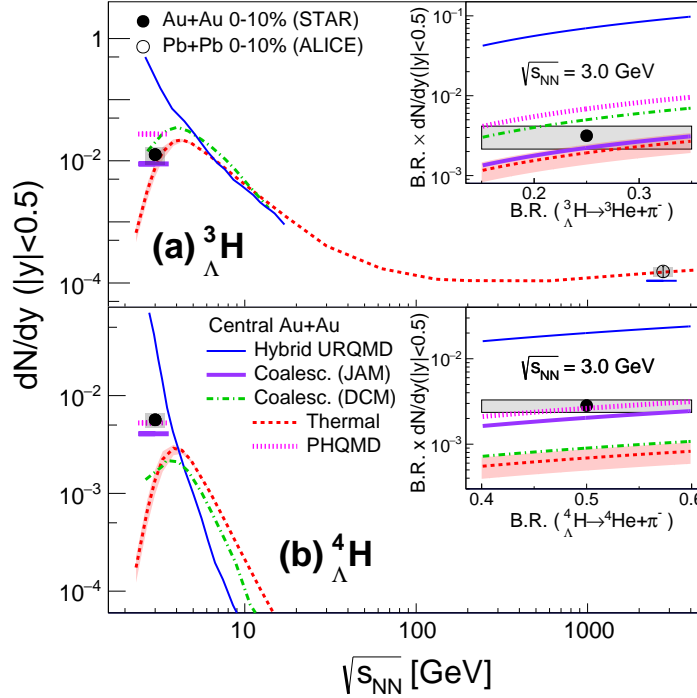


Figure 29: (Top) ${}^3_{\Lambda}H$ and (bottom) ${}^4_{\Lambda}H$ yields at $|y| < 0.5$ as a function of collision energy in central heavy-ion collisions. The solid black circle represent the new STAR measurements while the lines represent theoretical calculations.

690 as the interior of neutron stars or exotic hypernuclei.

691 1.2 Highlights from the Spin and Cold QCD Program

692 Introduction

693 The goal of the STAR Cold QCD program is to probe the spin and flavor structure of the
 694 proton and understand the role of spin in Quantum Chromodynamics, exploiting the unique
 695 capability of RHIC to provide longitudinally and transversely polarized $p+p$ collisions at mul-
 696 tiple energies. Measurements with longitudinal beam polarizations have given new insights
 697 into the helicity structure of the proton, while measurements with transverse polarizations
 698 have provided new ways to probe polarized parton distribution functions in the collinear and
 699 transverse momentum dependent frameworks. This program is complemented by studies
 700 of polarized $p+p$ elastic scattering and central exclusive production, in which a far-forward
 701 proton is detected intact.

702 Since 2009, RHIC STAR has completed several highly successful polarized $p+p$ runs
 703 both at $\sqrt{s} = 200$ GeV and $\sqrt{s} = 500/510$ GeV. Moreover, $p+Au$ and $p+Al$ data sets with a
 704 transversely polarized proton beam have been recorded in 2015 at $\sqrt{s} = 200$ to address im-
 705 portant physics problems, including the underlying non-perturbative mechanism responsible
 706 for large forward transverse single spin asymmetries, the ridge phenomenon and the possible

707 onset of gluon saturation effects. Table 3 summarizes the STAR sampled luminosity and the
 708 luminosity averaged beam polarization as measured by the hydrogen jet (H-jet) polarimeter.

Table 3: Summary of polarized $p+p$ and $p+A$ running periods at RHIC since 2009, including center-of-mass energy, STAR’s integrated luminosity and the average beam polarization for blue (B) and yellow (Y) beams from the H-jet polarimeter.

Year	System	\sqrt{s} (GeV)	Recorded Lumi. (pb^{-1})	Polarization Orientation	B/Y $\langle P \rangle$ (%)
2009	$p+p$	200	25	Longitudinal	55/55
2009	$p+p$	500	10	Longitudinal	39/39
2011	$p+p$	500	12	Longitudinal	48/48
2011	$p+p$	500	25	Transverse	48/48
2012	$p+p$	200	22	Transverse	61/56
2012	$p+p$	510	82	Longitudinal	50/53
2013	$p+p$	510	300	Longitudinal	51/52
2015	$p+p$	200	52	Transverse	53/57
2015	$p+p$	200	52	Longitudinal	53/57
2015	$p+\text{Au}$	200	0.45	Transverse	60/-
2015	$p+\text{Al}$	200	1	Transverse	54/-
2017	$p+p$	510	320	Transverse	55/55
2022	$p+p$	510	400	Transverse	52

709 Since the last PAC meeting, there are three very mature analyses, which have either
 710 been accepted, submitted or are about to be submitted for publication. One analysis on
 711 di-jet spin asymmetry which probes the contribution of gluon spin to the proton spin has
 712 been accepted to Phys. Rev. D. [59] Another which investigates non-linear gluon effects
 713 has been submitted to Phys. Rev. Lett. [60] and is in the second round of journal review.
 714 Finally, the analysis of the Collins asymmetry which is sensitive to transversity and the
 715 Collins fragmentation function is nearing submission to Phys. Rev. D. Additionally, the
 716 Sivers dijet analysis, which is sensitive the the quark Sivers functions have just formed GPC,
 717 which will work to have these results published in Phys. Rev. Lett.

718 **Inclusive Jet and Dijet A_{LL}**

719 Studies of the polarized gluon distribution function ($\Delta g(x)$) of the proton to gain deeper
 720 insight into its spin structure and dynamics, have been possible due to the unique longitudi-
 721 nally polarized proton-proton collision data provided by RHIC.

722 The STAR experiment collected several longitudinally polarized $p+p$ collision data sets,
 723 mainly dedicated to studying $\Delta g(x)$, which can be accessed by measuring the longitudinal
 724 double-spin asymmetry (A_{LL}) of inclusive jet and dijet production. The data were collected
 725 at center-of-mass energies of 200 GeV [61–63] and 510 GeV [64] at mid-rapidity, allowing

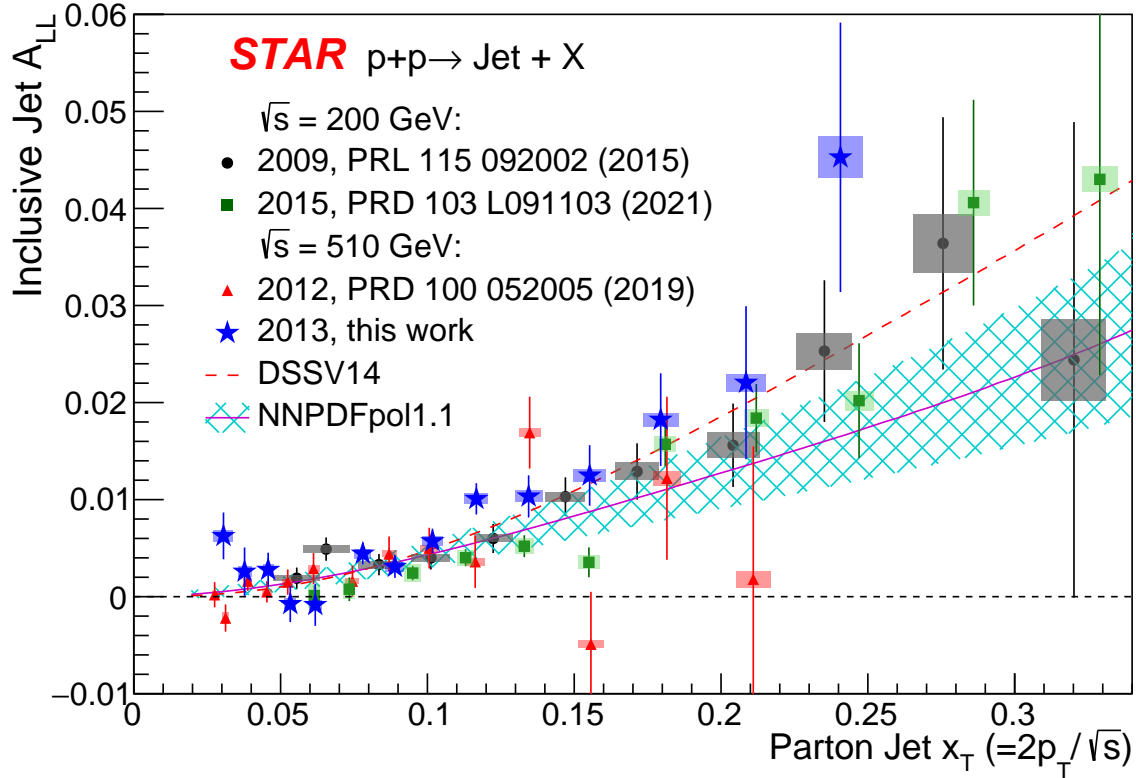


Figure 30: Inclusive jet A_{LL} versus x_T , compared to previous STAR results at $\sqrt{s} = 200$ GeV [61,63] and 510 GeV [64], and evaluations from DSSV14 [65] and NNPDFpol1.1 (with its uncertainty) [66] global analyses. The vertical lines are statistical uncertainties. The boxes show the size of the estimated systematic uncertainties. Scale uncertainties from polarization (not shown) are $\pm 6.5\%$, $\pm 6.6\%$, $\pm 6.4\%$ and $\pm 6.1\%$ from 2009 to 2015, respectively.

726 to probe a broader kinematic coverage in the partonic momentum fraction x . In 2015, the
 727 STAR concluded the collection of longitudinally polarized proton-proton collision data.

728 The recently published results on longitudinal double-spin asymmetry for inclusive jet
 729 and dijet production in polarized proton collisions at $\sqrt{s} = 510$ GeV in Phys. Rev. D [59],
 730 provides the last STAR A_{LL} measurements for inclusive jets at mid-rapidity, with data col-
 731 lected in 2013. These measurements complement and improve the precision of previous STAR
 732 measurements at the same center-of-mass energy that probe the polarized gluon distribution
 733 function at partonic momentum fraction $0.015 < x < 0.25$. The inclusive jet measurements
 734 A_{LL} , as shown in Figure 30, are in agreement with previous STAR measurements and with
 735 predictions from current next-to-leading-order global analyses. [65,66]

736 Additionally, results for dijet production are presented in Fig. 31. These measurements
 737 provide a better determination of the functional form of $\Delta g(x)$, compared to inclusive ob-
 738 servables, because better constraints on the underlying kinematics. At leading order, the
 739 dijet invariant mass is proportional to the square root of the product of the partonic momen-

Table 4: The four dijet topology bins A-D.

Bin	η_3 and η_4 Regions	Physics description
A	$0.3 < \eta_{3,4} < 0.9; \eta_3 \cdot \eta_4 > 0$	Forward-Forward
B	$ \eta_{3,4} < 0.3; 0.3 < \eta_{4,3} < 0.9$	Forward-Central
C	$ \eta_{3,4} < 0.3$	Central-Central
D	$0.3 < \eta_{3,4} < 0.9; \eta_3 \cdot \eta_4 < 0$	Forward-Backward

tum fractions, $M_{inv} = \sqrt{s x_1 x_2}$, and the pseudorapidity sum of the two jets is proportional to the logarithmic ratio of the x values, $\eta_3 + \eta_4 \propto \log(x_1/x_2)$ ¹. The individual jets in a dijet were separated into three pseudorapidity regions: forward $0.3 < \eta < 0.9$, central $-0.3 < \eta < 0.3$, and backward $-0.9 < \eta < -0.3$. The A_{LL} measurements for dijets are presented in four topology bins A-D (Table 4), as in [64], which allows discrimination between symmetric and asymmetric collisions in terms of the partonic momentum fractions x_1 and x_2 . With a redesigned and optimized set of triggers in 2013, we were able to increase the statistics in the low dijet mass region by approximately an order of magnitude, which is critical to enable a controlled extrapolation of the polarized gluon distribution function in this gluon-rich region, with x down to 0.015. Preliminary results of dijet measurements from 2012 [67] and 2013 [68] data at intermediate-pseudorapidity, will probe even lower values of x . These high precision measurements motivate the natural step forward for an Electron Ion Collider in order to study the gluon-rich region of the proton in even greater detail.

753 Di-Hadron Correlations

754 The STAR Collaboration recently submitted a paper [60] on measurements of back-to-back
 755 azimuthal correlations of di- π^0 s in $p+p$, $p+Al$, and $p+Au$ collisions at a center-of-mass energy
 756 of 200 GeV. The forward π^0 s ($2.6 < \eta < 4.0$) were reconstructed from the STAR forward
 757 meson spectrometer (FMS), with data recorded in 2015.

758 The correlation function is defined as $C(\Delta\phi) = \frac{N_{\text{pair}}(\Delta\phi)}{N_{\text{trig}} \times \Delta\phi_{\text{bin}}}$, where N_{pair} is the yield of the
 759 correlated trigger and associated π^0 pairs, N_{trig} is the trigger π^0 yield, $\Delta\phi$ is the azimuthal
 760 angle difference between the trigger π^0 and associated π^0 , and $\Delta\phi_{\text{bin}}$ is the bin width of $\Delta\phi$
 761 distribution. After the mixed event correction is applied, the correlation function is fitted
 762 with two individual Gaussians at the near- and away-side peak, together with a constant
 763 for the pedestal in the whole $\Delta\phi$ range. The area of the away-side peak used to describe
 764 the suppression, is defined as the integral of the correlation function from $\Delta\phi = \pi/2$ to
 765 $\Delta\phi = 3\pi/2$ after pedestal subtraction. The corresponding width is defined as the σ of the
 766 away-side peak according to the fit.

767 We observe a clear suppression of the correlated yields of back-to-back π^0 pairs in $p+Al$
 768 and $p+Au$ collisions compared to the $p+p$ data at low p_T , see the top panel of Fig. 32.
 769 The suppression disappears at high p_T where x (Q^2) is not sufficiently small to reach the
 770 nonlinear regime (bottom panel of Fig. 32). These results are the first measurements of

¹the kinematics of the initial partons and final jets are denoted by subscripts 1,2 and 3,4, respectively

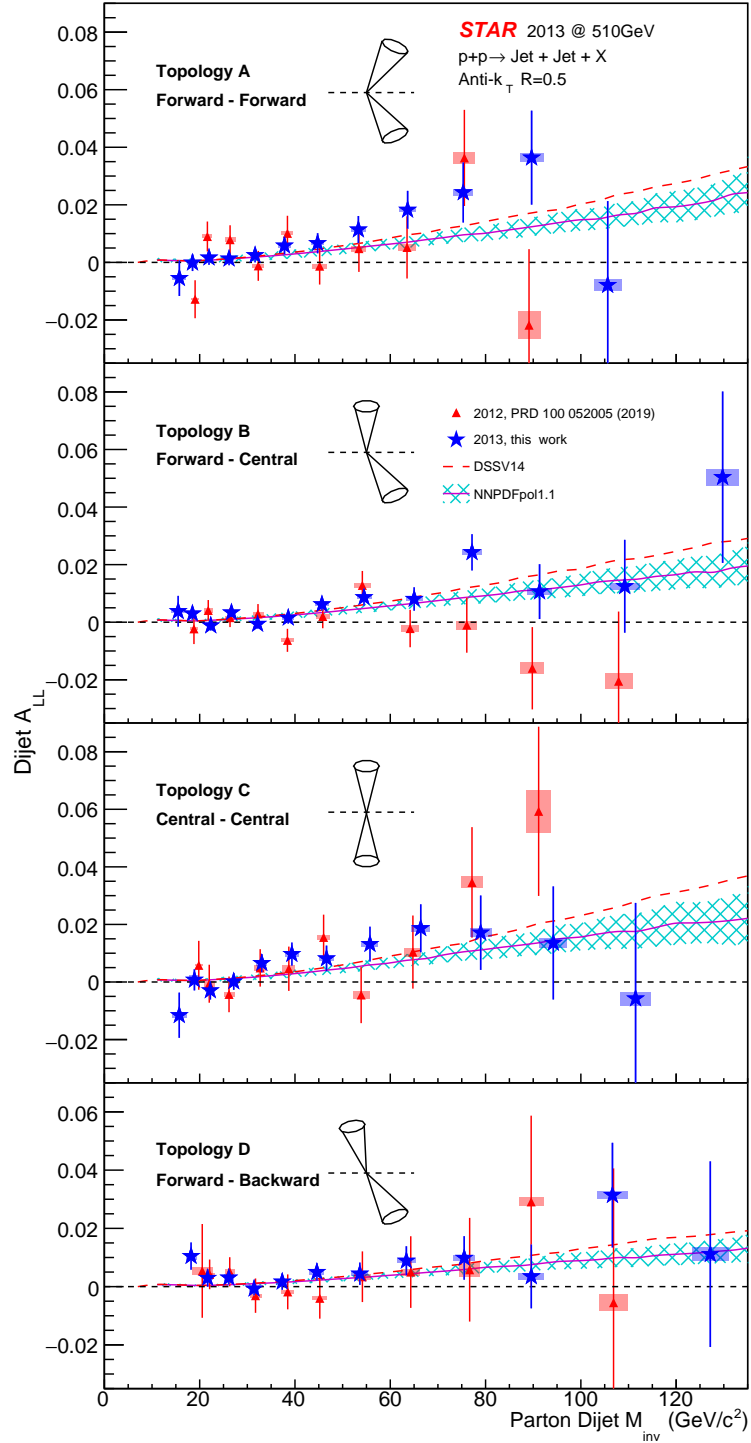


Figure 31: Dijet A_{LL} versus M_{inv} for the A, B, C and D (top to bottom) topological configurations as explained in the text. They are compared to previous STAR results from 2012 data [64] and predictions from DSSV14 [65] and NNPDFpol1.1 (with its uncertainty) [66] global analyses. The vertical lines are statistical uncertainties. The boxes show the size of the estimated systematic uncertainties. Topological configurations are shown for each jet orientation relative to the beam line. Scale uncertainties from polarization (not shown) are $\pm 6.6\%$ and $\pm 6.4\%$ for 2012 and 2013, respectively.

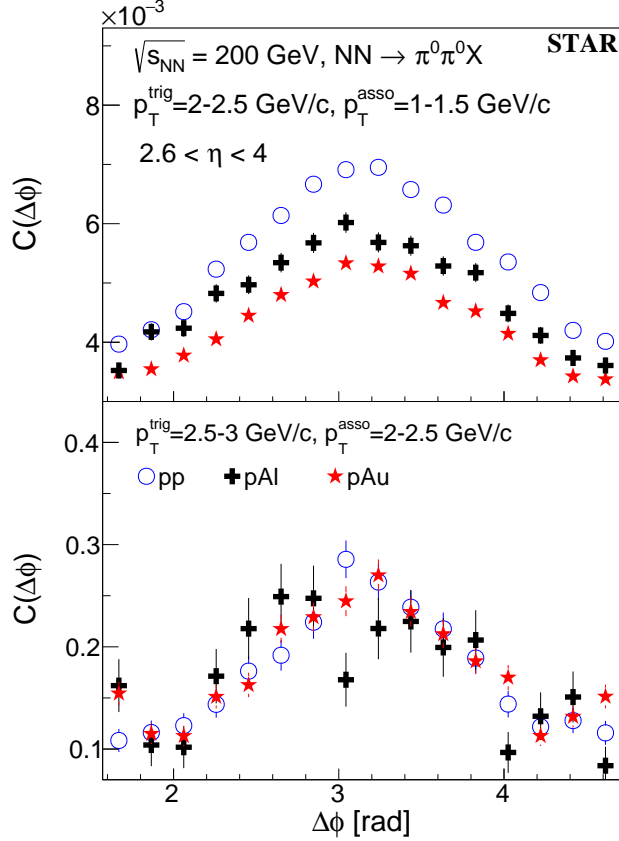


Figure 32: Comparison of the correlation functions (corrected for nonuniform detector efficiency in ϕ ; not corrected for the absolute detection efficiency) vs. azimuthal angle difference between forward ($2.6 < \eta < 4.0$) π^0 s in $p+p$, $p+Al$, and $p+Au$ collisions at $\sqrt{s_{NN}}=200$ GeV. Upper panel: the trigger π^0 's p_T ($p_T^{\text{trig}} = 2-2.5$ GeV/c and the associated π^0 's p_T ($p_T^{\text{asso}} = 1-1.5$ GeV/c; Bottom panel: $p_T^{\text{trig}} = 2.5-3$ GeV/c and $p_T^{\text{asso}} = 2-2.5$ GeV/c.

771 the A -dependence of the cold nuclear effect; the suppression is enhanced with higher A
772 and scales with $A^{1/3}$, see Fig. 33. The suppression is analyzed for various event activities
773 (E.A.) selections and found to be larger with higher E.A. The E.A. describes the degree of
774 violence of the collision and is defined as the energy deposition in the backward (aluminum
775 and gold going direction) inner sectors of the beam beam counter (BBC, $3.3 < -\eta < 5.0$).
776 The measured suppression in high E.A. $p+Au$ collisions is consistent with the predictions
777 calculated from the gluon saturation model [69]. Meanwhile, the broadening predicted in the
778 color glass condensate (CGC) framework in Ref. [70, 71] is not observed. This observation
779 agrees with a similar measurement in d+Au collisions by the PHENIX experiment. [72] The
780 pedestals in $p+A$ and $p+p$ collisions are found to be stable.

781 The comparison of the correlation function from $p+p$, $p+Au$, and d+Au collisions pro-
782 vides opportunities to understand the contributions from multiple parton scatterings [73].
783 From the preliminary results of d+Au collisions, we found much higher background in d+Au
784 collisions compared to $p+p$ and $p+Au$ collisions reconstructing the π^0 candidates. The gener-
785 ated combinatoric correlation dominates in d+Au collisions, which makes it very challenging
786 to identify the signal correlation. The forward di- π^0 correlation measurement favors for the
787 cleaner $p+A$ collisions rather than d+Au collisions.

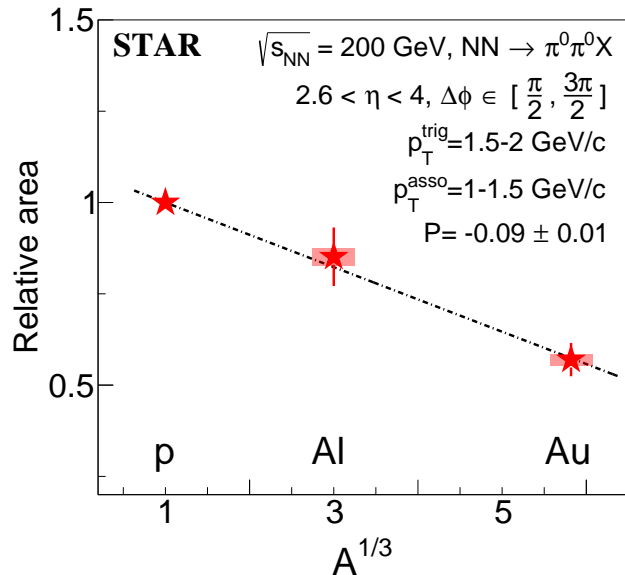


Figure 33: Relative area of back-to-back di- π^0 correlations at forward pseudorapidities ($2.6 < \eta < 4.0$) in p +Au and p +Al relative to p + p collisions for $p_T^{\text{trig}} = 1.5\text{--}2$ GeV/ c and $p_T^{\text{asso}} = 1\text{--}1.5$ GeV/ c . The vertical bars for the Al and Au ratios indicate the statistical uncertainties and the vertical bands indicate the systematic uncertainties. The data points are fitted by a linear function, whose slope (P) is found to be -0.09 ± 0.01 .

Collins Asymmetry

Recently, we finalized the measurement of the transverse single-spin asymmetries for charge pions inside a jet at p + p 200 GeV based on the data from 2012 and 2015 running [?]. These observables, so called the Collins asymmetries, combine the quark transversity in the proton with the transverse momentum dependent Collins fragmentation function. Both of them are important topics in the transverse-momentum-dependent (TMD) frameworks.

Figures 34 shows the j_T , momentum transverse to the jet axis, dependence of the Collins asymmetry in six jet- p_T bins, with the average hadron z about 0.22. DMP+2013 model and KPRY model expectations are also presented in the plot. The DMP+2013 model uses the leading order TMD approach, and is based on a fit to transversity and Collins fragmentation function measurements from SIDIS and e^+e^- processes [74]. The KPRY model is also based on the global analysis of SIDIS and e^+e^- data and then treat TMD evolution up to the next-to-leading logarithmic effects using the soft-collinear-effective theory framework [75]. Our results slightly favor the KPRY model, however significant discrepancies exist between the data and both model calculations.

We also presented the first measurement of the Collins asymmetries for charged kaons and protons inside jets at p + p collisions as shown in Fig. 35. These results are plotted with jet- p_T , hadron- z , and hadron- j_T dependence from left to right panels. Due to the limited statistics, they are not further divide into multi-dimensional bins. The asymmetries of K^+ has similar magnitude to those for π^+ ; while for K^- , proton and anti-proton, the asymmetries are consistent with zero at the one sigma level.

1.3 Run-22 Performance

All RHIC runs are challenging, however Run 22 seemed to have been more challenging than most. Despite the many set backs in the first half of the run, STAR ended up achieving

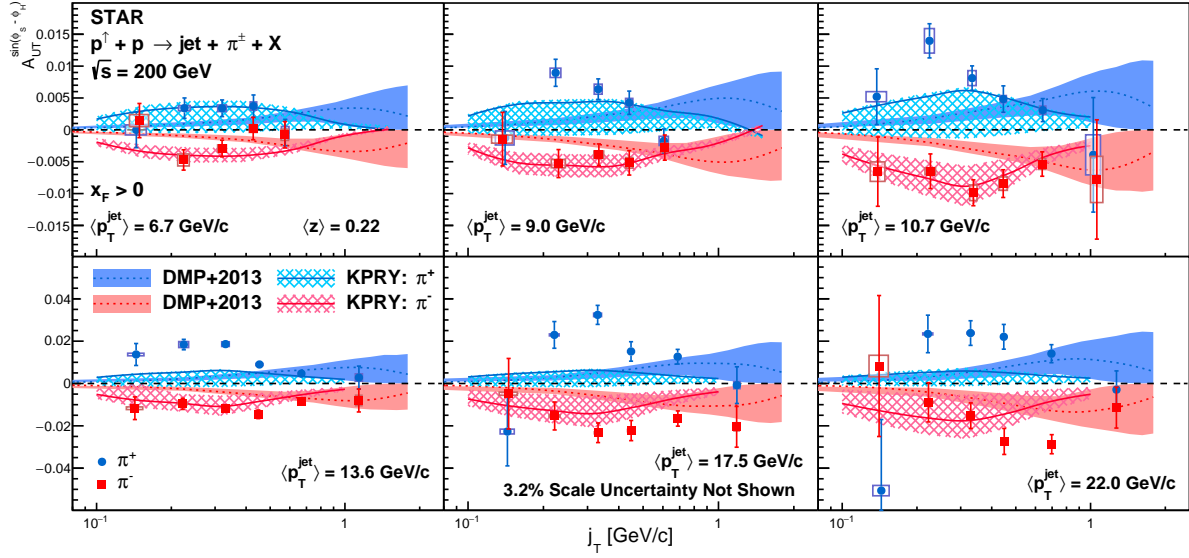


Figure 34: Collins asymmetries, $A_{UT}^{\sin(\phi_S - \phi_H)}$, as a function of the charged pion's momentum transverse to the jet axis, j_T , for in different jet p_T bins. The bars show the statistical uncertainties, while the size of the boxes represent the systematic uncertainties. The asymmetries are shown in comparison to calculations with the DMP+2013 model from Ref. [74] and the KPRY model from Ref. [75].

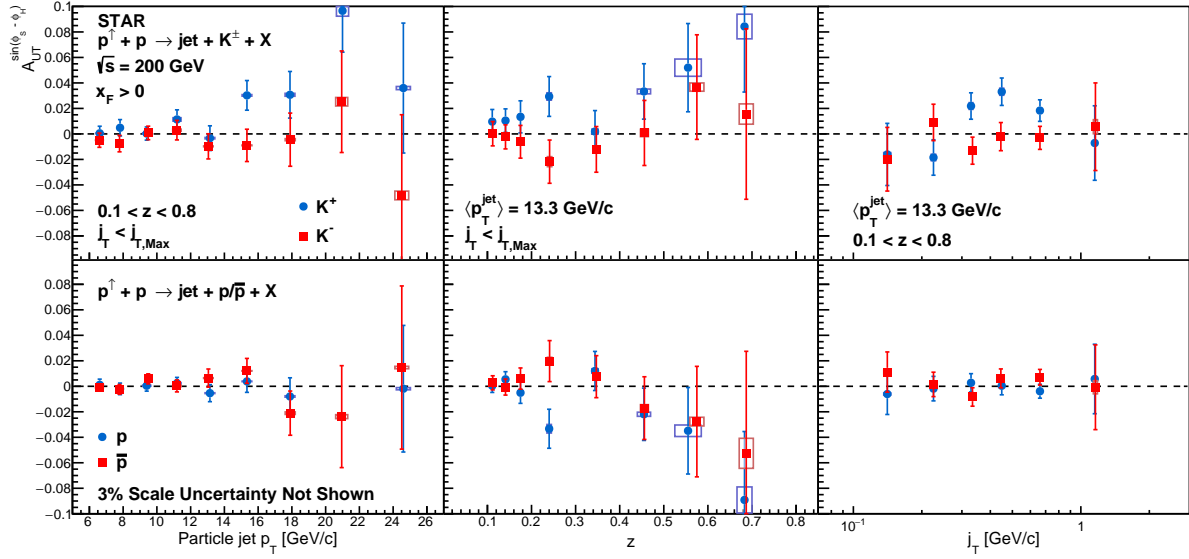


Figure 35: Collins asymmetries, $A_{UT}^{\sin(\phi_S - \phi_H)}$, as a function of particle jet- p_T , hadron- z , and hadron- j_T for charged kaons (upper panels) and protons (lower panels) inside jets. In both cases, the p_T dependence is shown integrated over the full ranges of z and j_T , while the z and j_T dependences are shown integrated over detector jet- $p_T > 9.9$ GeV/c. The bars show the statistical uncertainties, while the size of the boxes represent the systematic uncertainties

812 107% of the forward goal and 98% of the mid-rapidity goal.

813 Let us start by reviewing the goals and request for Run 22. The run was planned for
814 20 cyro-weeks. These weeks included time for cool-down and warm-up, sixteen days of CeC
815 running, and the remaining time for the STAR physics program. The specific requests for
816 the STAR physics program were:

- 817 • Sampled luminosity of 400 pb^{-1} . This was achieved by April 6th.
- 818 • Luminosity leveling for a maximum ZDC rate of 330 kHz. The leveling worked well,
819 especially with the addition of a second beta squeeze to maintain luminosity through
820 the end of the stores.
- 821 • A peak luminosity of $135 \times 10^{30} \text{ cm}^{-2} \text{ s}^{-1}$. This was achieved in early February.
- 822 • A polarization of 55% in both beams. This goal turned out to be extremely challenging
823 due to the loss of two coils in the Siberian Snake and the loss of the Siemens Motor
824 Generator from January 12th to March 8th. Despite these challenges, polarizations
825 close to 55% were achieved for the final six weeks of the run.
- 826 • Spin pattern and abort gaps similar to those for Run 17.
- 827 • Commissioning of the Forward Upgrade Detectors. This was expected to take place in
828 the initial two weeks of the run with beam, however the start-up of operations was de-
829 layed due to the cyro-system upgrades. STAR was able to complete the commissioning
830 using cosmics during this initial period when beams were not available.
- 831 • A few special runs were required with low luminosity and a small number of bunches
832 for calibrations. These runs were completed early in the run, before RHIC achieved
833 peak luminosities.
- 834 • Optimized time sharing with CeC. This was efficiently planned and executed to mini-
835 mize the impacts on the STAR physics program.

836 From the STAR operation point of view, the run started on schedule. The Forward
837 Upgrade Detector systems were all installed on schedule and ready for the start of the run.
838 The STAR magnet power supply heat runs were conducted from November 5th to 9th,
839 verifying that the STAR magnet was ready for operations. STAR started two-person shift
840 crews in November 9th when gas was introduced to the TPC. Initial cosmic ray data taking
841 to test the detector systems was started on November 11th. The full four-person (plus
842 trainees) were in place on November 16th and STAR was ready to take data. Much credit
843 should go to the shift coordination as STAR is still operating under COVID precautions and
844 many international institutions were unable to travel to the US.

845 Unfortunately, beam operations did not start as expected due to a delay caused by
846 the RHIC cyro-systems upgrade. STAR made use of this period of time without beam to
847 commission the Forward Upgrade Detectors using cosmic rays. Although commissioning with
848 cosmics was less efficient than commissioning with beam there was ample time to complete

849 the process. Beams were first injected into the blue ring on December 3rd and into the yellow
850 ring on December 7th.

851 The start of RHIC operations was further effected by two significant power dips. The first
852 on December 3rd and the second on December 12th. After the first power dip, which was an
853 86 second long site-wide power dip, a superconducting helical coil in the blue Siberian Snake
854 was found to be damaged. The second power dip damaged a second coil in the snake. The
855 snakes are essential for maintaining the polarization of the beams. Credit must go to the
856 CAD experts who were able to determine how to operate the damaged snake, which allowed
857 the run to go forward, however initially the polarizations were achievable were only 45% in
858 both rings, and as the figure of merit is polarization squared times integrated luminosity, the
859 reduction in polarization significantly impacted the ability to achieve the physics goals. On
860 December 18th, physics running started. This was almost one month behind the expectation.

861 Another major set-back occurred on January 12th when the Siemens Motor Generator
862 failed. CAD was able to quickly switch to the Westinghouse, however polarizations dropped
863 to 40%. At the end of January, STAR was projecting to only reach 30% on our goals.

864 In early February, due to improvements in injection and optimization, RHIC was able
865 to achieve the luminosity goals, however the polarization was still low. This improved on
866 March 8th when the repaired Siemens Motor generator was out back online. With the return
867 of the Siemens and much optimization, the polarizations in both beams finally reached 55%
868 and remained at that level for the remainder of the run.

869 On March 8th, the run was extended by an additional two weeks. The new end date for
870 beam operations was scheduled for April 18th. The sampled luminosity goal of 400 pb^{-1}
871 was achieved on April 6th. However due to the reduced polarizations at the start of the run,
872 the figure of merit polarization squared times sampled luminosity goal was delayed. At the
873 end of beam operations, STAR had achieved 98% of the figure of merit goal (as is shown in
874 Fig. 36, which is quite remarkable considering the challenges which needed to be overcome.

875 Throughout the run period, STAR operations achieved the expected up-time of twelve
876 hours per day of data taking.

877 Tremendous credit must go to CAD for overcoming the series of challenges. These chal-
878 lenges caused an initial delay of almost a month and reduced polarizations for the first half of
879 the run. STAR was able to commission the Forward Upgrades with cosmics which allowed
880 us to start taking data as soon as beams were available. The two week extension to the run
881 was essential. By April, everything was running extremely well.

882 1.3.1 Forward Upgrade

883 The forward upgrade consists of four major new subsystems an electromagnetic and hadronic
884 calorimeter and a tracking system, including a silicon and a small-strip Thin Gap Cham-
885 bers tracking detector. The calorimeter subsystems were fully installed, instrumented, and
886 commissioned during the 2021 RHIC running period. The tracking detectors were installed
887 in summer and fall 2021, on schedule and ready for the start of Run 22. All the systems
888 were further commissioned at the beginning of the run. They performed well and took data
889 smoothly throughout Run 22.

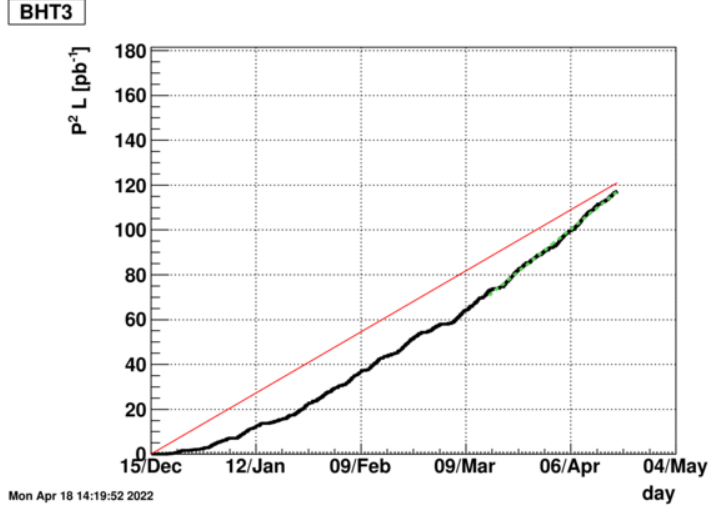


Figure 36: The figure of merit polarization squared times integrated luminosity as a function of date. The red line represents the rate necessary to achieve the physics goals. The black line displays the actual accumulation of the data. The trigger which is displayed uses a 3 GeV signal in a single barrel calorimeter tower (BHT3).

890 **FCS Run-22 Summary** Forward Calorimeter System (FCS) consists of Electro-Magnetic
 891 Calorimeter (Ecal) with 1486 towers, and Hadronic Calorimeters (HCal) with 520 towers.
 892 Ecal was installed in 2019 and Hcal was installed in 2020, both on the west platform at STAR.
 893 All SiPM sensors, front-end electronics boards and readout & triggering boards called DEP
 894 were installed, commissioned and calibrated during Run-21. Signal splitter boards for west
 895 EPD detector were installed before Run-22 and West EPD was used as pre-shower detector
 896 in electron triggers. FPGA code for FCS triggers was developed in fall 2021, and total of 29
 897 triggers, including triggers for di-electron (J/Ψ and DY), jets, hadrons, and photons were
 898 commissioned and verified within few days after RICH started delivering stable pp collisions
 899 and then used for data taking throughout Run-22 successfully. Calibration of Ecal was
 900 quickly done with reconstructing π^0 , and calibration of Hcal was done by MIP peak from
 901 $<1\%$ of hadrons from pp collisions which did not start hadronic shower in Hcal, together
 902 with cosmic muon signals with Hcal module oriented vertically outside STAR. FCS operation
 903 during Run-22 was successful and smooth, besides 3 LVPS modules needed to be replaced,
 904 and occasional power cycling of electronics were needed due to beam related radiation upsets
 905 in the electronics. All 1486 channels of Ecal were working without any bad channels, and
 906 Hcal had only a couple of dead channels. Radiation damage to the SiPM sensors due to beam
 907 was within the expectation. There was unexpected loss of signal amplitudes of 20% per
 908 week in Ecal near beam, which turned out to be radiation damage in the front-end electronics
 909 boards. The loss of signal was compensated during Run-22 by changing gain factor on DEP
 910 boards, attenuator setting in the front-end electronics, and raising voltage settings tower by
 911 tower based on LED signals. Details on the radiation damage on the front-end electronics
 912 are currently under investigation.

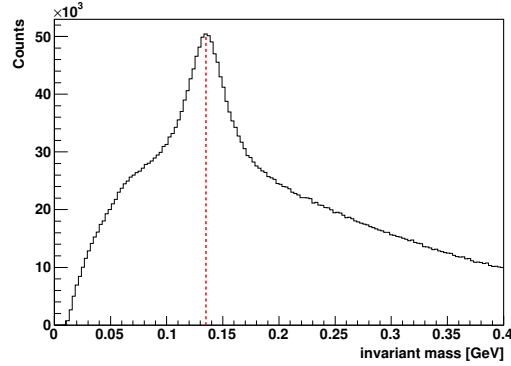


Figure 37: Invariant mass distribution and π^0 peak reconstructed with Ecal from pp collision data taken during Run-22.

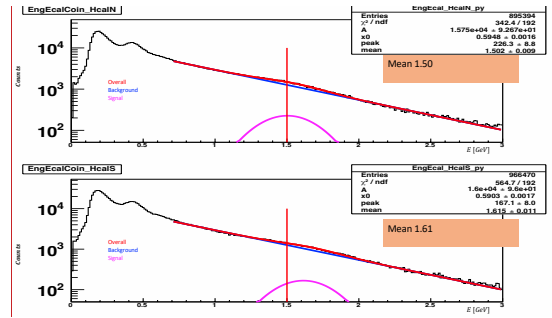


Figure 38: MIP peak in Hcal from $< 1\%$ of charged hadrons which did not start hadronic shower in Hcal from pp collision data taken during Run-22

913 **sTGC Run-22 Summary** The sTGC has four identical planes, each plane has four identical
 914 pentagonal shaped gas chambers. These gas chambers are made of double-sided and
 915 diagonal strips that give x, y, u in each plane. Sixteen chambers and about 5 spare chambers
 916 are built at Shandong University in China. Custom designed and fabricated aluminum frame
 917 allowed to fit the detector inside the pole-tip of the STAR magnet and around the beam-pipe
 918 on the west side of the STAR.

919 The sTGC chambers are operated with a quenching gas mixture of n -Pentane and CO_2 at
 920 a ratio of 45%:55% by volume at a typical high voltage of 2900 V. This gas mixture allowed
 921 the chambers to operate at high amplification mode. This mixture forms a flammable gas
 922 and the n -Pentane is liquid at normal atmospheric pressure and temperature. This made
 923 building the gas mixing systems extremely challenging. The supply chain issue caused by the
 924 pandemic added another layer of difficulty in completing the gas system. Allowable maximum
 925 pressure tolerance for the sTGC chambers are about 4 mBar above the atmospheric pressure
 926 and gas flow rate is extremely low, about 50 cc/min per chamber. In house, a newly designed

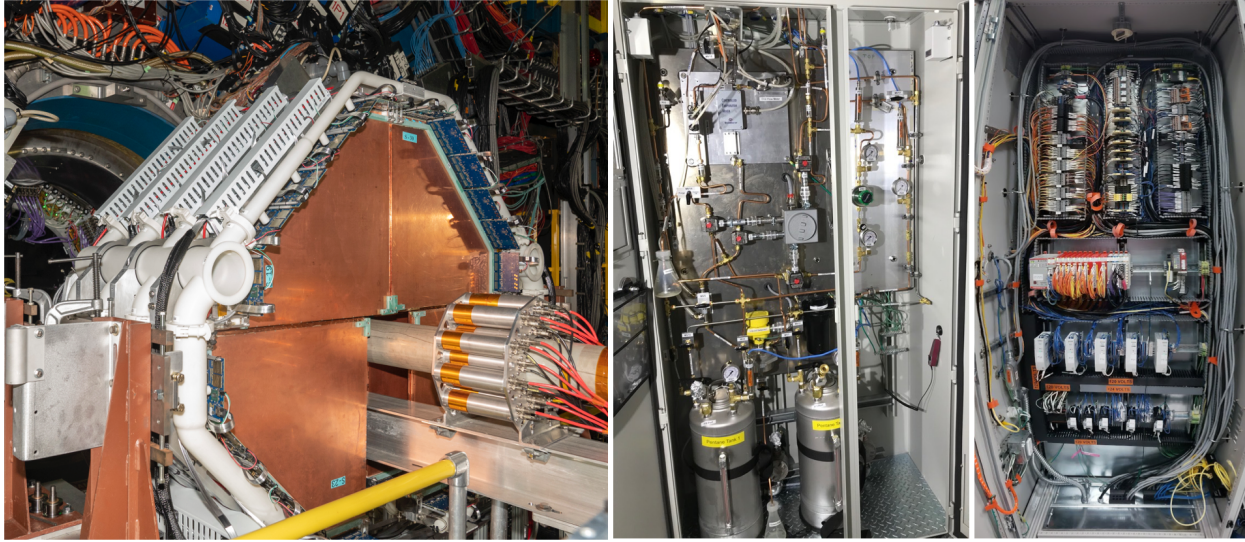


Figure 39: Left: sTGC detector module on a temporary platform prior to insertion into the pole-tip, middle: sTGC gas mixing system cabinet, right: sTGC interlock system cabinet.

927 and built gas system for mixing, and supplying the gas along a long-heated path to deliver
 928 to the chambers, met above requirements, and performed exceptionally well during the Run-
 929 22. Specially with many storms and power outages during the run, these systems performed
 930 uninterrupted. Added independent binary gas analyser during Run22 ensures that the gas
 931 mixture is at the right ratio. Since the gas mixture is flammable and liquefaction is possible
 932 inside the gas tubing, an independent redundant interlock system was designed and developed
 933 according to the industry standards. This system places the gas system in a safe state
 934 during any unforeseen situation such as flammable gas leak, fire, pentane liquefaction or
 935 over pressure occurs inside the chambers. Left panel of Fig. 39 shows the sTGC detector
 936 module on a temporary platform built prior to insertion into the pole-tip. Middle and right
 937 pictures show the gas mixing system and the interlock system respectively.

938 The sTGC readout is based on ATLAS VMM chips designed for ATLAS sTGCs. FEE
 939 cards were directly mounted on the edge of the sTGC chambers. This location is subjected to
 940 high radiation and magnetic field but the FEE cards performed exceptionally well during the
 941 operation. The sTGC was fully installed prior to the start of Run-22, and the detector was
 942 fully commissioned during the first few weeks of the run. Operating point of the high voltage
 943 was scanned for optimum efficiency. Gas chambers were stable at the desired operational
 944 high voltage and at the high luminosity, also the leakage current is well within the operational
 945 limits. sTGC has exceeded designed hit efficiency of 97% with the VMM chip gain set 3
 946 mV/pC and the high voltage 2900V. During the running four chambers were lost, the reason
 947 for losing the chambers are still known. But, during initial training of the chambers, these
 948 four chambers performed poorly compared to the rest. These chambers will get replaced for
 949 the next run.

950 **FST Run-22 Summary** Forward Silicon Tracker (FST) consists three identical disks
951 and each disk contains 12 modules. Each module has 3 single-sided double-metal Silicon
952 mini-strips sensors and readout by 8 APV chips. The installation of FST was completed on
953 August 13th, 2021 and the first collision p+p 510 GeV collision data recorded on December
954 15, 2021. The FST has been running smoothly through the whole run22 and the detector
955 operation via slow control software remains minimum to the shift crew.

956 To find the optimal operation high voltage, a voltage scan was performed with low lu-
957 minosity runs on December 17th, 2021. The operation high voltage decided to be 140V and
958 160V for inner and outer silicon sensors separately. The FST was running with 9 time bins
959 initially for the detector commissioning and tuned to 3 time bins on December 21st, 2021 to
960 increase the maximum DAQ rates of FST to 4.5kHz.

961 The noise level of FST silicon sensors is 10 to 20 ADC depending on position of the
962 silicon strip and the average signal to noise ratio is about 25. Due to irradiation damage, the
963 leakage current of silicon sensors increased from 2 uA to around 10uA (inner silicon sensor)
964 and 15uA (outer silicon sensor) after 4 months of p+p 510 GeV data taking. This increase
965 is consistent with expectation. There are 2 inner sectors and 2 outer sectors were operating
966 at a lower high voltage value due to abnormal bias current behavior. Those modules will be
967 investigated during the shutdown.

968 The FST readout chips are kept at room temperature by the cooling crate (same crate
969 also used by Intermediate Silicon Tracker) running 3M NOVEC. The leak rate of the whole
970 cooling system increased from 0.6% per day to 0.9% per day at the end of run22. The coolant
971 tank were refilled every 6 weeks by expert.

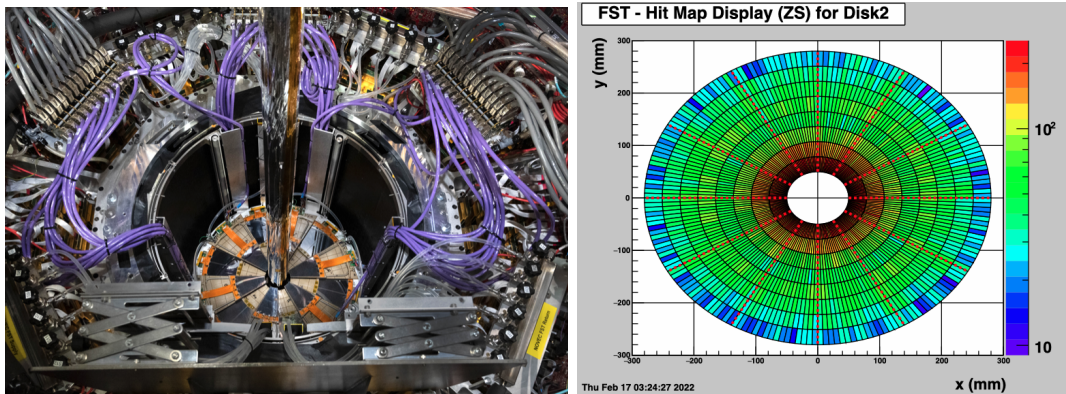


Figure 40: Left: FST after installation; Right: event display for p+p 510 GeV collisions.

972 **Software and Tracking**

2 Run-23 and Run-25 Requests for Au+Au Collisions at 200 GeV

2.1 Explore the Microstructure of the QGP

The completion of RHIC's scientific mission involves the two central goals: (i) mapping out the phase diagram of the QCD, and (ii) understanding the inner workings of the QGP by resolving its properties at varying length scales [76]. The former goal is addressed by the BES-II/FXT program. For the latter goal, the complementarity of the RHIC and LHC facilities is scientifically as essential as is having more than one experiment independently study the microstructure of the QGP at RHIC. With several years of operating the iTPC upgrade and commissioning and operation of the forward detectors in Run 22, the STAR collaboration is in an excellent position to take advantage of its vastly improved detection capabilities. Combining this with the prospect of a substantial increase in beam luminosities, RHIC will be uniquely positioned to fully engage in a detailed exploration of the QGP's microstructure.

Through careful discussions in its physics working groups, the STAR collaboration has identified a number of topics that together with the expected sPHENIX results in 23-25 make up a comprehensive study of the QGP microstructure, and successfully complete RHIC's scientific mission. In this section, we present a selection of those topics that will take full advantage of both STAR and RHIC's unique capabilities and address the following important questions about the inner workings of the QGP. We enumerate questions below that follow the chronology of an event; from questions addressing the QCD vacuum and the initial conditions, to the formation, temperature, and properties of the QGP, to the quenching of jets in said QGP, to its phase transition back to hadronic matter, and finally to the interactions of those final state hadrons.

1. **What is the nature of the 3-dimensional initial state at RHIC energies?** How does a twist of the event shape break longitudinal boost invariance and decorrelate the direction of an event plane? Can the v_1 of the J/ψ tell us about the initial tilt angle of the source? Can the Wigner distributions of photon tell us about the magnetic field effects in the initial state?
2. **What is the precise temperature dependence of the shear η/s , and bulk ζ/s viscosity?** Can combining precision flow results with those from other energies can help determine the temperature dependence of the viscosity.
3. **What can we learn about confinement from charmonium measurements?** Can the elliptic flow of J/ψ tell us the charmed quarks are deconfined?
4. **What is the temperature of the medium?** Do the Υ and $\psi(2s)$ melt at RHIC energies, and if so can their suppression be used to determine the temperature of the QGP? The thermally produced di-leptons are also produced in the plasma. Does their temperature agree with that found via quarkonium suppression?

- 1010 **5. What are the electrical, magnetic, and chiral properties of the medium?**
 1011 How is global vorticity transferred to the spin angular momentum of particles on such
 1012 short time scales? And, how can the global polarization of hyperons be reconciled with
 1013 the spin alignment of vector mesons? Can dilepton production in the low mass region
 1014 tell us about the electrical conductivity of the plasma? Can clear observation of the
 1015 ρ^0 - a_1 mixing tell us about the degrees of freedom therefore the chirality of the plasma?
 1016 Is there local parity violation and chiral magnetic effect?
- 1017 **6. What are the underlying mechanisms of jet quenching at RHIC energies?**
 1018 What do jet probes tell us about the microscopic structure of the QGP as a function
 1019 of resolution scale?
- 1020 **7. What is the precise nature of the transition near $\mu_B = 0$?** Where does the
 1021 sign-change of the susceptibility ratio χ_6^B/χ_2^B take place?
- 1022 **8. What can we learn about the strong interaction?** Can correlation functions
 1023 between baryons emitted at the surface of the fireball tell us how they interact in free
 1024 space.

1025 The event statistics projections that are used in this section will rely on the CAD's 2023E
 1026 and 2025E Au+Au luminosities [77] and the improved iTPC readout speed, and are listed
 1027 in Table 5. For each year we presume 24 weeks of physics data taking, and based on past
 1028 run operations an overall average of 85% \times 60% (STAR \times RHIC) uptime, respectively.

1029 It was realized that it will be possible to improve the readout speed of the iTPC detector
 1030 as deployed in BES-II, to a substantial higher rate for the runs 23-25 program. The upgrade
 1031 is primarily firmware and software development. It consists of the following components:

- 1032 • Rewrite the FPGA firmware for FEEs and RDOs. The FPGAs are different for the
 1033 outer sectors (TPX) and inner sectors (iTPC)
- 1034 • Rewrite DAQ online software for TPC in framework as for FCS
- 1035 • Redo and evaluate cluster finder
- 1036 • Improve network connectivity
- 1037 • Add some DAQ PC and event builders to handle increased data volume.
- 1038 • The original gating grid driver that had a limit of 2.2 kHz was replaced for Run 22
 1039 and can now easily handle more than 5 kHz.

1040 The expectation is that data rate can be approximately doubled with nominal downtime.
 1041 Thus:

- 1042 • Minimum Bias data taken at low luminosity should be able to record 5 kHz with 30%
 1043 downtime.

1044
1045

- High luminosity data for rare triggers should be able to be recorded at 3 kHz at 20%
deadtime.

1046
1047
1048
1049

The coding has already begun and is being developed and tested on the actual hardware using one of the TPC sectors. Performance is being evaluated using actual Au+Au low luminosity data from Run 19. Progress is good and expect that the development and system testing will be completed by end of the year.

1050
1051
1052
1053
1054
1055
1056
1057
1058
1059
1060
1061
1062
1063
1064
1065

In order to achieve a balance between those physics observables which are acquired with a minimum bias trigger (and negatively impacted by excess tracks in the TPC) and the rare probes which require specialized triggers (high tower (HT), dimuon) and the highest luminosity which can be accommodated with the TPC, the collaboration will optimize the interaction rates at STAR by allocating high and low luminosity periods within fills. CAD can offset the beam to independently control the maximum luminosity in each IR. Such periods, in which low interaction rates (specialized triggers) are sampled in the early part of a fill and high interaction rates (min bias trigger) typically in the later part, will allow us to collect clean, low pile-up, minimum bias events, while at the same time not burn beam luminosities that could affect interaction rates for sPHENIX. Clean minimum bias events will improve tracking efficiencies which in turn are expected to benefit many of the proposed correlation analyses. Optimization of the available bandwidth for rare triggers would allow us to push for lower p_T thresholds, thus further reducing biases. The impact of such an optimization will lead to some reduction in the projected rates, while still enabling a significant improvement in the precision and kinematic reach of current STAR measurements, and making important measurements that are yet more differential possible.

year	minimum bias [$\times 10^9$ events]	high- p_T int. luminosity [nb^{-1}]		
		all vz	$ vz < 70\text{cm}$	$ vz < 30\text{cm}$
2014	2	27	19	16
2016				
2023	20	40	36	24
2025				

Table 5: STAR minimum bias event statistics and high- p_T luminosity projections for the 2023 and 2025 Au+Au runs. For comparison the 2014/2016 event statistics and luminosities are listed as well.

1066
1067
1068
1069
1070
1071
1072
1073
1074

It is possible to build detectors that can span from mid-rapidity to beam rapidity – with the BES-II upgrades and the recent Forward upgrade STAR is able to achieve this unique capability. STAR’s BES-II upgrade sub-systems comprised of the inner Time Projection Chamber (iTPC, $1.0 < |\eta| < 1.5$), endcap Time Of Flight (eTOF, $1 < \eta < 1.5$) and Event Plane Detectors (EPDs, $2.1 < |\eta| < 5.1$), that are all fully operational since the beginning of 2019 [7, 78, 79]. The STAR Collaboration has commissioned and operated a forward rapidity ($2.5 < \eta < 4$) upgrade that includes charged particle tracking and electromagnetic/hadronic calorimetry [80]. Charged particle tracking is achieved using a combination of silicon detectors and small strip thin gap chamber detectors. The combination of these two tracking

1075 detectors is referred to as the forward tracking system (FTS). The FTS is capable of dis-
 1076 criminating the hadron charge sign. It can measure p_T of charged particles in the range of
 1077 $0.2 < p_T < 2$ GeV/ c with 20 – 30% momentum resolution.

1078 In what follows, we will refer to the combination of the existing TPC ($|\eta| < 1$) and the
 1079 iTPC upgrade as iTPC ($|\eta| < 1.5$) for simplicity.

1080 The impetus for running STAR during runs 23 and 25 in Au+Au 200 GeV collisions
 1081 comes from gains via: i) extended acceptance, ii) enhanced statistics, and iii) low material
 1082 budget. The extended acceptance is important for analyses that probe the η dependencies
 1083 and especially so for those that require correlations between particles (CME, $v_2(\eta)$, $r_n(\eta)$,
 1084 and $P_H(\eta)$). The enhanced statistics through longer running time and higher luminosities is
 1085 especially important for the rare probes (jets, J/ψ , CME, net-p C_6). In the previous 200 GeV
 1086 runs in 2014-2016 STAR included inner silicon detectors (the Heavy Flavor Tracker). This
 1087 has since been removed and by comparison in run 23-25 STAR will have a reduced material
 1088 budget between the beam and the iTPC, and will be uniquely positioned to perform dielec-
 1089 tron measurements. With these measurements, we propose to study the initial conditions
 1090 (Wigner functions, photoproduction of J/ψ), the degrees of freedom of the medium (excess
 1091 yield), and its transport properties (temperature through slope in the IMR).

1092 A synopsis of the proposed analyses, which questions they address, whether they will be
 1093 part of the minimum bias (low luminosity) or specialized trigger (high luminosity) program,
 1094 which coverage is essential, and the required trigger is shown in Fig. 41.

1095 The following subsections will address the specific analyses which are proposed to answer
 1096 the questions outlined previously in the section. The questions sequentially step through the
 1097 chronology of an event.

1098 What is the nature of the 3D initial state?

1099 Pseudorapidity-dependent Azimuthal Correlations to Constrain the Longitudi- 1100 nal Structure of the Initial State ($v_n(\eta)$)

1101 Initial-state longitudinal fluctuations and the fluid dynamical response of the medium formed
 1102 in heavy ion collisions can lead to de-correlations of the direction of the reaction planes Ψ_n
 1103 (which determines the orientation of the harmonic anisotropies) with pseudorapidity (see
 1104 Fig. 42). Such effects are often referred to as a torque or twist of the event shape [19, 81, 82]
 1105 that eventually leads to a breaking of longitudinal/boost/rapidity invariance. The magni-
 1106 tude of the de-correlation is determined by the details of the dynamics of initial state, and
 1107 the distribution of nucleons and partons inside the colliding nuclei.

1108 Several promising observables have been proposed to study this effect, Fig. 42 shows one
 1109 which can be expressed as $r_n(\eta_a, \eta_b) = V_{n\Delta}(-\eta_a, \eta_b)/V_{n\Delta}(\eta_a, \eta_b)$, where $V_{n\Delta}$ is the Fourier
 1110 coefficient calculated with pairs of particles taken from three different pseudorapidity re-
 1111 gions $-\eta_a$, η_a and η_b . The observable $r_n(\eta_a, \eta_b)$ was originally introduced and measured
 1112 by CMS collaboration in Ref. [84] and also been measured by the ATLAS collaboration
 1113 in [85]. An observable using three-particle correlations that is sensitive to this effect is
 1114 the relative pseudorapidity dependence of the three-particle correlator $C_{m,n,m+n}(\eta_a, \eta_b, \eta_c) =$
 1115 $\langle \cos(m\phi_1(\eta_a) + n\phi_2(\eta_b) - (m+n)\phi_3(\eta_c)) \rangle$ [86]. Another, very similar to r_n in terms of design

Observable	Question	PWG	MB/H \mathcal{L}	Coverage	Trigger
$v_2(\eta)$ Twist	1) Initial State	FCV	Min bias	iTPC, TOF, EPD, FTS	VPDMB
$r_n(\eta_a, \eta_b)$	1) Initial State	FCV	Min Bias	iTPC, TOF, EPD, FTS	VPDMB
$J/\psi v_1$	1) Initial State	HF	Luminosity	iTPC, TOF, EPD	Dimuon
Photon WF	1) Initial State	LFSUPC	Min Bias	iTPC, TOF	VPDMB
$v_2(\eta)$	2) Viscosity	FCV	Min bias	iTPC, TOF, EPD, FTS	VPDMB
$J/\psi v_2$	3) Deconfinement	HF	Luminosity	iTPC, TOF, EPD	Dimuon
Υ suppression	4) Temperature	HF	Luminosity	iTPC, TOF	Dimuon
$\psi(2s)$ suppress.	4) Temperature	HF	Luminosity	iTPC, TOF	Dimuon
Di-elec IMR	4) Temperature	LFSUPC	Min bias	iTPC, TOF	VPDMB
$P_H(\eta)$	5) Properties	FCV	Min Bias	iTPC, TOF, FTS	VPDMB
P_H of J/ψ	5) Properties	FCV	Luminosity	iTPC, TOF	dimuon
$\rho^0 a_1$ mixing	5) Properties	LFSUPC	Min Bias	iTPC, TOF	VPDMB
Di-elec LMR	5) Properties	LFSUPC	Min Bias	iTPC, TOF	VPDMB
CME	5) Properties	FCV	Min Bias	iTPC, TOF	VPDMB
$\gamma_{Dir} + \text{jet } I_{AA}$	6) Jet quenching	Jet Corr	Luminosity	Bcal, Ecal, Fcal	BHT3
$\gamma_{Dir} + \text{jet acopl.}$	6) Jet quenching	Jet Corr	Luminosity	Bcal, Ecal, Fcal	BHT3
Jet substruct.	6) Jet quenching	Jet Corr	Luminosity	Bcal, Ecal, Fcal	BHT3
Net-p C_6	7) Phase Transition	CF	Min Bias	iTPC, TOF	VPDMB
Baryon CF	8) Strong Interact.	CF	Min Bias	iTPC, TOF	VPDMB

Figure 41: A tabulation of the proposed analysis. The columns indicate which of the nine questions a given analysis addresses, which physics working group will lead the analysis effort, whether the analysis will be part of the low or high luminosity program, which detector systems are essential, and the required trigger for that analysis.

1116 but involving four-particle correlations, is: $R_{n,n|n,n}(\eta_a, \eta_b)$ [17]. As shown in Fig. 42, CMS
1117 measurements of r_n show strong de-correlation ($\sim 16\%$ for $n=3$, $\sim 8\%$ for $n=2$) in central
1118 events within the range of their acceptance. In the 3D-Glasma model of initial state, the
1119 breaking of boost invariance is determined by the QCD equations which predict the evolu-
1120 tion of gluons in the saturation regime with Bjorken-x. At the LHC such models predict
1121 weaker de-correlation as compared to when the initial state is described by wounded nucleon
1122 models. The 3D-Glasma model does a good job in explaining the r_2 data from CMS [18]
1123 but over-predicts the r_3 results. One expects the nature of the initial state to change from
1124 LHC to RHIC, in particular the region of Bjorken-x probed is very different. It is there-
1125 fore extremely important to utilize the enhanced acceptance of the STAR detector with a
1126 Au+Au 200 GeV run to study this effect. In Fig. 42 STAR's projections using preliminary
1127 Run-19 results to estimate the uncertainties for 10 B events are shown for the measurement
1128 of r_n within the acceptance $|\eta| < 1.5$. The colored regions show that the current and future
1129 capabilities at STAR (with iTPC+EPD+FTS) can extend such measurements using observ-

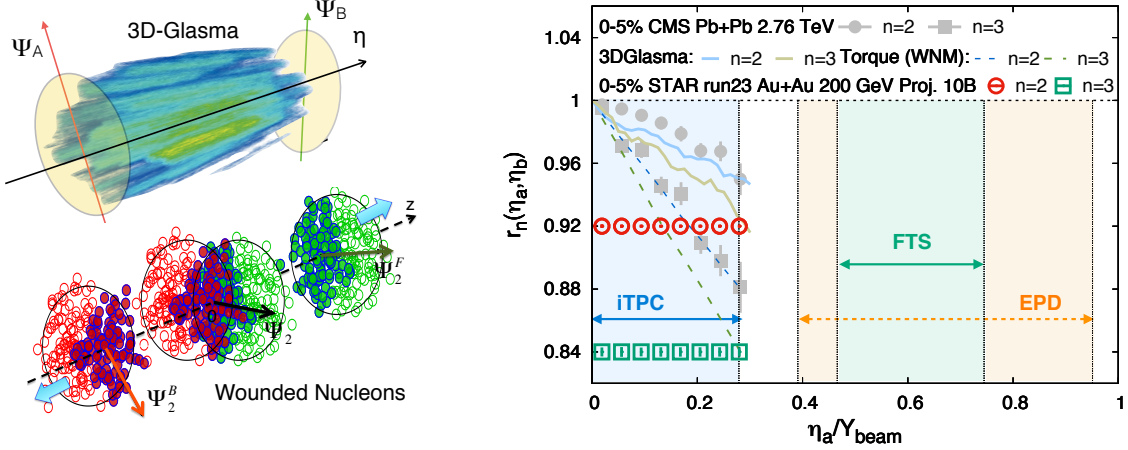


Figure 42: (Left) Cartoon to demonstrate the de-correlation of event planes in the longitudinal direction of a collision from a gluon saturation based 3D-Glasma model [18] and a wounded nucleon model (WNM) [19, 83]. (Right) The longitudinal de-correlation of the elliptic anisotropy plane as a function of pseudorapidity in units of beam rapidity. CMS results are compared to predictions from two models in the left with STAR projection for Run 23 (using preliminary Run 19 results) from an anticipated 10 B min-bias events. The colored regions show that the current and future capabilities at STAR (with iTPC+EPD+FTS) can extend such measurements with good precision by covering a large fraction of the beam rapidity at 200 GeV – this demonstrates the unique strength of STAR to study the physics of 3D initial state.

1130 ables $r_n, C_{m,n,m+n}, R_{n,n|n,n}$ with good precision by covering either an equal (iTPC only) or
 1131 larger (iTPC+FTS+EPDs) fraction of the beam rapidity at 200 GeV compared to the LHC
 1132 measurements. This unique measurement capability will help pin down the nature of the
 1133 3-D initial state of heavy ion collisions. It will also help constrain different models of QCD
 1134 that predict the rapidity (or Bjorken-x) dependence of valance quark and gluon distributions
 1135 inside colliding nuclei as has been demonstrated by theoretical calculations in Ref. [18, 20].

1136 J/ψ v_1 to Study the Initial Tilt

1137 Studies of the directed flow, v_1 , as a function of rapidity provide crucial information to un-
 1138 derstand the initial tilt of the medium produced in heavy-ion collision [87, 88]. Heavy quarks
 1139 are produced in the early stage of a heavy-ion collision and thus are of particular interest
 1140 for the medium initial asymmetry studies. STAR recently reported the first measurement of
 1141 D-meson v_1 in Au+Au collisions at 200 GeV where the magnitude of the heavy-flavor meson
 1142 v_1 is about 25 times larger than the v_1 for charged kaons. With the runs 23 and 25 data,
 1143 STAR would have a unique opportunity to study the v_1 of a bound $c\bar{c}$ state, the J/ψ mesons,
 1144 for which even larger directed flow can be expected [89]. In addition to STAR’s excellent
 1145 capability to reconstruct low- p_T J/ψ , the iTPC will improve the momentum resolution and
 1146 extend the pseudorapidity coverage. This will provide better precision for the slope extrac-
 1147 tion of the v_1 vs y measurement, that quantifies the strength of directed flow. The expected
 1148 precision of a J/ψ v_1 measurement vs p_T at STAR in runs 23 and 25, assuming 20 B MB

1149 events and 40 nb^{-1} of HT trigger sampled luminosity, in 0-80% central Au+Au collisions at
 1150 200 GeV is shown in Fig. 43.

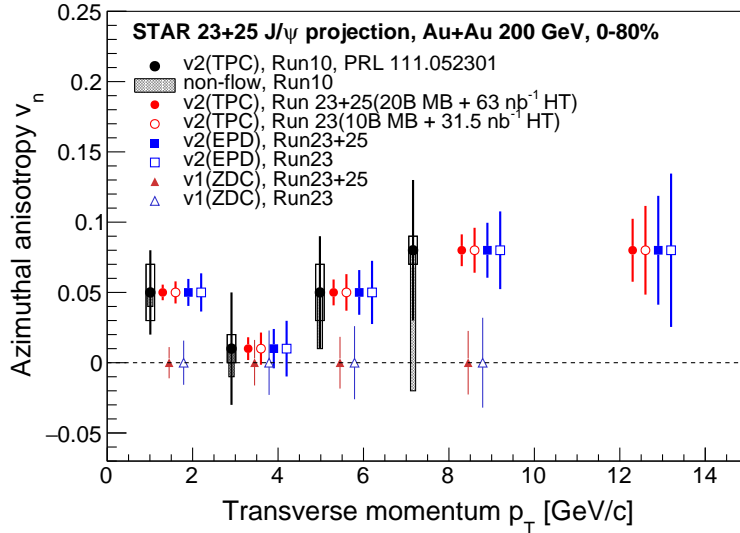


Figure 43: Projections for the J/ψ ($J/\psi \rightarrow e^+e^-$) directed (v_1) and elliptic (v_2) flow vs J/ψ p_T in 0-80% Au+Au collisions at 200 GeV, assuming 20 B MB events and 40 nb^{-1} of HT trigger sampled luminosity.

1151 Studying the Photon Wigner Function and Final-state Magnetic Fields in the 1152 QGP (photon WF)

1153 The unsuccessful description of STAR data by the STARLight model led to the attribution
 1154 of the broadening to the possible residual magnetic field trapped in an electrically conducting
 1155 QGP [90]; which is key information to the study of the chiral magnetic effect.

1156 Similarly, ATLAS quantified the effect via the acoplanarity of lepton pairs in contrast
 1157 to the measurements in UPC and explained the additional broadening by multiple electro-
 1158 magnetic scatterings in the hot and dense medium [91], which is analogous to the medium
 1159 P_\perp -broadening effects for jet quenching.

1160 These descriptions of the broadening in hadronic collisions are based on the assumption
 1161 that there is no impact parameter dependence of the p_T distribution for the electromagnetic
 1162 production. Recent lowest-order QED calculations, in which the impact parameter depen-
 1163 dence is recovered, could reasonably describe the broadening observed by STAR and ATLAS
 1164 without any in-medium effect. To solve the puzzle, we propose to precisely study the initial
 1165 P_\perp -broadening for the dilepton pair in ultra-peripheral collisions. Different neutron emission
 1166 tags serve as the centrality definition, and will allow us to explore the broadening baseline
 1167 variation with impact parameter. Furthermore, the differential spectrum as a function of
 1168 pair P_\perp , rapidity, and mass enable us to study the Wigner function of the initial electromag-
 1169 netic field, which provide the information to extract the momentum and space correlation
 1170 of EM field.

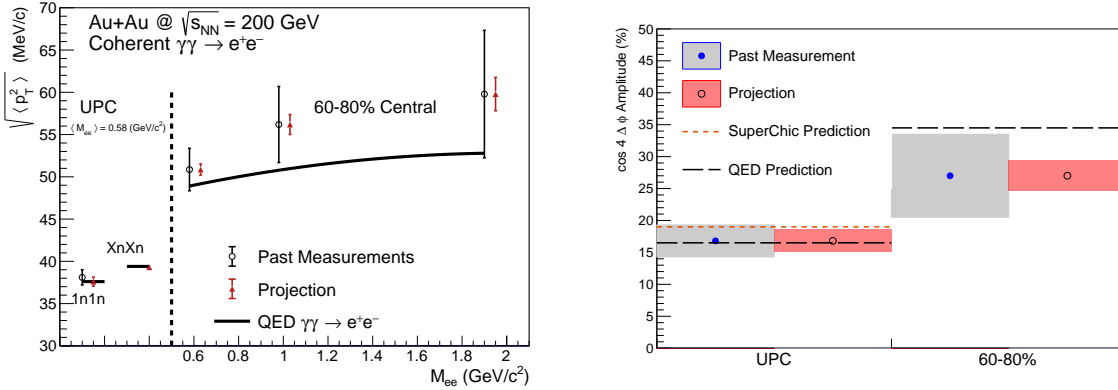


Figure 44: Projections for measurements of the $\gamma\gamma \rightarrow e^+e^-$ process in peripheral and ultra-peripheral collisions. Left: The $\sqrt{\langle p_T^2 \rangle}$ of di-electron pairs within the fiducial acceptance as a function of pair mass, M_{ee} , for 60–80% central and ultra-peripheral Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. Right: The projection of the $\cos 4\Delta\phi$ measurement for both peripheral (60–80%) and ultra-peripheral collisions.

1171 As shown in Fig. 44, comparing with the latest QED calculation, there still exists addi-
 1172 tional broadening in peripheral collisions, although the significance is only about 1σ , which
 1173 still leave room for the medium effect. In Runs 23 and 25, as projected in the figure, we
 1174 could judge the existence of additional broadening with much higher precision and further
 1175 constrain the strength of final-state magnetic field in the QGP.

1176 Precision measurement of the amplitude of the recently observed $\cos 4\Delta\phi$ modulation
 1177 of the $\gamma\gamma \rightarrow e^+e^-$ process will allow precision mapping of the photon Wigner function
 1178 and provide additional constraints on possible final-state effects, thereby complementing
 1179 the P_\perp broadening measurement. Figure 44 right panel shows the projected precision for a
 1180 measurement of the $\cos 4\Delta\phi$ modulation in Runs 23 and 25. The modulation is a direct result
 1181 of the mismatch in initial and final spin configuration of the $\gamma\gamma \rightarrow e^+e^-$ process. Any final-
 1182 state effect that modifies the P_\perp will necessarily reduce the amplitude of the modulation.
 1183 Assuming the same central value as previously measured, evidence for suppression of the
 1184 $\cos 4\Delta\phi$ modulation will be visible at the $> 3\sigma$ level (stat. & syst. uncertainty). Precision
 1185 measurement of the $\cos 4\Delta\phi$ modulation in Runs 23 and 25 may also allow a first direct
 1186 experimental measurement of the impact parameter dependence of this new observable (by
 1187 comparing UPC and 60–80%). Assuming the same central values as previously measured,
 1188 the improved precision will provide evidence for impact parameter dependence at the $> 3\sigma$
 1189 level (stat. & syst. uncertainty). Assuming the central value predicted by QED would lead
 1190 to a $> 5\sigma$ difference between the UPC case and the 60–80% case.

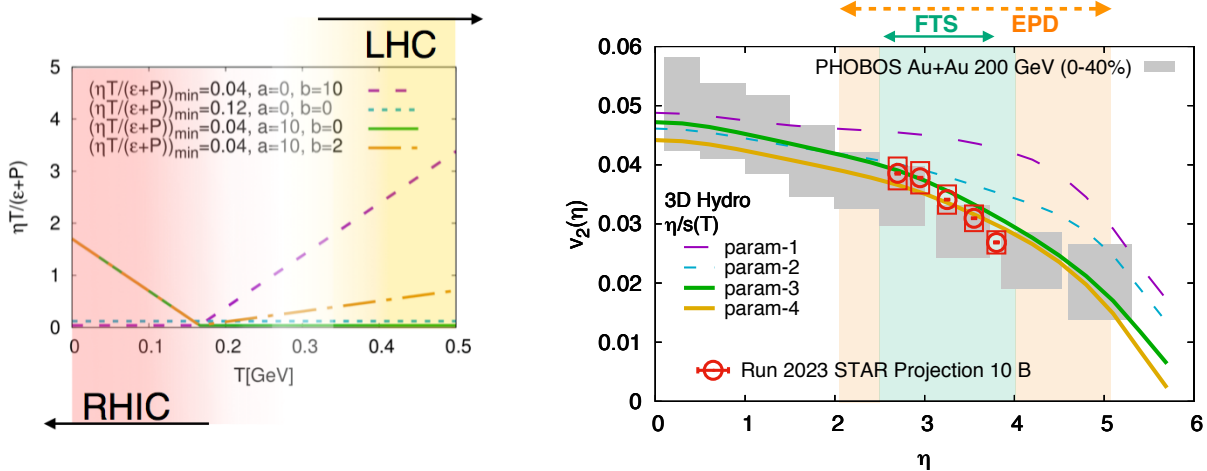


Figure 45: (Left) Different parameterizations of the temperature dependence of the shear viscosity to entropy η/s (T) (at $\mu_B = 0$) used in the hydrodynamical simulation of Ref. [95]. It has been demonstrated in Ref. [96] that the region of lowest η/s is the one that can be probed at RHIC. (Right) Effects on the elliptic flow co-efficient v_2 due to different parameterizations of the viscosity parameter, indicating better constraints on η/s (T) can only be performed by measurements at forward rapidities at RHIC. The interpretation of the existing PHOBOS data is limited by the large uncertainties. Projections for STAR measurements are shown on the same plot.

1191 What is the precise temperature dependence of viscosity?

1192 The idea of tightly constraining the temperature dependent viscosity of the QGP was envi-
 1193 sioned in the 2015 Long Range Plan for Nuclear Science [76]. The QCD matter formed at
 1194 RHIC shows nearly perfect fluidity characterized by the smallest viscosity to entropy ratio
 1195 η/s known in nature. One major aim is to perform precision measurements to constrain the
 1196 temperature dependence of the shear η/s (T) and bulk ζ/s (T) viscosities. Recent state-
 1197 of-the-art Bayesian analyses of flow and spectra data within sophisticated event-by-event
 1198 hydrodynamics models has show strong evidence for temperature dependence of η/s and
 1199 ζ/s [92–94], but the uncertainties are still quite large. On the other hand, hydrodynamic
 1200 simulations have demonstrated that since the temperature of the produced fireball varies
 1201 with rapidity, the measurement of the rapidity dependence of flow harmonics can provide
 1202 additional constraints on the η/s (T) and ζ/s (T) [95]. For this, RHIC measurements have
 1203 an advantage over the LHC since the smaller beam rapidity at RHIC provides stronger
 1204 variations of the temperature with rapidity. The beam energy scan at RHIC provides an
 1205 additional handle on temperature to map η/s (T), and ζ/s (T) over a wide range of tem-
 1206 peratures. Indeed, the hydrodynamic simulation of Ref. [95] indicates that η/s (T) at lower
 1207 temperatures, near its possible minimum ($T = T_c$), can be better constrained by RHIC mea-
 1208 surements. Results from such simulations are shown in Fig. 45. In this simulation, a number
 1209 of QCD-motivated parameterizations of the temperature dependence of the shear viscosity
 1210 were assumed, as shown in Fig. 45 (left).

Existing data from the PHOBOS collaboration suffer from large uncertainties, therefore only limited constraints on the temperature dependence of the transport parameters can be achieved. The BES-II upgrades and the FTS will provide precise estimations of different azimuthal correlation observables: $v_n(\eta)$ and other higher-order ($n > 2$) flow coefficients $v_n(\eta)$, its fluctuations $\sigma(v_n)/v_n$ that have never been measured at forward rapidity, are essential in terms of constraining η/s (T) near its possible minimum. These quantities previously measured at mid-rapidity are not enough for discriminating different parameterizations of η/s (T) as shown in the hydrodynamic simulation of Ref. [95]. While p_T integrated quantities at forward rapidity can constrain the shear viscosity, measurement of the p_T of particles at forward rapidity (i.e. FTS) is essential to constrain the bulk ζ/s – in particular the information of $\langle p_T \rangle$ is needed to constrain $\zeta/s(T)$. With the FTS it will be possible to measure the p_T dependence of v_n in Au+Au collisions in runs 23 and 25.

What can charmonium tell us about deconfinement?

The strong collectivity of the QGP is studied by measuring the azimuthal anisotropy of the produced particles in heavy-ion collisions. A positive elliptic flow coefficient (v_2) of the light flavor hadrons, and also D -mesons and electrons from heavy-flavor hadron decays are observed in Au+Au collisions at $\sqrt{s_{NN}} = 54.4$ and 200 GeV at RHIC. This corroborates that, like light-flavor, the charm quarks are (partially) thermalized and show collectivity in the QGP. On the other hand, the v_2 of heavier J/ψ reported by STAR based on the 2010 Au+Au 200 GeV data sample was found to be consistent with zero, albeit within large statistical uncertainties and systematic uncertainties due to non-flow effects. The precision of the measurement was also not enough to distinguish between theoretical model calculations that assume only primordial J/ψ production and ones that include additional J/ψ production via recombination. This calls for a larger sample of heavy-ion data at 200 GeV, as will be provided by RHIC in runs 23 and 25, in order to observe a possible non-zero J/ψ v_2 at RHIC energies and put more constraints on the J/ψ production models especially regarding its regeneration. Particularly important for these studies is STAR's potential to measure low p_T J/ψ with a very good precision. This excellent low- p_T performance at STAR can be achieved thanks to its low material budget and great particle identification capabilities.

The second order event plane will be reconstructed using the EPDs which will significantly decrease the contribution from the non-flow effects and consequently the measurement's systematic uncertainties. Also, an inverse of the EP resolution enters directly the J/ψ v_2 uncertainty calculation. Due to the use of the EPD, the resolution of the EP at forward rapidity for the J/ψ v_2 measurement at STAR will improve. Figure 43 presents statistical projections for the J/ψ v_2 measurement in 0-80% central Au+Au collisions assuming 20 B MB events and 40 nb^{-1} of HT trigger sampled luminosity. Both cases of the second order EP reconstruction, using the forward EPD and mid-rapidity TPC, are considered and shown. A significant improvement in the precision of the J/ψ v_2 can be seen across the experimentally accessible J/ψ p_T coverage. In addition, the larger dataset would allow to extend the measured p_T range beyond 10 GeV/ c .

1251 **What is the temperature of the medium?**

1252 **Υ Suppression**

1253 In the QGP, the confining potential of a heavy quark-antiquark pair is predicted to be
 1254 screened by the surrounding partons leading to the quarkonium dissociation. Within this
 1255 static picture, a quarkonium state dissociates if its size is larger than the Debye screening
 1256 length of the medium that is inversely proportional to the medium temperature. Conse-
 1257 quently, different quarkonium states, depending on their sizes, are expected to dissociate at
 1258 different temperatures, which is usually referred to as the quarkonium sequential suppres-
 1259 sion. Quarkonia are therefore considered excellent probes of the medium thermodynamic
 1260 properties. In particular, differences in the dissociation temperatures between $\Upsilon(1S)$, $\Upsilon(2S)$
 1261 and $\Upsilon(3S)$ states are larger compared to the charmonium states, providing a longer lever
 1262 arm. In addition, the regeneration contribution for bottomonia is expected to be negligibly
 1263 small at RHIC energies. Figure 46 presents statistical projections for $\Upsilon(1S)$ and $\Upsilon(2S)$ R_{AA}
 1264 as a function of p_T and N_{part} (centrality), compared to STAR's latest results from the 2011,
 1265 2014 and 2016 datasets. The projections are done combining the di-electron and di-muon
 1266 Υ decay channels and for an integrated luminosity of 40 nb^{-1} that corresponds to the runs
 1267 23 and 25 data samples. One can see a clear improvement of the statistical precision for
 1268 both Υ states. Due to the larger suppression of the $\Upsilon(3S)$ state, only an upper limit on the
 1269 R_{AA} , 0.29 at 99% confidence level, was obtained so far. With an integrated luminosity of 40
 1270 nb^{-1} we expect a precision of about 30% for $\Upsilon(3S)$ that may allow us to extract the $\Upsilon(3S)$
 1271 signal if the meson is not fully dissociated in the medium or significantly improve precision
 1272 of our upper limit. The requested luminosity is therefore crucial to obtain a full picture of
 1273 the bottomonium family suppression at the RHIC top energy.

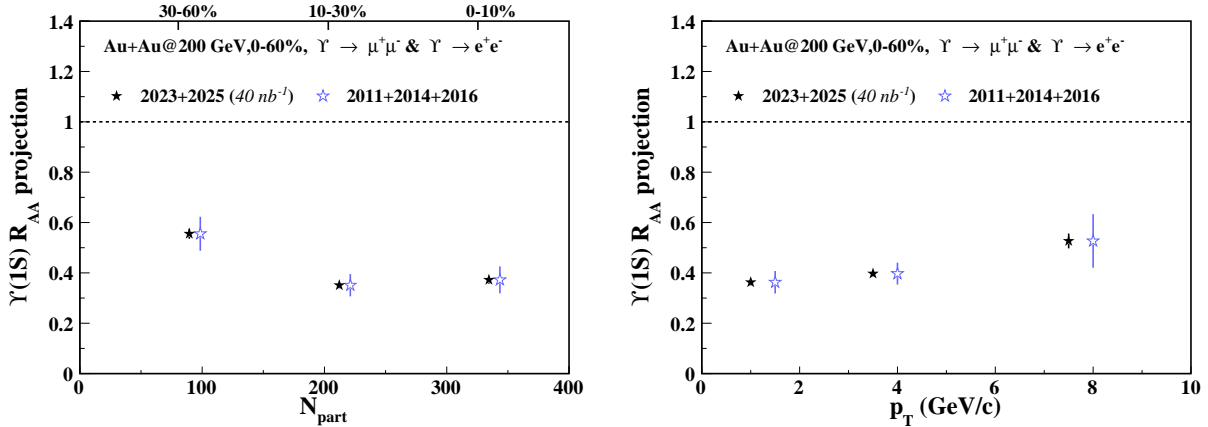


Figure 46: Statistical projections for the $\Upsilon(1S)$ and $\Upsilon(2S)$ R_{AA} as a function of N_{part} (left) and p_T (right) in 0-60% Au+Au collisions at 200 GeV, assuming 40 nb^{-1} of HT triggered events. The projections are done combining the di-electron and di-muon decay channels and are compared to the STAR results from 2011, 2014 and 2016 datasets. $\Upsilon(2S)$ to be added

1274 $\psi(2S)$ Suppression

1275 $\psi(2S)$ is the most loosely bounded quarkonium state currently accessible to heavy-ion col-
1276 lision experiments. Its dissociation temperature is predicted to be around, or below, the
1277 critical temperature, and is much less than that of J/ψ and Υ states. It is therefore more
1278 likely to be dissociated in the early stage and in the core of the fireball, and those $\psi(2S)$
1279 that are measured may have significant contributions from regeneration at a later stage in
1280 the evolution of the fireball. The relative suppression of $\psi(2S)$ and J/ψ is sensitive to the
1281 temperature profile of the fireball produced in heavy-ion collisions and its space-time evo-
1282 lution. It is also argued that the charmonium formation process from a $c\bar{c}$ pair may be
1283 affected by both the QGP and the initial strong external magnetic field, altering the relative
1284 yields among different charmonium states [97,98]. The measurement of $\psi(2S)$ is much more
1285 difficult than that of J/ψ due to a much smaller production cross-section and dilepton decay
1286 branching ratio, resulting in a very low signal-to-background ratio. The ALICE Collabora-
1287 tion successfully measured the relative suppression of $\psi(2S)$ and J/ψ in Pb+Pb collisions at
1288 forward rapidity [99], and the ATLAS and CMS Collaborations published the relative sup-
1289 pression in Pb+Pb collisions at mid-rapidity and high p_T [100,101]. Attempts to measure
1290 $\psi(2S)$ suppression in heavy-ion collisions at RHIC have not been successful to date. The low
1291 material budget and excellent particle identification capability of STAR together with the
1292 combined large data sample in runs 23 and 25 will provide a unique opportunity to measure
1293 the suppression of $\psi(2S)$ at low p_T and mid-rapidity in heavy-ion collisions. Figure 47 shows
1294 the projections of $\psi(2S)$ signal and the yield ratio of $\psi(2S)$ and J/ψ from 20 B MB events
1295 in Au+Au collisions. Here the $\psi(2S)/J/\psi$ ratio is assumed to be 0.02, and the performance
1296 of detectors from existing data before STAR iTPC upgrade is used for the projection. As
1297 shown in the figure, the $\psi(2S)$ signal significance will be around 3σ level in the 0-20% cen-
1298 trality bin. This significance could become even smaller depending on the level of further
1299 suppression for $\psi(2S)$ compared to J/ψ . Despite the improvement of momentum and dE/dx
1300 resolution thanks to the STAR iTPC upgrade, it is crucial to have both the runs 23 and 25
1301 data for a significant $\psi(2S)$ measurement.

1302 QGP Temperature from Di-lepton in the IMR

1303 The dilepton mass spectrum has many contributions. A cocktail of known processes is
1304 subtracted to find the excess radiation. To gain a deeper understanding of the microscopic
1305 origin of the excess radiation, we will

- 1306 • separate early from later time radiation by measuring dilepton elliptic flow (v_2) as a
1307 function of dilepton mass;
- 1308 • identify the source of dilepton radiation by studying dilepton polarization versus in-
1309 variant mass (helicity angle);
- 1310 • measure precisely the lifetime of the interacting fireball. As an observable we will use
1311 integrated low-mass yield but also compare explicit model calculations with various
1312 $\tau_{fireball}$;

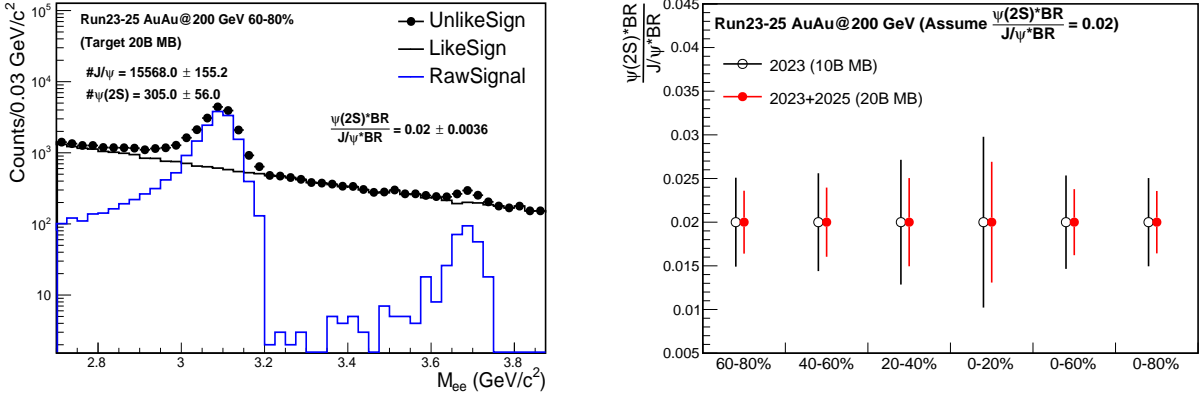


Figure 47: Projections for the J/ψ and $\psi(2S)$ signals in 60-80% Au+Au collisions at 200 GeV and the yield ratio in various centrality bins.

- extract an average radiating source temperature from the fit of a Boltzmann distribution to the invariant mass slope in the range 1.1 - 2.5 GeV/c^2 spectrum. The higher the invariant mass, the stronger the QGP contribution to the spectrum, the higher the chance to measure temperature of the QGP.

The di-lepton intermediate mass region, between the peaks from the decays of the ϕ and $J\psi$, is dominated by thermal emission from the QGP. The slope of the spectrum in this region can be used as a blue-shift free measurement of the temperature at the time of di-lepton emission. As was shown in the Highlight section, di-lepton IMR temperatures of 301 ± 60 and 338 ± 59 were found for the $\sqrt{s_{NN}} = 27$ and 54.4 GeV systems respectively. Extraction of a di-lepton temperature at $\sqrt{s_{NN}} = 200$ GeV will be directly comparable to the temperatures suggested by the Υ and $\psi(2s)$.

Last, but not least, concerning direct-photon emission, the existing difference, on the order of a factor of two, between the low momentum spectra from PHENIX and STAR in 200 GeV Au+Au collisions, has to be resolved. In order to clarify the direct photon puzzle we will measure with precision the direct virtual photon yield as well as its elliptic flow coefficient. We will particularly focus on low p_T η measurement which might be instrumental in clarifying this long standing question.

What are the electrical, magnetic, and chiral properties of the medium?

The QGP medium which is created during the collision of two heavy ions has significant electric fields, magnetic fields, vorticity, and chirality.

Pseudorapidity Dependence of Global Hyperon polarization ($P_H(\eta)$)

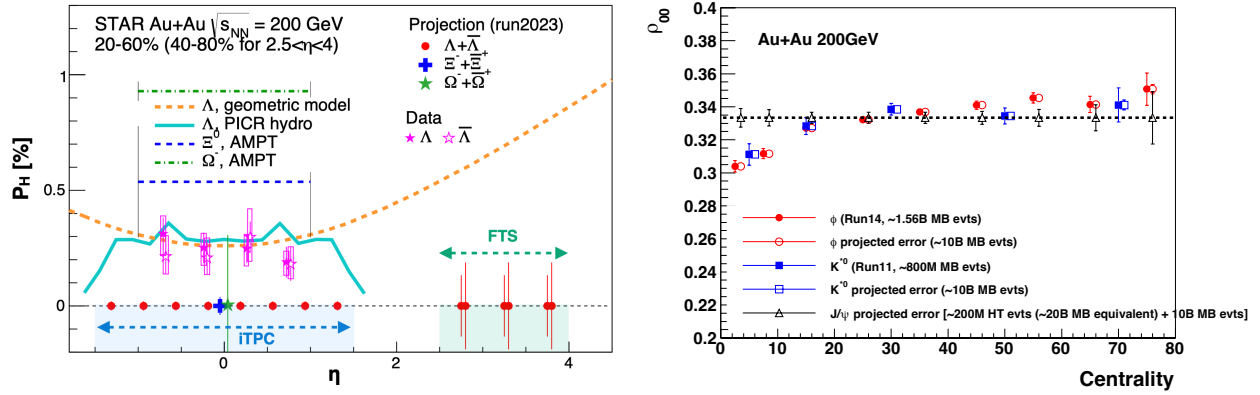


Figure 48: (Left) Projections (along with preliminary data) for differential measurements of Λ ($\bar{\Lambda}$) polarization over the extend range of pseudorapidity with the iTPC and FTS detectors of STAR that will help resolve tension between different theoretical model predictions (shown by curves) of polarization with η . In addition, projections for the measurements of spin-1/2 Ξ and spin-3/2 Ω particles are also shown. (Right) Spin alignment co-efficient ρ_{00} as a function of centrality, with projected errors. The enhanced statistics from run 23 and 25, combined with the excellent dilepton capabilities of STAR, will enable us to measure J/ψ alignment along with increasing the significance of the ϕ and K^{*0} measurements.

1336 The global polarization of hyperons produced in Au+Au collisions has been observed
 1337 by STAR [102]. The origin of such a phenomenon has hitherto been not fully understood.
 1338 Several outstanding questions remain. How exactly is the global vorticity, and its associated
 1339 strong magnetic fields, generated by the two incident heavy ions dynamically transferred
 1340 to the fluid-like medium on the rapid time scales of a collision? Then, how does the local
 1341 thermal vorticity of the fluid gets transferred to the spin angular momentum (magnetic mo-
 1342 ment) of the produced particles during the process of hadronization and decay? In order
 1343 to address these questions one may consider measurement of the polarization of different
 1344 particles that are produced in different spatial parts of the system, or at different times. A
 1345 concrete proposal is to: 1) measure the Λ ($\bar{\Lambda}$) polarization as a function of pseudorapidity
 1346 and 2) measure it for different particles such as Ω and Ξ . Both are limited by the current
 1347 acceptance and statistics available as recently published by STAR [103]. However, as shown
 1348 in Fig. 48 with the addition of the iTPC and FTS, and with high statistics data from runs
 1349 23 and 25 it will be possible to perform such measurements with a reasonable significance.
 1350 iTPC (+TPC) has excellent PID capability to measure all these hyperons. Although the

1351 FTS has no PID capability we can do combinatorial reconstruction of $\Lambda(\bar{\Lambda})$ candidates via
 1352 displaced vertices. A similar analysis was performed and published by STAR using the pre-
 1353 vious FTPC [104]. In order to make a conservative projection we assume similar momentum
 1354 resolution of 10 – 20% for single charged tracks, similar overall tracking efficiency, charge
 1355 state identification capability for the FTS and FTPC. We also assume the FTS, with it's
 1356 novel-tracking framework, will be able to measure a minimum separation of 20 cm between
 1357 the all pairs of one positive and one negative track (a possible decay vertex) from the main
 1358 vertex of the event. This will give rise to about 5% efficiency of $\Lambda(\bar{\Lambda})$ reconstruction with
 1359 about 15 – 20% background contribution from $K_S^0 \rightarrow \pi^+ + \pi^-$ [104]. With this we can
 1360 make projections for a polarization measurement in Au+Au 200 GeV 40 – 80% as shown in
 1361 Fig. 48. The two different error bars correspond to lower and upper limits considering cur-
 1362 rent uncertainties on the efficiency of charged track reconstruction and the final efficiency of
 1363 Λ reconstruction. Currently theoretical models predict contradictory trends for the pseudo-
 1364 rapidity dependence of Λ polarization. If the initial local orbital angular momentum driven
 1365 by collision geometry [105] plays a dominant role it will lead to increases of polarization with
 1366 pseudorapidity. On the other hand if the local thermal vorticity and hydrodynamic evolu-
 1367 tion [106] play a dominant role it will predict decreasing trend or weak dependence with
 1368 pseudorapidity. Such tensions can be easily resolved with the future proposed measurement
 1369 during runs 23 and 25.

1370 **Global Spin Alignment of J/ψ**

1371 Surprisingly large signals of global spin alignment of vector mesons such as $\phi(1020)$ and
 1372 $K^{*0}(892)$ have been measured via the angular distribution of one of their decay products.
 1373 These experimental observations of vector meson spin alignment have yet to be interpreted
 1374 satisfactorily by theory calculations. It has been realized that the mechanism driving the
 1375 global polarization of hyperons can have its imprint on vector meson spin alignments albeit
 1376 the observed strength of signals for the two measurements cannot be reconciled. In fact
 1377 the large quantitative difference between the measurements of $\phi(1020)$ and $K^{*0}(892)$ spin
 1378 alignment as shown in Fig 4 cannot be simultaneously explained by conventional mechanisms
 1379 of spin-orbit coupling, driven by angular momentum, without invoking strong force fields.
 1380 It is argued that the strong force field makes a dominant contribution to the spin-alignment
 1381 coefficient ρ_{00} of ϕ , while for K^{*0} , the contribution is diminished due to the mixing of quark
 1382 flavors (averaging-out of different meson fields) [11, 107]. Therefore, the current experimental
 1383 data from STAR [108] supports the role of strong force field as a key mechanism that leads to
 1384 global spin alignment. An extended test of such a prediction can be performed by measuring
 1385 the spin alignment of J/ψ . This is because similar arguments apply for both and J/ψ , i.e.
 1386 like s and \bar{s} , the strong field component also couples to c and \bar{c} quarks leading to large ρ_{00}
 1387 for J/ψ . ALICE recently reported $\rho_{00} \approx 0.37$ for J/ψ at forward rapidity ($2.5 < y < 4$) with
 1388 a 3.9σ significance, seemingly supporting this argument. STAR can definitely contribute to
 1389 this study by measuring J/ψ global spin alignment at mid rapidity with large data set taken
 1390 during runs 23 and 25.

1391 In Fig 48 we present the projected uncertainties for ρ_{00} of J/ψ estimated for various

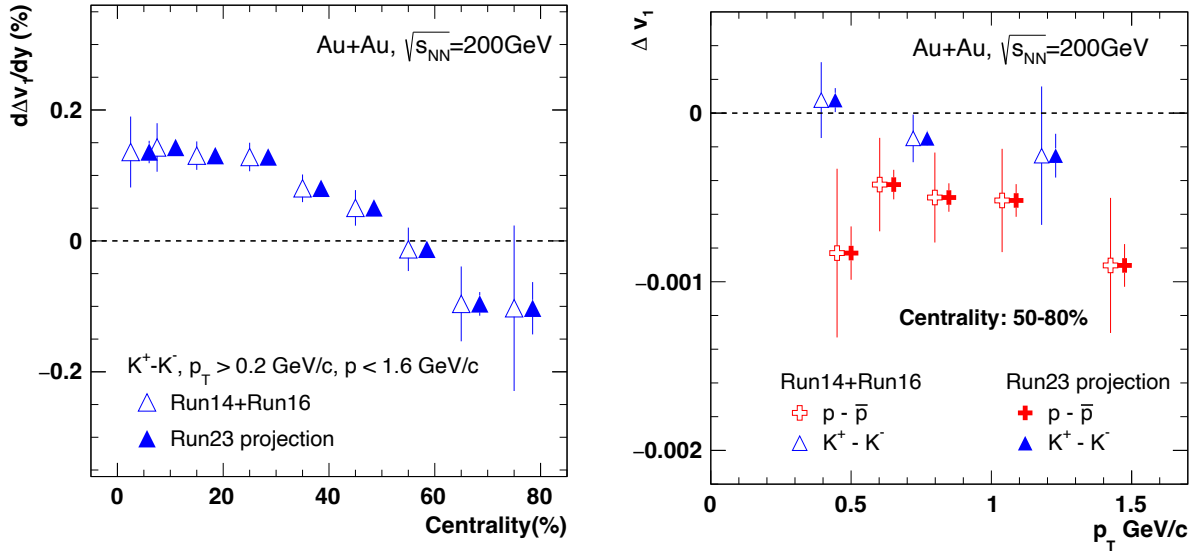


Figure 49: (Left) Plot shows the $\Delta dv_1/dy$ between K^+ and K^- as a function of centrality Au+Au collisions at $\sqrt{s_{NN}} = 200 \text{ GeV}$. Open markers indicate the projection for Run 23. (Right) The $\Delta dv_1/dy$ of proton and kaon as a function of p_T in 50 – 80% Au+Au collisions at 200 GeV.

1392 centralities assuming: 1) 10 B min-bias events for low p_T J/ψ measurements and, 2) 200 M
 1393 events implementing High Tower (BHT3) triggers with the Barrel Electromagnetic Calorimeter
 1394 for the high p_T J/ψ . Both assume 24 weeks running anticipated in Run 23. It is worth to
 1395 mention that apart from J/ψ spin alignment, such a large statistics dataset will also allow
 1396 addition differential study of global spin alignment of ϕ and K^{*0} and help to further elucidate
 1397 the mechanism behind vector meson spin alignment.

1398 Probing Electromagnetic Effect via Charge-dependent Directed Flow

1399 One of the features in high energy heavy-ion collisions is the generation of an ultra-strong
 1400 magnetic field, which is predicted to have the strength of 10^{18} Gauss [109–113]. The interplay
 1401 between magnetic field and QGP may induce many interesting phenomena, such like the
 1402 CME and CMW. Recent studies suggest that the charge dependent directed flow can be
 1403 the probe to search for it in experiment [38, 114]. It predicts a negative $\Delta dv_1/dy$ between
 1404 positively and negatively charged particles due to the influence of electromagnetic field. Some
 1405 experimental efforts have been made for searching this effect, such as the charge dependent
 1406 v_1 measurements presented by LHC-ALICE collaboration [115], and the directed flow of D^0
 1407 and \bar{D}^0 in RHIC-STAR experiment [116]. Results of light flavors in Pb+Pb collisions at
 1408 $\sqrt{s_{NN}} = 5.02 \text{ TeV}$ show large discrepancy to theoretical calculations, which gives an order
 1409 of magnitude larger and positive Δv_1 slope. Similar results have been obtained in Au+Au

1410 collisions at several energies at RHIC, which measured positive Δv_1 slope between proton
 1411 and anti-proton in semi-central collisions owing to the transported quark contributions.

1412 Recent analysis in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV and isobar collisions shows
 1413 striking centrality dependence of this Δv_1 slope. It was found that the $\Delta dv_1/dy$ between
 1414 proton and anti-proton changes from positive to negative as centrality going from central to
 1415 peripheral. The negative value in peripheral collisions, with the significance of 5σ , qualita-
 1416 tively agree with theoretical calculations. However, the $\Delta dv_1/dy$ between K^+ and K^- , π^+
 1417 and π^- are less significant because of the limitation of statistics. If 20 B events in Au+Au
 1418 collisions at 200 GeV could be collected, the $\Delta dv_1/dy$ between K^+ and K^- will have the
 1419 significance $> 5\sigma$, as illustrated in left panel of Fig. 49. Moreover, the EM-field prediction
 1420 shows nontrivial p_T dependence, but this measurements are limited by current statistics. As
 1421 illustrated in right panel of Fig.49, with the data accumulated in Runs 23 and 25, we will
 1422 be able to measure the p_T dependence of $\Delta dv_1/dy$ with higher precision.

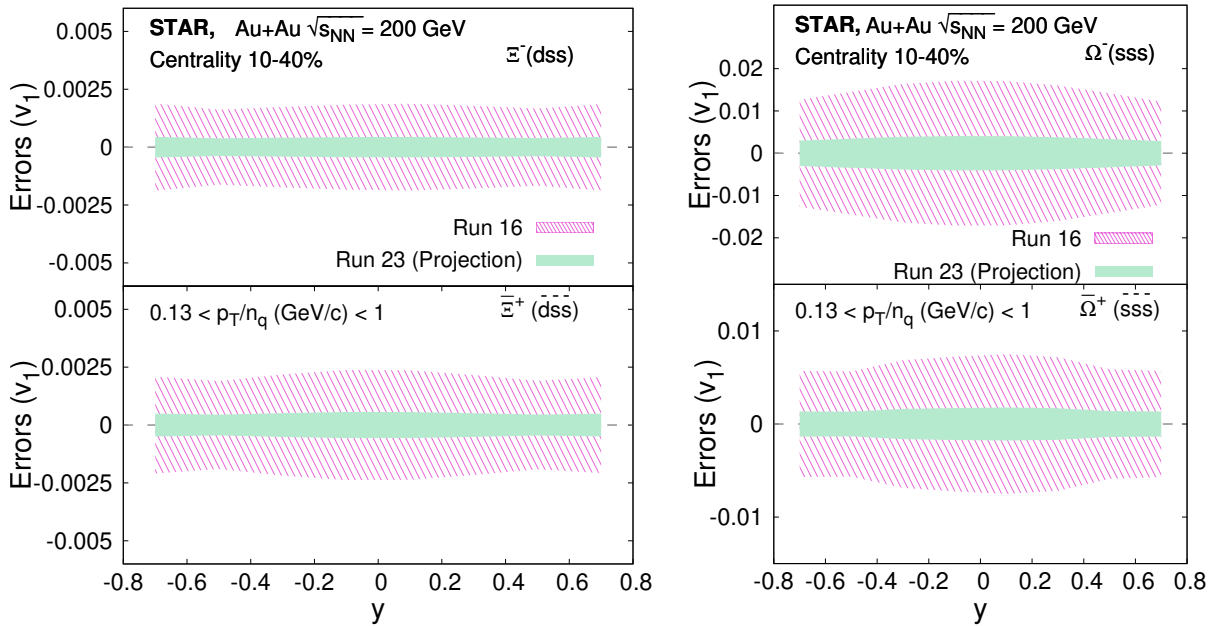


Figure 50: Projection of directed flow (v_1) of Ξ^- , Ξ^+ , Ω^- and $\bar{\Omega}^+$ as a function of rapidity (y) for 10%-40% centrality in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The projections are made by assuming 20B events will be collected in runs 23 and 25.

1423 The existing measurements of v_1 for Ξ and Ω in Au+Au collisions at $\sqrt{s_{NN}} = 200$
 1424 GeV have large uncertainties. There is a hint of a large v_1 for Ω baryons from recent
 1425 measurements, however, as shown in Fig.50, the statistical uncertainties of the current STAR
 1426 measurements are large. There are also measurements for electric charge and strangeness
 1427 dependent splitting in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. These measurements also
 1428 suffer from insufficient statistics. The EM field is expected to lead to increasing splitting
 1429 with increasing electric charge difference. Recent STAR measurements using data from Run
 1430 16 were presented at the Quark Matter 2022 conference. Statistical uncertainties from such

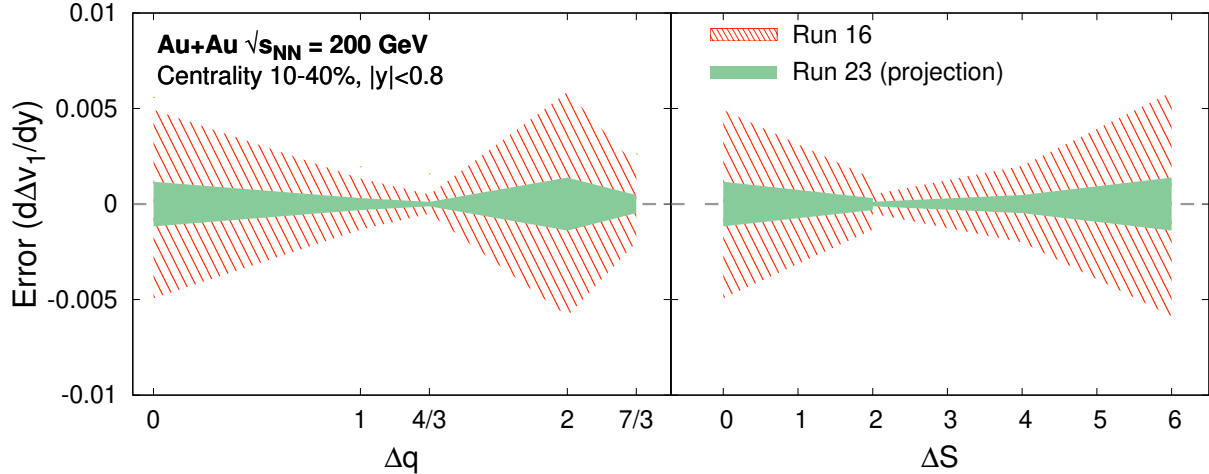


Figure 51: Δv_1 slope ($d\Delta v_1/dy$) at midrapidity as a function of electric charge difference (Δq) and strangeness difference (ΔS) for 10-40% centrality in Au+Au collisions at $\sqrt{s_{NN}}=200$ GeV. The projections are made by assuming 20 B events will be collected in runs 23 and 25.

1431 measurements, as are shown in Fig.51, are limited by statistics. A large dataset (20B) from
 1432 the upcoming Au+Au Runs 23 and 25 at $\sqrt{s_{NN}} = 200$ GeV will definitely help improve
 1433 the precision of these measurements. The projection plots, obtained by assuming that 20B
 1434 events will be collected in the future runs (runs 23 and 25), are shown in Figs. 50 and 51.

1435 **Chiral Properties: ρ - a_1 Mixing**

1436 At $\mu_B \sim 0$ Lattice QCD works and can be directly tested against experimental results.
 1437 In case the measured in-medium spectral function merges into the QGP description this
 1438 would indicate a transition from hadrons into a structure-less quark-antiquark continuum,
 1439 thus providing the manifestation of chiral symmetry restoration. We will continue to search
 1440 for a direct signature of chiral symmetry restoration via chiral ρ - a_1 mixing. The signal is
 1441 predicted to be detectable in the dilepton intermediate mass range. Difficulties are related
 1442 to the fact that correlated charm-anticharm and QGP saturate the invariant mass region
 1443 of 1.1 — 1.3 GeV/ c^2 . Therefore an accurate measurement of the excess dilepton yield,
 1444 i.e. dilepton yield after subtraction of the cocktail of contributions from final-state decays,
 1445 Drell-Yan and those from correlated heavy-flavor decays, up to invariant mass of 2.5 GeV/ c^2
 1446 is required. The challenging analysis on charmed-decayed dielectron is ongoing from the
 1447 datasets taken with the Heavy Flavor Tracker at STAR [117]. Thus deeper understanding
 1448 of origin of thermal radiation in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV from \sim zero mass up
 1449 to 2.5 GeV/ c^2 will become possible with rigorous theoretical efforts and improved dielectron
 1450 measurements. Figure 52 shows the expected statistical and systematic uncertainties of the
 1451 dielectron excess mass spectrum with all the detector upgrades and for the anticipated total
 1452 Runs 23 and 25 statistics of 20 B events.

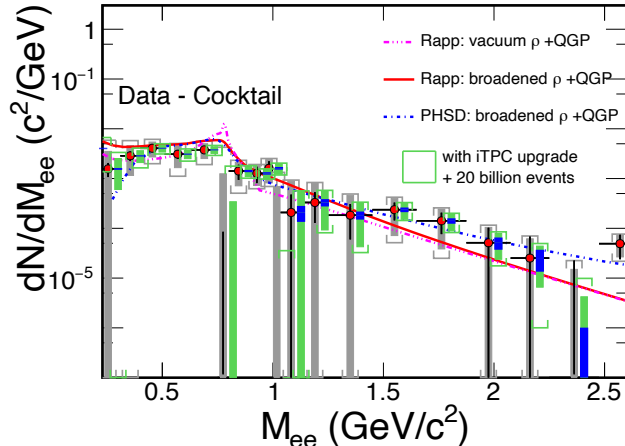


Figure 52: The expected statistical and systematic uncertainties on the dielectron excess mass spectrum with the iTPC upgrade compared to the current TPC case. The data are from our measurements in $\sqrt{s_{NN}} = 200$ GeV Au+Au collisions [118]. Model comparisons are also shown. The boxes represent systematic uncertainties from data and the brackets represent the total systematic uncertainties including those from cocktails. The grey ones are for the current case while the green ones are for the Runs 23 and 25 case. The blue bands represent statistical uncertainties from 20 B min-bias events with the iTPC upgrade.

1453 Electrical Conductivity (Dielectron LMR)

1454 Another application of dielectrons is to use them to measure transport coefficients. The
 1455 electrical conductivity can be directly obtained as the low-energy limit of the EM spectral
 1456 function. We aim to extract such information by studying excess dielectron yields at the low-
 1457 energy regime of the dilepton spectra and the conductivity peak at small invariant masses,
 1458 i.e. at low invariant mass and low p_T^{ee} . Measurement of Drell-Yan in $p+A$ collisions at low
 1459 p_T would provide an important reference to constrain the dilepton cocktail.

1460 **Local Parity Violation and the Chiral Magnetic Effect** A decisive experimental test
 1461 of the Chiral Magnetic Effect (CME) has become one of the major scientific goals of the
 1462 heavy-ion physics program at RHIC. The existence of CME would be a leap towards an
 1463 understanding of the QCD vacuum, establishing a picture of the formation of a deconfined
 1464 medium in which chiral symmetry is restored, and it would also provide unique evidence that
 1465 the strongest known electromagnetic fields are created in relativistic heavy-ion collisions [119,
 1466 120]. The impact of such a discovery would go beyond the community of heavy-ion collisions
 1467 and will possibly be a milestone in physics. The remaining few years of RHIC running and
 1468 analyses of previously-collected data will likely provide the only chance for CME searches in
 1469 heavy-ion collisions in the foreseeable future.

1470 The isobar collisions provided an unique opportunity to search for the CME because of
 1471 the $\sim 15\%$ difference in B^2 and hence the CME correlation signals between Ru+Ru and
 1472 Zr+Zr collisions. No CME signal has been observed in isobar data even with an improved
 1473 understanding of background baseline. The signal/background ratio is expected on general

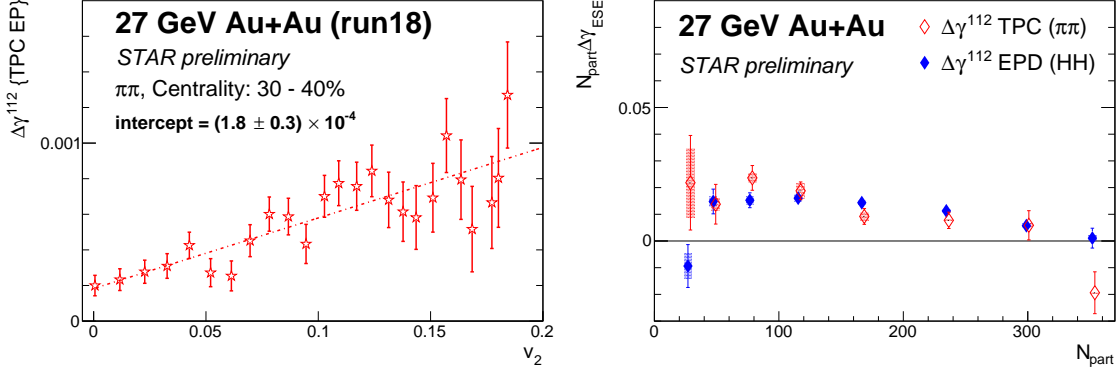


Figure 53: (Left) $\Delta\gamma_{112}$ for π - π vs v_2 measured with the TPC event plane in 30–40% Au+Au collisions at 27 GeV. (Right) $\Delta\gamma_{112,ESE}$ scaled by N_{part} as a function of N_{part} for π - π using the TPC event plane, and for hadron-hadron using the EPD event plane in Au+Au collisions at 27 GeV.

1474 ground to be smaller in isobar collisions than in Au+Au collisions by an approximately factor
 1475 of 3 [6]. This would be in line with the Au+Au data which indicate a positive CME signal
 1476 of $\sim 8\%$ with $\sim 2\sigma$ significance using the spectator/participant plane method [3].

1477 The current Au+Au data statistics are total 2.4 B events from Runs 2011, 2014 and
 1478 2016 [3]. In order to achieve 5σ significance with the same analysis one needs to have 15
 1479 B events. Therefore, with the proposed 20 B events that can be collected by STAR during
 1480 runs 23 and 25, one can achieve more than 5σ significance provided the possible CME signal
 1481 remains at 8%. A stringent upper limit will be possible on the CME.

1482 This estimate does not account for two important facts that can lead to higher signif-
 1483 icance. The first is that the iTPC upgrade enhances the charge particle multiplicity by
 1484 50% and therefore triplet ($\sim dN/d\eta^3$) (pair $\sim dN/d\eta^2$) statistics by a factor of 3.4 (2.3).
 1485 The second one is the addition of the EPD detector which will significantly reduce nonflow
 1486 contaminations in the measurements with respect to the participant plane. Our estimate
 1487 assumes that the systematic uncertainty can be controlled to be smaller than the statistical
 1488 uncertainty, i.e. below 1%.

1489 Running STAR in runs 23 and 25, concurrently with sPHENIX, would be essential to per-
 1490 form precision measurements to further investigate and characterize the CME phenomenon.
 1491 Such a program will have a strong discovery potential.

1492 The dominant background in the CME-sensitive $\Delta\gamma_{112}$ correlator is caused by the cou-
 1493 pling of elliptic flow and other mechanisms such as resonance decays and local charge con-
 1494 servation. Accordingly, the event-shape engineering (ESE) method aims to project $\Delta\gamma_{112}$ to
 1495 a class of events with minimal flow to suppress the v_2 -related background. We adopt an ESE
 1496 technique [121] that uses the flow vector ($q_{2,x} = \frac{1}{\sqrt{N}} \sum_i^N \cos(2\phi_i)$, $q_{2,y} = \frac{1}{\sqrt{N}} \sum_i^N \sin(2\phi_i)$) to
 1497 select spherical sub-events with almost zero v_2 . Observables like v_2 and γ_{112} are measured
 1498 as a function of q_2^2 from the particles of interest (POIs), and then $\Delta\gamma_{112}$ is plotted against
 1499 v_2 in the same q_2^2 interval to yield a reliable projection to the zero-flow mode.

1500 Figure 53 (left) demonstrates the application of the ESE approach to the STAR data

1501 of 30–40% Au+Au collisions at 27 GeV (run 2018), and the decrease of $\Delta\gamma_{112}$ for π - π with
 1502 decreased v_2 illustrates how the v_2 -related background is suppressed. Figure 53 (right)
 1503 shows the centrality dependence of $N_{\text{part}} \Delta\gamma_{112\text{ESE}}$ for π - π using the TPC event plane, and
 1504 for hadron-hadron using the EPD event plane in Au+Au collisions at 27 GeV.

1505 The ESE method will be applied to the 200 GeV Au+Au data from Run-23 and Run-25.
 1506 With the large data set of anticipated 20 B events, we are able to perform more differential
 1507 measurements and involve identified particles such as kaons and protons.

1508 Event-by-event correlations between CME charge separation and other parity-odd fea-
 1509 tures of the event will be studied. One such analysis is motivated by the idea that the local
 1510 parity violation (characterized in each event by a net topological charge Q) that is expected
 1511 to work with the spectator-produced magnetic field to give the CME should also cause a net
 1512 helicity of $\Lambda(\bar{\Lambda})$ in the event. Importantly, even though both of these parity-odd signatures
 1513 switch handedness event-by-event, in any given event they should have the same handedness
 1514 as one another and so can be compared with one another in a correlation analysis. To do
 1515 this, the charge separation with respect to the first-order reaction plane must be determined
 1516 in each event.

1517 We are looking for evidence of an event-by-event correlation between these two parity-
 1518 odd effects. A measured event-by-event correlation between Δa_1 and ΔN would be strong
 1519 evidence for the CME and underlying local parity violation, and would extend the mea-
 1520 surement into other parity-odd effects. Note also that the flow-related backgrounds that
 1521 plague charge-separation measurements are not expected to affect ΔN and so should not be
 1522 a background for this correlation measurement.

1523 We use a similar toy model to that used in [122] to estimate the number of events required
 1524 to see non-zero correlations between Δa_1 and ΔN with different CME signal fraction in the
 1525 $\Delta\gamma$ measurement (see Fig. 54). The chief unknown in this estimate is the extent to which
 1526 strange quarks may be counted as light quarks and so will have a net handedness imparted by
 1527 the parity-odd domain. Recent theoretical work [123] makes a direct connection between the
 1528 net lambda helicity and the axial chemical potential developed from local parity violation.
 1529 Such work holds promise of leading to an improved estimate of the expected signal size, but
 1530 it is likely that this will remain a speculative measurement in which a non-observation will
 1531 be difficult to interpret quantitatively but a positive observation would be a very significant
 1532 result.

1533 Figure 54 suggests that this will be a topic requiring the large datasets of runs 23 and
 1534 25. To explore this correlation, we have analyzed the Run-18 Au+Au collision data at
 1535 $\sqrt{s_{NN}} = 27$ GeV. The $\Lambda(\bar{\Lambda})$ baryons are reconstructed by their decay daughter tracks and
 1536 identified by the `KFPARTICLE` package. Each Λ handedness is estimated by decay kinematics.
 1537 After a purity correction, N_L and N_R are calculated for both Λ and $\bar{\Lambda}$ in each event, and
 1538 then Δn (normalized ΔN , $\Delta n = \frac{N_L - N_R}{(N_L + N_R)}$) is calculated. The observable Δa_1 can be calcu-
 1539 lated from primordial particles' azimuthal angles w.r.t. the first-order EP measured by the
 1540 EPD. The covariance between Δn and Δa_1 is then calculated for the event sample. In this
 1541 exploratory measurement, the covariance is consistent with zero, and so no significant cor-
 1542 relations have been observed (see Fig. 55). However, this event-by-event correlation method

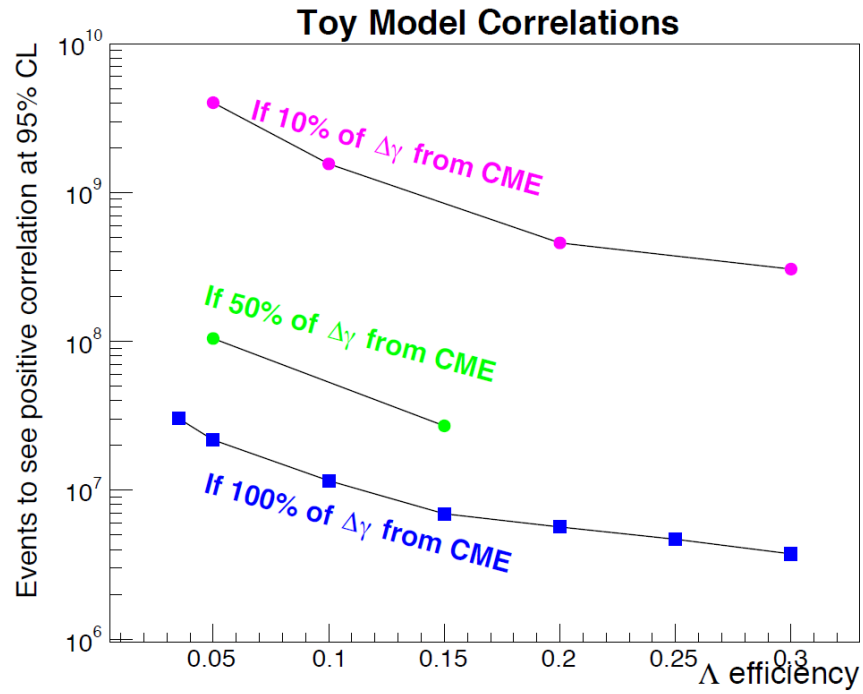


Figure 54: Estimation of the number of events required to see positive correlation between net Λ helicity with out-of-plane charge separation sensitive to local parity violation at 95% confidence level, plotted against the efficiency of $\Lambda(\bar{\Lambda})$ reconstruction (see Ref. [122] for details).

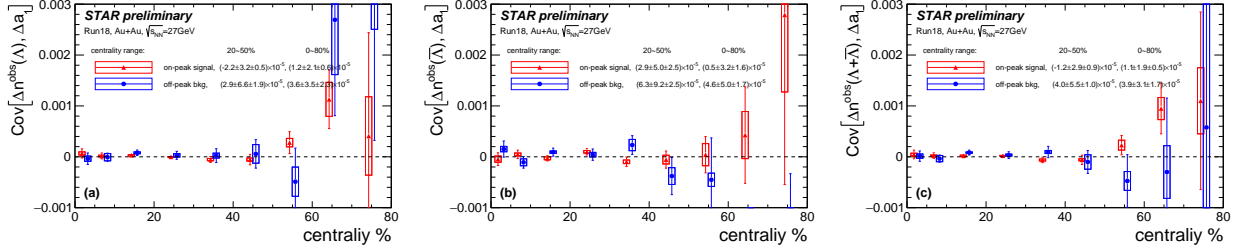


Figure 55: The covariance between Δa_1 and measured Δn for Λ (Left), $\bar{\Lambda}$ (Middle), and the sum of them (Right) as functions of centrality. The red markers come from the Λ ($\bar{\Lambda}$) mass peak region with purity correction and blue markers come from the side bands for pure background.

1543 holds great potential with future high statistics data from Runs 23 and 25 by a qualitatively
 1544 new technique different from all existing analyses.

1545 What are the underlying mechanisms of jet quenching?

1546 The dependence of jet energy loss on the jet p_T and/or resolution or angular scale tagged
 1547 by jet substructure observables are key tools in discriminating various jet quenching mecha-
 1548 nisms [124–127]. In addition, the measurement of jet acoplanarity is a sensitive probe of p_T
 1549 broadening and medium-induced radiative effects [128], particularly for jets at low p_T which
 1550 are accessible at STAR by selecting a given momentum transfer via a photon trigger. Such a
 1551 measurement is also affected by background arising from vacuum Sudakov radiation at RHIC
 1552 energies [129, 130], potentially enabling a precise extraction of in-medium jet scattering.

1553 STAR’s unique geometry allows collection of events over a wide range of vertex positions
 1554 along the beam direction (v_z) for jet analyses, thereby efficiently sampling the provided
 1555 RHIC luminosity. Optimization of the v_z range used in the various analyses involves a
 1556 balance between statistical precision and complexity of corrections, with the latter predom-
 1557 inantly contributing to the systematic uncertainties of the measurement. Recent STAR
 1558 jet measurements in Au+Au collisions have employed two classes of v_z cuts: the inclusive
 1559 charged-particle jet analysis [131] utilizes $|v_z| < 30$ cm, whereas the $\gamma_{\text{dir}} + \text{jet}$ analysis uti-
 1560 lizes $|v_z| < 70$ cm. With the $\gamma_{\text{dir}} + \text{jet}$ measurement successfully utilizing the broad v_z range
 1561 with controlled systematic precision, we are exploring similar event selections maximizing
 1562 the available statistics for future jet measurements, including the inclusive/differential jet
 1563 analyses. In the following discussions, we assume an integrated luminosity of 40 nb^{-1} , which
 1564 is roughly a factor 4 increase in trigger statistics relative to the current analyses based on
 1565 Run 14 data.

1566 To quantify the effect of the marked increase in integrated luminosity, we utilize two
 1567 mature jet measurements and discuss their expected improvements. These analyses are the
 1568 semi-inclusive distribution of charged-particle jets recoiling from a high- E_T direct-photon
 1569 trigger ($\gamma_{\text{dir}} + \text{jet}$); and the differential measurement of energy loss for jet populations se-
 1570 lected by varying a substructure metric. Since these analyses are mature, their analysis
 1571 methodologies and correction schemes are optimized, so that their projections based on

1572 increased statistics are meaningful.

1573 **Semi-inclusive $\gamma_{\text{dir}} + \text{jet}$ measurements (I_{AA})**

1574 Figure 56 shows I_{AA} for fully-corrected semi-inclusive distributions of charged-particle jets
 1575 (anti- k_T , $R = 0.5$) recoiling from a direct-photon trigger with $15 < E_T < 20$ GeV in central
 1576 Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV, for the current analysis based on 10 nb^{-1} [132] within
 1577 $|vz| < 70$ cm. The projected uncertainties, including the previous years and runs 23 and 25,
 1578 are shown in green bands. Significant reduction in the uncertainty band is seen resulting
 1579 from the increase in integrated luminosity, together with a significant increase in kinematic
 1580 reach as indicated by the extended green band along the x-axis.

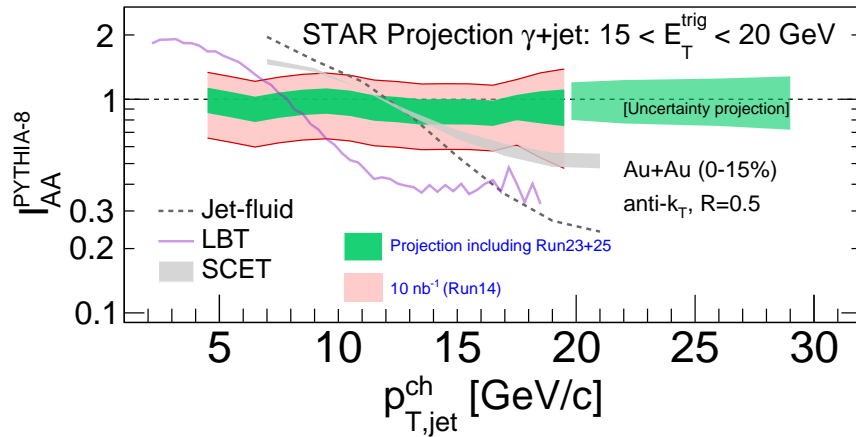


Figure 56: Projections of the I_{AA} for semi-inclusive anti- k_T , $R = 0.5$ jets recoiling from a direct-photon trigger with $15 < E_T < 20$ GeV for central (0-15%) Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV. The colored bands show the cumulative uncertainties for the current analysis and projections for future analysis with the higher statistics datasets.

1581 The luminosity projection of 40 nb^{-1} is expected to reduce the systematic uncertainty
 1582 band by a factor of 2 from the current measurement since systematic uncertainty of this mea-
 1583 surement, dominated by the unfolding procedure, is correlated with the statistical precision.
 1584 Due to this correlation, the improvement shown in Fig. 56 should be regarded as a con-
 1585 servative estimate of the improvement in precision of this measurement with the projected
 1586 integrated luminosity increase.

1587 **Jet acoplanarity**

1588 The p_T broadening due to medium effects not only modifies the shape but also introduces a
 1589 decorrelation between the di-jet angular distributions. The vacuum QCD process (Sudakov
 1590 radiation) makes such measurements challenging in heavy-ion collisions, but at RHIC the
 1591 Sudakov effect is smaller than at the LHC as it depends on the virtuality Q^2 [129, 130]. A
 1592 detailed study is needed to understand both effects (medium-induced and vacuum radiation)
 1593 in a wide range of jet $p_{T,\text{jet}}$ both at RHIC and the LHC energies. Such measurements are

1594 crucial to probe \hat{q} and/or quest for the predicted large-angle jet scattering off of quasi-
 1595 particles in the QGP [133].

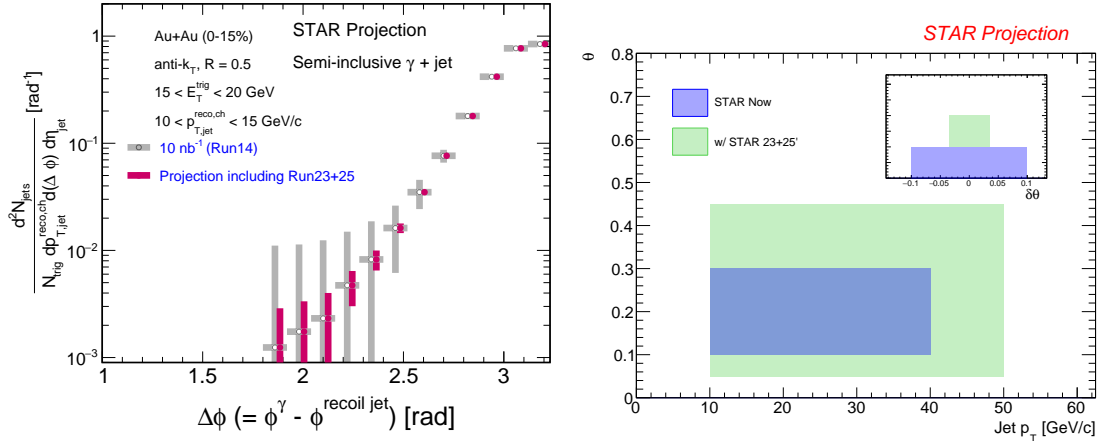


Figure 57: Left: Projections of the acoplanarity for semi-inclusive anti- k_T , $R = 0.5$ jets recoiling from a direct-photon trigger with $15 < E_T < 20$ GeV and $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/c for central (0-15%) Au+Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV. The colored bands show the cumulative uncertainties for the current analysis and projections for future analysis with the higher statistics datasets. Right: The subjet opening angle as a function of jet $p_{T,\text{jet}}$ in 0-20% central Au+Au collisions. The inset is the corresponding resolution of θ . Blue and green represent current (10 nb^{-1}) and future (including runs 23 and 25) analyses, respectively.

1596 In this direction, the STAR experiment reports the first signature of medium-induced
 1597 acoplanarity in the central Au+Au collisions as discussed in section 1.1.5 Fig. 20 (right
 1598 figure). This measurement is performed for both γ_{dir} and π^0 triggers with $11 < E_T < 15$
 1599 GeV and charged-particle jets (anti- k_T , $R = 0.2$ and 0.5) with $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/c.
 1600 To have a better understanding of the nature of this acoplanarity, we plan to extend both
 1601 E_T^{trig} and recoil jet $p_{T,\text{jet}}$ kinematic ranges which demands high statistics datasets. On the
 1602 other hand, the STAR experiment also reports the same measurements in $p+p$ collisions to
 1603 study the shape of this acoplanarity in vacuum. In this direction, both $\gamma_{\text{dir}}+\text{jet}$ and $\pi^0+\text{jet}$
 1604 measurements would be crucial to study trigger dependence of $\Delta\phi$ decorrelation between the
 1605 trigger and recoil jets in $p+p$ collisions and sets a baseline for Au+Au collisions. However,
 1606 due to limited statistics we only report $\pi^0+\text{jet}$ measurement in $p+p$ collisions as shown in
 1607 Fig. 20 left. Furthermore, this measurement could exploit forward triggering using forward
 1608 calorimeter to explore a relatively small x region, compared to mid-rapidity measurement,
 1609 in $p+p$ collisions. This is important to study various pQCD effects like NLO corrections,
 1610 ISR/FSR, and MPI effects. Upcoming Run-24 $p+p$ collision data-taking is very important
 1611 in this direction.

1612 The left plot of Fig. 57 shows the semi-inclusive distribution of the azimuthal separation
 1613 between a direct-photon trigger with $15 < E_T < 20$ GeV and a charged-particle jet (anti- k_T ,
 1614 $R = 0.5$) with $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/c, in central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV with

1615 only statistical uncertainties. The azimuthal smearing of this observable due to uncorrelated
 1616 background is small, and such acoplanarity measurements are therefore strongly statistics-
 1617 limited [134,135]. The grey vertical bars represent the statistical uncertainty with the current
 1618 preliminary measurement based on 10 nb^{-1} , whereas the red vertical bars correspond to
 1619 including Run-23+25.

1620 Differential Measurement of Energy Loss Tagged with a Substructure Metric

1621 Systematic exploration of parton energy loss controlled for variations in the jet shower forms
 1622 an integral part of the jet program at STAR. Since parton showers are inherently probabilis-
 1623 tic, a jet population contains patterns of radiation varying in both angle and momentum
 1624 scales which can be extracted via jet substructure measurements defined based on jet con-
 1625 stituents' angle and/or momentum via algorithms or correlations. By selecting jets based
 1626 on their substructure, STAR can differentially measure jet-medium interactions for various
 1627 types of energy loss e.g. color coherence, dead cone, etc. In other words, the STAR jet
 1628 program for Run-23+25 will focus on jet substructure as a jet-tagger.

1629 Theory calculations show significant differences between energy loss signatures for jets
 1630 perceived by the medium as a single or multiple color charges [126]. The integrated luminosity
 1631 from the Runs 23 and 25 datasets not only provide a substantial increase in statistics in the
 1632 current measurements of jet substructure, they also increase the available phase space for
 1633 rare processes such as wide angle emissions from high- p_T jets. This enables STAR to extend
 1634 our current measurements of differential energy loss from a resolution of $\delta\theta = 0.1$ to finer
 1635 resolution $\delta\theta \approx 0.025$ in the jet opening angle as shown in Fig. 57 (right), and also extend to
 1636 jets of higher momenta. By extending to high energy splittings within jets at varied opening
 1637 angles, we can probe earlier formation times whereby vacuum-like emissions and medium
 1638 induced radiations are expected to occur.

1639 Given the unique nature of jet-medium interactions at RHIC, with the jet and sub-jet
 1640 scales sufficiently closer to the medium scale than the LHC, the aforementioned measure-
 1641 ments bolster the importance of the STAR jet program with the goal of extracting the
 1642 microscopic properties of the QGP as outlined in the 2015 LRP.

1643 What is the nature of the phase transition near $\mu_B = 0$?

1644 LQCD calculations [136,137] predict a sign change of the susceptibility ratio χ_6^B/χ_2^B with
 1645 temperature (T at $\mu_B = 0$) taking place in the range of 145-165 MeV. The observation of
 1646 this ratio going from positive to negative values is considered to be a signature of a crossover
 1647 transition. Interestingly, as shown in Section ??, values of net-proton C_6/C_2 are found to be
 1648 negative systematically from peripheral to central Au+Au 200 GeV collisions within large
 1649 statistical uncertainties. The observation of negative C_6/C_2 is intriguing and so far only
 1650 hinted at in the 200 GeV data, the current result has less than 2.3σ significance for 30-
 1651 40% centrality in terms of statistical uncertainties. The current systematic uncertainty is of
 1652 similar order as the statistical uncertainty and if based off of combining datasets from Run-10
 1653 and Run-11. As shown in the projection plot of Fig. 58 it is possible to establish definitive

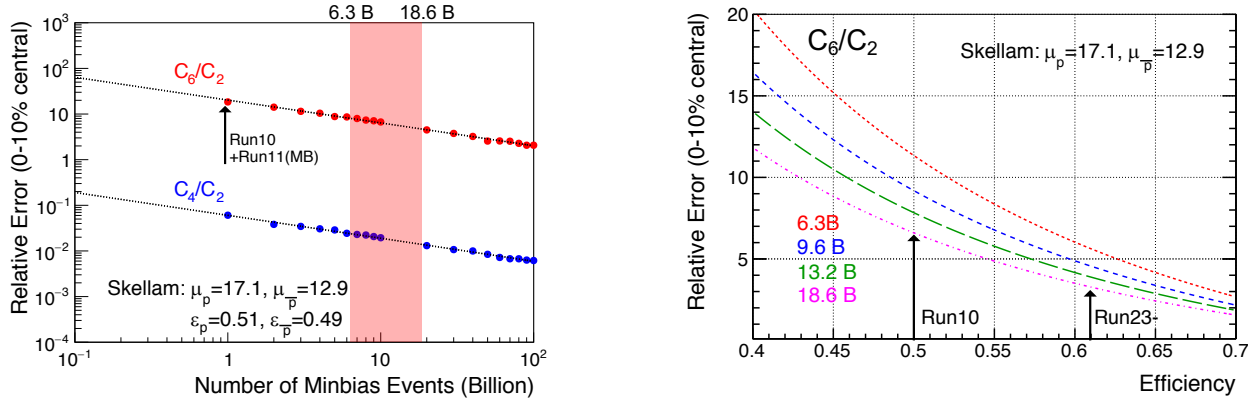


Figure 58: Projection for measurement of ratio of sixth order over second order cumulants of net-proton distribution.

1654 observation of negative C_6/C_2 at 200 GeV the 20 B minimum-bias events to be collected
 1655 during the run 23 and 25 with 15% increase in the reconstruction efficiency and enhanced
 1656 acceptance of the iTPC detector upgrade. A similar measurement can be performed at the
 1657 LHC for vanishing baryon chemical potential, while only STAR measurements can explore
 1658 the finite μ_B region. Our measurement at $\sqrt{s_{NN}}=200$ GeV has the potential to establish the
 1659 first experimental observation of QCD chiral crossover transition at $\mu_B \approx 20$ MeV.

1660 What can we learn about the strong interaction?

1661 The strong interaction between baryons leads to a residual force; the most common example
 1662 is NY . The same force is responsible for binding $n-p$ into d . So far, understanding the
 1663 strong interaction has been limited to the effective theories related to nucleons and the
 1664 scattering experiments, which are very challenging due to the short lifetime of those baryons
 1665 (a few cm decay length). One of the current challenges is to evaluate the strong interaction
 1666 between hyperons, as experimentally still very little is known about NY and YY interactions.
 1667 Hypernuclei (a hyperon bound inside an atomic nucleus) are proof of a positive, attractive
 1668 interaction of NY . Measurements of NN and NY interactions have crucial implications for
 1669 the possible formations of bound states. Studies of the strong interaction potential via two-
 1670 particle correlations in momentum space measured in relativistic heavy-ion and elementary
 1671 collisions have proven to be useful to gain access to the interactions between exotic and rare
 1672 particles. Possible combinations can be: $p\Lambda$, $p\Sigma$, $p\Omega$, $p\Xi$, $\Lambda\Lambda$, $\Xi\Xi$. In contrast to $p\Lambda$, the
 1673 nature of $p\Sigma$, $p\Omega$, $\Lambda\Lambda$ still need experimental verification. Even if scattering experiments are
 1674 available, they are not very conclusive.

1675 Figure 59 shows the preliminary $p\Xi$ correlations function. All available statistics, 3 B
 1676 events accumulated over all previous runs, were used for the $p\Xi$ and $p\Omega$ cases. Combining
 1677 such datasets leads to the run-to-run variations resulting in larger total systematic uncer-
 1678 tainties in the detector responses. A long run with similar detector settings during the runs
 1679 23 and 25 will avoid such issues. Statistical uncertainties of the current measurements re-

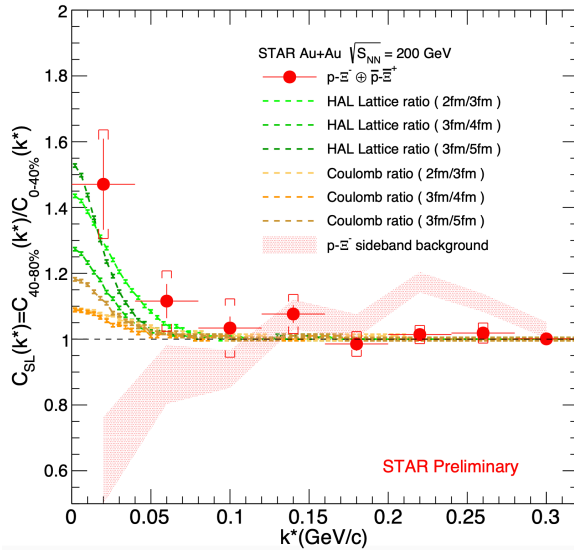


Figure 59: Solid circles represent the ratio (R) of the small system (40-80% collisions) to the large system (0-40% collisions) for proton- Ξ and \bar{p} - Ξ correlations. The bars correspond to the statistical uncertainties. The shaded area represents R for background candidates from the side-band of the Ξ invariant mass. Coulomb-induced R are shown in yellow and orange colors. HAL Lattice predictions of R are shown in green.

1680 main high, and the number of points that build the correlation function is minimal. This
 1681 means that the current results are not conclusive enough to study in detail the parameters
 1682 of the strong interaction. The collection of 10 B events from run 23 will make possible the
 1683 construction of correlation functions of the $p\Xi$ case with double the number of points and
 1684 smaller statistical uncertainties than the current measurement.

1685 The $p\Omega$ system is more statistics hungry, and will require 20 B events, from combining
 1686 runs 23 and 25. Previous STAR measurements of $p\Omega$ correlations show that the parameters
 1687 of the strong interaction can be studied. However, with higher data collections, more precise
 1688 and detailed studies would be possible.

1689 The description of the $\Lambda\Lambda$ interaction is still an open issue. Such a description is funda-
 1690 mental since it plays a decisive role in understanding the nature of hyperons that appear in
 1691 neutron stars. If many hyperons appear close to each other and their fraction becomes signif-
 1692 icant, the YY interactions are expected to play an essential role in describing the equation of
 1693 state of the dense system. An alternative way to study hypernuclei is two-particle momentum
 1694 correlations of $\Lambda\Lambda$ pairs produced in hadron-hadron collisions thanks to femtoscopy. Figure
 1695 60 shows primary $\Lambda\Lambda$ (left) and $\Xi\Xi$ (right) correlation functions. For current $\Lambda\Lambda$ and $\Xi\Xi$
 1696 systems also data from all previous runs were combined. Due to differences between individ-
 1697 ual runs, a significant source of systematic uncertainties exist now, and it will disappear with
 1698 all events collected during run 23 for the $\Lambda\Lambda$ case. More critical seems to be the increased
 1699 statistics for the $\Xi\Xi$ case, and having 20 B events from runs 23 25 enables the reduction of

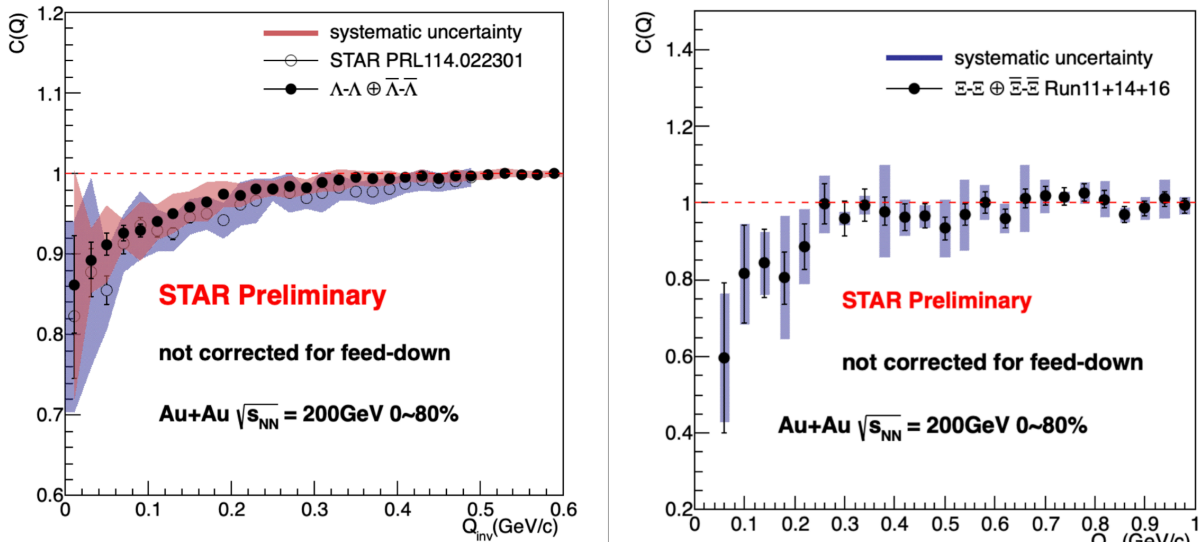


Figure 60: Left: combined $\Lambda\Lambda$ and $\bar{\Lambda}\bar{\Lambda}$ preliminary correlation functions with systematic uncertainties compared with already published previous STAR results. Right: combined $\Xi\Xi$ and $\bar{\Xi}\bar{\Xi}$ correlation functions with systematic uncertainties.

1700 statistical uncertainties significantly and makes it possible to determine parameters of the
 1701 strong interaction with higher precision. Having combined data from the runs 23 and 25 will
 1702 also allow the hypotheses about possible bound states to be verified.

1703 2.2 Ultra-Peripheral Collisions

1704 One of the most important scientific goals in high-energy nuclear physics is to understand the
 1705 nuclear structure under extreme conditions. Thanks to ultra-relativistic heavy-ion collider
 1706 facilities, e.g., the Relativistic Heavy-Ion Collider, one direction is to create a system that has
 1707 an extremely high temperature of partons, and study its deconfined properties of strongly
 1708 interacting quarks and gluons. However, the other direction is to reveal the property of
 1709 nucleons and nuclei before such violent collision happens, where the initial-state dynamics
 1710 inside these particles may provide ultimate understanding of the Quantum Chromodynamics
 1711 (QCD) in generating the visible matter. These two aspects are usually known as the heavy-
 1712 ion hot Quark-Gluon-Plasma (QGP) physics and cold QCD physics, respectively. Both of
 1713 them are indispensable building blocks of our fundamental understanding of nuclear physics.
 1714 In this section, we will focus on the initial-state physics program via the ultra-peripheral
 1715 collision in nucleus-nucleus (AA) interactions.

1716 In relativistic heavy-ion collisions, a large fraction of the total cross section or interaction
 1717 between the two colliding nucleus is provided by photon-induced reactions. Most of these
 1718 events are removed by the requirement of inelastic collisions, because the hot quark-gluon-
 1719 plasma (QGP) can be more likely, if not only, to be produced in such high parton density

1720 system. However, these events are difficult to understand if one wants to separate effects
1721 related to the initial state, e.g., nuclear parton distribution functions (nPDFs), from final-
1722 state interactions, such as fragmentation, medium-induced collective effects, etc. One way to
1723 overcome this difficulty is to “turn off” the QGP and use a simple and clean probe to examine
1724 the nuclear target - photon-nucleus collisions, which is also known as the “ultra-peripheral
1725 collisions” (UPC).

1726 Typically, the UPC takes places when the impact parameter between the two colliding
1727 nucleus is greater than the sum of their radii. The interaction is initiated by one or multiple
1728 photons emitted from the moving charged ions, where the photon interacts with the other
1729 nucleus. Due to the large mass of the heavy nucleus, the emitted photons have very small
1730 virtualities or very small p_T . This process is regarded as *photoproduction*. For example,
1731 diffractive Vector Meson (VM) photoproduction has been extensively studied at the RHIC
1732 and at the LHC, where the gluon density distribution of the nucleon and nucleus target
1733 can be directly probed. In recent analyses carried out by the LHC collaborations [138–
1734 145], photoproduction of the J/ψ meson has been measured in UPCs of heavy ions. The
1735 resulting cross sections were found to be significantly suppressed with respect to that of a free
1736 proton [138, 139, 143, 144]. Leading Twist Approximation (LTA) calculations strongly suggest
1737 that the suppression is caused by the gluon shadowing effect [146–148], while other models,
1738 e.g., the Color Dipole Model with gluon saturation and nucleon shape fluctuations [149], can
1739 also describe the UPC data qualitatively. The mechanism of gluon density modification in
1740 the nuclear environment remains unknown.

1741 However, there are other processes of photoproduction that are sensitive to the nPDFs.
1742 For example, inclusive and diffractive back-to-back jets (dijets) in nuclei are sensitive to
1743 both quark and gluon distribution, and it is theoretically easier to be used in the global
1744 PDF analysis. Recent studies from Refs. [150–152] have shown the uncertainty of nPDFs
1745 can be reduced by a factor of 2 by having these experimental measurements. In addition,
1746 the incoming low-virtuality photons can have properties of a point-like particle (direct pro-
1747 cess) or a hadron with partonic substructure (resolved process). The dijets photoproduction
1748 process can be extremely useful in constraining the photon structure, which still remains
1749 poorly known to-date. Finally, the diffractive dijets contribution is a sensitive experimental
1750 observable to understand the QCD factorisation breaking and the diffractive nPDFs.

1751 Last but not least, inclusive particle photoproduction at high energy provides important
1752 insights to the *soft physics* of photon-nucleus interactions, where cold nuclear matter and
1753 Intra-Nuclear Cascade can be studied via fragmentation in both current and target frag-
1754 mentation regions. One recent study led by Chang et al [153] has shown the difficulty of
1755 describing the charged particle production in nuclei of existing E665 experimental data. Al-
1756 though the experimental data is with higher photon virtualities, not many data exists at high
1757 energy at all for photoproduction. Furthermore, inclusive charged particle photoproduction
1758 can be a baseline for comparison to the diffractive VM production, where different theoretical
1759 models have drastically different prediction, e.g., gluon saturation model [149] verse nuclear
1760 shadowing model [146–148]. Together with the VM production and with different level arm
1761 of photon virtualities, this measurement is one of the most important scientific goals at the

1762 upcoming US-based Electron-Ion Collider (EIC).

1763 Hereby, we propose to utilize the unique capability of the RHIC experimental program
 1764 in the upcoming 2023-2025 runs with the STAR detector and its recent forward upgrades,
 1765 to study photoproduction processes in details. The major advantage is that the top RHIC
 1766 energy can access a kinematic regime that is hardly, if not at all, accessible by the LHC
 1767 experiments, and provide a seamless transition to the physics at the EIC.

1768 Photoproduction of Vector-Meson

1769

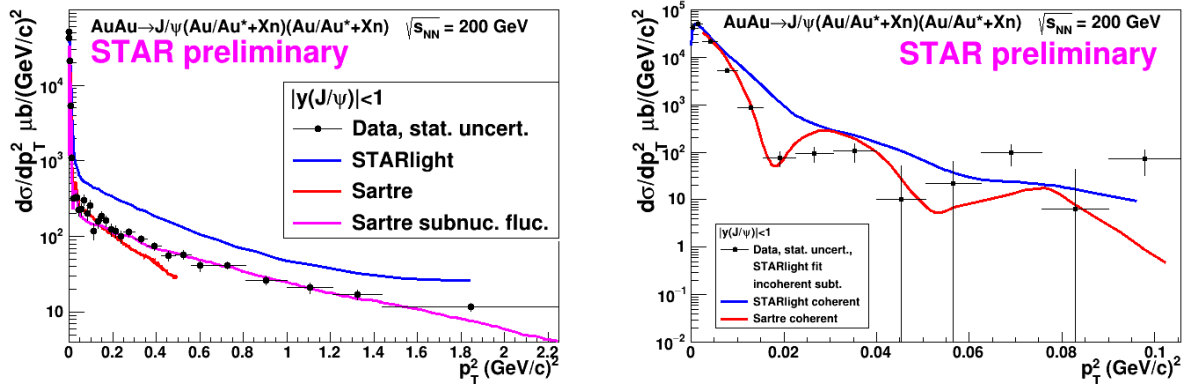


Figure 61: Left: differential cross section $d\sigma/dp_T^2$ of J/ψ photoproduction as a function of p_T^2 in Au+Au UPC at 200 GeV. Right: the same cross section but with incoherent contribution subtracted.

1770 One of the most important and direct measurements of the gluon density in the initial-
 1771 state of nuclei is the photoproduction of Vector-Meson, e.g., ρ^0 , ϕ , and J/ψ . The process
 1772 can be generally considered in a color dipole picture, where the quasi-real photon emitted
 1773 from the heavy nucleus fluctuates into a quark and anti-quark pair (leading order). The
 1774 quark and anti-quark pair scatters off the nucleus with a Pomeron exchange and becomes a
 1775 Vector-Meson; the cross section of this process is directly sensitive to the gluon density and
 1776 its spatial distribution.

1777 In previous STAR publications, there has been studies on ρ^0 meson, e.g., the most re-
 1778 cent analysis in Ref. [140] for coherent photoproduction. Although the measurement has
 1779 provided important insights to the structure of the gold nucleus, e.g., the impact parameter
 1780 distribution from a Fourier transform of the momentum transfer $-t$ distribution, the general
 1781 theoretical concern is that the process lacks of a hard scale because the mass of ρ^0 is rather
 1782 small. Therefore, perturbative calculations of QCD are difficult to be carried out. In addi-
 1783 tion, the scale dependence of the photoproduction process is also of great interest, which can
 1784 be only achieved by varying the mass of the Vector-Meson in photoproduction. Therefore,
 1785 heavier vector-mesons, e.g., ϕ and J/ψ , are important to be measured.

1786 In Fig. 61, the STAR preliminary results on J/ψ photoproduction are shown in Au+Au
 1787 UPC at 200 GeV. The differential cross section of $d\sigma/dp_T^2$ as a function of p_T^2 is presented,

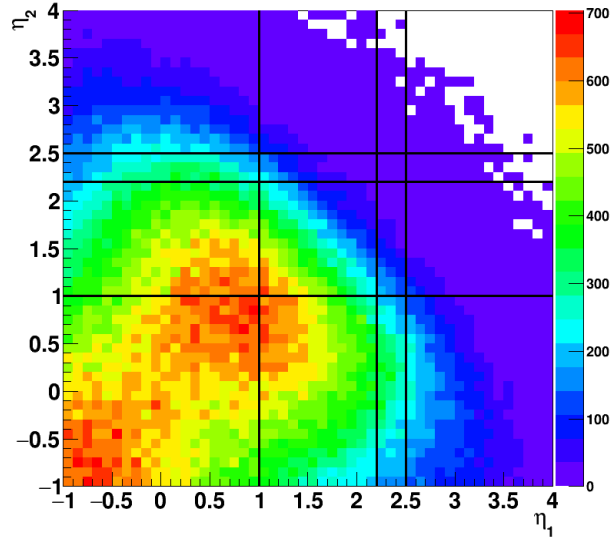


Figure 62: Pseudorapidity distribution of daughter electrons from the J/ψ decay using STARLight MC simulations. Lines are boundary acceptance of Barrel, Endcap, and Forward upgrade detectors.

1788 with the total contribution (left) and coherent contribution only (right). The data has
 1789 been compared with leading Monte Carlo models STARlight and Sartre, where a much
 1790 better description by Sartre is found. This is the first differential measurement of J/ψ
 1791 photoproduction off gold nucleus at the center-of-mass energy between photon and nucleon
 1792 (proton or neutron), $W \sim 25$ GeV, which provides important constraints to the gluon density
 1793 and its spatial distribution at this kinematic region, $x_g \sim 0.01$. The observed suppression of
 1794 the gluon density from this data, compared to the Impulse Approximation, is found to be
 1795 15-20%.

1796 Since the data presented above was taken in 2016, the acceptance of J/ψ is limited
 1797 to rapidity $y < 1$ due to the η acceptance of the daughter electrons. However, this can be
 1798 significantly improved in Run 2023 and 2025 Au+Au at 200 GeV with the endcap EMC, inner
 1799 TPC, and forward upgrade detectors. The extension of acceptance in rapidity to $1 < y < 1.5$
 1800 can lead to a lower x down to 4×10^{-3} , which overlaps with the LHC kinematics, as well as
 1801 going to higher x up to 0.05. With the forward upgrades, $y > 2.5$, the kinematic coverage
 1802 will be even wider, where STAR can cover a regime that is complementary to the LHC, e.g.,
 1803 the anti-shadowing region $x_g \sim 0.1$.

1804 In Fig. 62, it shows the pseudorapidity distribution of both daughter electrons from the
 1805 J/ψ decay, simulated by the STARLight MC model. The lines are boundaries of the barrel,
 1806 endcap, and forward detector acceptances. By extending to the endcap and forward, there is
 1807 a significant improvement in the J/ψ acceptance. Based on the established UPC J/ψ trigger
 1808 using both barrel and endcap, a high statistics event sample can be collected.

1809 When extending the acceptance of J/ψ to higher rapidity, there is a long standing issue
 1810 of photon energy ambiguity. At a J/ψ rapidity that $y \neq 0$, the photon energy can be

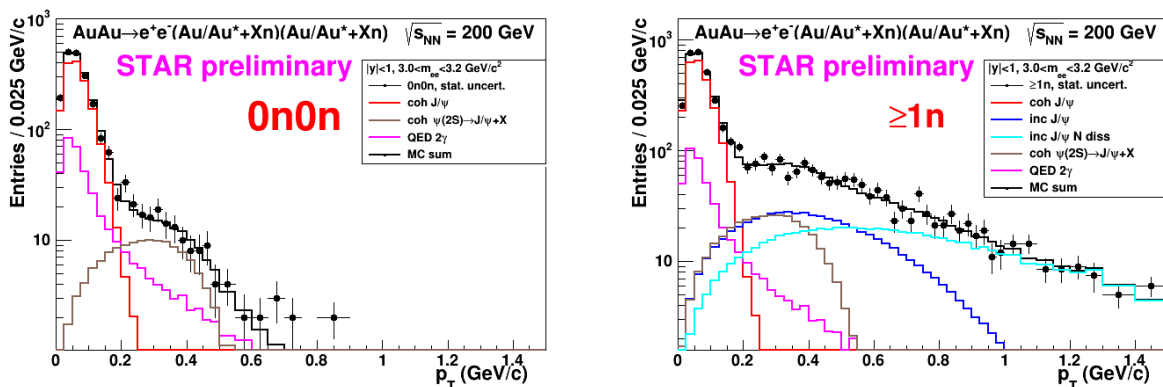


Figure 63: Uncorrected p_T of J/ψ mesons fitted with different contributions in Au+Au UPC at 200 GeV with no neutron on either side (left) and at least 1 neutron on either side (right).

1811 $(M_J/2)e^{\pm y}$, which corresponds to a higher and lower photon energy, respectively. However,
 1812 thanks to the neutron tagging in the ZDCs, this ambiguity can be resolved by considering
 1813 different neutron multiplicities and their theoretical expected photon fluxes [154]. The STAR
 1814 analysis using this method has just begun. In order to qualitatively see the difference by
 1815 introducing different neutron tagging classes, see Fig. 63. For details of this method, see
 1816 Ref. [154].

1817 Finally, for the STAR upcoming Run 2023 and 2025, there is an opportunity for measur-
 1818 ing the photoproduction of ϕ meson for the first time. The experimental challenge of this
 1819 measurement is that ϕ is usually reconstructed via the kaon channel. However, for photo-
 1820 production process, the momentum of the kaon daughters are very soft, ~ 100 MeV/c, such
 1821 that reconstructing the daughter tracks has been impossible with only the TPC. However,
 1822 for the upcoming runs, the inner TPC could push the low momentum tracking down to
 1823 ~ 100 MeV/c. There are two ways to achieve a statistical significant event sample of UPC
 1824 ϕ meson.

1825 The first one is to use ZDC coincidence trigger with no TOF requirement at the full
 1826 magnetic field in STAR, while the second one is to use the standard TOF-base UPC Vector-
 1827 Meson trigger at half-field. At full field, although the inner TPC can reconstruct tracks
 1828 down to ~ 100 MeV/c, it would not reach TOF for triggers due to the small bending radius.
 1829 Therefore, events can be collected without a dedicated UPC ϕ trigger. This requires a large
 1830 integrated luminosity to reach a few thousand raw ϕ events, based on the recent study using
 1831 2019 Au+Au data. However, if STAR can be run at half field, the TOF-base trigger might
 1832 be applicable. See Fig. 64 for illustration of the TOF-based trigger acceptance in kaon p_T .
 1833 Detail simulations will be followed up for the half-field running.

1834 With all three Vector-Meson (ρ^0 , ϕ , and J/ψ) measured at STAR in Au+Au UPC, they
 1835 will provide an unprecedented understanding of the diffractive process off the gold nucleus
 1836 in photoproduction, laying the foundation for such physics at the EIC.

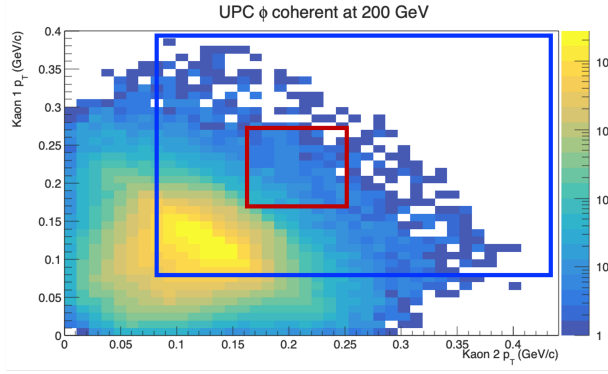


Figure 64: UPC ϕ meson decay p_T distributions of daughter 1 vs 2. The red box is the acceptance in p_T if requiring track to reach the location of TOF at STAR's full magnetic field; blue box is showing the same but with STAR at the half-field running.

1837 Vector-Meson decay: probing gluon distribution inside the nucleus

1838

1839 STAR recently observed a significant $\cos 2\Delta\phi$ azimuthal modulation in $\pi^+\pi^-$ pairs from
 1840 photonuclear ρ^0 and continuum production. The structure of the observed modulation as
 1841 a function of the $\pi^+\pi^-$ pair P_\perp , appears related to the diffractive pattern. Recent theoret-
 1842 ical calculations [155], which implemented linearly polarized photons interacting with the
 1843 saturated gluons inside a nucleus, have successfully described the qualitative features of the
 1844 observed modulation (see Fig. 65), and indicate that the detailed structure of the $\cos 2\Delta\phi$
 1845 modulation vs. P_\perp is sensitive to the nuclear geometry and gluon distribution. Data from
 1846 Run-23 and Run-25 would allow the additional statistical reach needed to perform multi-
 1847 differential analysis, providing stronger theoretical constraints. Specifically, multi-differential
 1848 analysis of the $\cos 2\Delta\phi$ modulation with respect to pair rapidity and pair mass are needed.
 1849 Multi-differential analysis with respect to pair mass is needed to separate the ρ^0 produc-
 1850 tion from the continuum Drell-Soding production. Multi-differential analysis with respect
 1851 to the pair rapidity is needed to quantitatively investigate how the double-slit interference
 1852 mechanism effects the structure of the observed azimuthal modulation. Additional statisti-
 1853 cal precision is also needed for measurement of the higher harmonics. Similar measurements
 1854 with $J/\Psi \rightarrow e^+e^-$ can be performed and such measurements at higher mass provide better
 1855 comparison with more reliable QCD calculation.

1856

1857 Ultraperipheral AA collisions, where photons generated by the Lorentz-boosted electro-
 1858 magnetic field of one nucleus interact with the gluons inside the other nucleus, can provide
 1859 certain 3D gluonic tomography measurements of heavy ions, even before the operation of
 1860 the future EIC. STAR has performed experimental measurements of the photoproduction
 1861 of J/ψ at low p_T in non-UPC heavy-ion collisions [156], accompanying the violent hadronic
 1862 collisions. A detailed study with p_T distributions has shown that the $|t|$ distribution in
 1863 peripheral collisions is more consistent with the coherent diffractive process than the inco-
 1864 herent process. Although models [157, 158] incorporating different partial coherent photon
 and nuclear interactions could explain the yields, it remains unclear how the coherent process

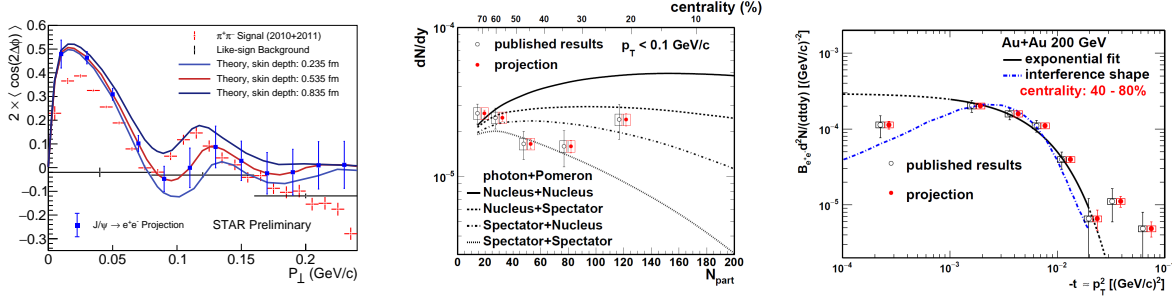


Figure 65: Left: Measurement of the $\cos 2\Delta\phi$ modulation of $\pi^+\pi^-$ pairs from photonuclear ρ^0 and continuum production compared to theoretical predictions [155]. Projections are shown for a similar measurement of the azimuthal modulation of e^+e^- pairs from photonuclear production of the J/ψ . Center: Projection of the dN/dy of photoproduced J/ψ in non-UPC events vs. the event centrality (N_{part}) compared to various theoretical production scenarios. Right: Projection of the t spectra of photoproduced J/ψ in 40 – 80% central collisions.

1865 happens and whether final-state effects play any role [159]. Resolving this puzzle with high
 1866 statistical data and detailed $|t|$ distributions at different centralities at RHIC as projected
 1867 for Run-23+25 in Fig. 65 may be important for understanding what defines the coherenceness
 1868 of the photoproduction, how vector mesons are formed in the process and how exclusive the
 1869 similar process has to be in future EIC experiments with forward neutron veto/tagging.

1870 Photoproduction of dijets

1871

1872 In addition to photoproduction of Vector-Meson, photoproduction of back-to-back jets
 1873 has been increasingly interested in the context of nuclear PDF. The process is a two-to-two
 1874 hard scattering between a direct or resolved photon from the projectile (photon from UPC)
 1875 and the quarks or gluons from the nucleus target. The final-state is a pair of back-to-back jet,
 1876 which is directly sensitive to the photon and nuclear structure in terms of parton distribution
 1877 functions. At the LHC, this process corresponds to the kinematic region $x_A \sim 10^{-3}$, which
 1878 is the gluon dominated regime. Here we propose to measure the photoproduction dijets
 1879 at STAR, where kinematic regions, e.g., the anti-shadowing and the EMC region, can be
 1880 reached. This measurement has never been done at RHIC and will provide a significant
 1881 constraints to the nPDFs of heavy nucleus at this kinematics for photoproduction.

1882 The pseudo-data from eA collisions used here is generated by BeAGLE (**Benchmark eA**
 1883 **Generator for LEptoproduction**) [153], based on the lepton and gold beam energy of 18×100
 1884 GeV, where the input PDF for the the exchanged photon is the CTEQ 5 from the LHAPDF
 1885 library [160] and EPS09 for the nuclear PDF.

1886 Jets are reconstructed by FastJet [161] with the anti- k_T algorithm, which is based on
 1887 the energy distribution of final state particles in the angular space. All the stable and
 1888 visible particles produced in the collisions with $p_T > 250$ MeV/c and $-1.5 < \eta < 1.5$ and
 1889 $2.5 < \eta < 4.0$ in the laboratory system are taken as input. The jet cone radius parameter
 1890 has been set to $R_{\text{jet}} = 1$ in the jet finding algorithm. To obtain the events in Au+Au UPC

1891 collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV from simulations of $e\text{Au}$ at 18×100 GeV, an event-by-event
 1892 weight is applied according to the photon flux difference between $e\text{A}$ and $\text{Au}+\text{Au}$ UPC
 1893 collisions.

1894 After reweighing we obtain the dijet events with the pseudorapidity of jets (η^{jet}) from -1.5
 1895 to 1.5 in middle rapidity region and $2.5 < \eta^{\text{jet}} < 4.0$ in the forward region. In each event, the
 1896 jet with the highest p_T is called the trigger jet, the jet with the second highest p_T is called
 1897 the associate jet. Events are selected with the requirement that the trigger jet has $p_T^{\text{trig}} > 5$
 1898 GeV/c and the associated one has $p_T^{\text{asso}} > 4.5$ GeV/c. 100 M event are generated, after all
 1899 cuts applied, we found ~ 5600 dijet events corresponding to the integrated luminosity $L =$
 1900 9 nb^{-1} . Therefore, with STAR Run 2023 and 2025 $\text{Au}+\text{Au}$ collisions, an event sample of
 1901 dijets of 50-60k is expected.

1902 In 200 GeV AuAu UPC collisions, the distributions of jets' pseudorapidity and p_T can
 1903 be found in Fig. 66. Jets dominate at $\eta \sim 0.5$ with the maximum $p_T \sim 20$ GeV/c.

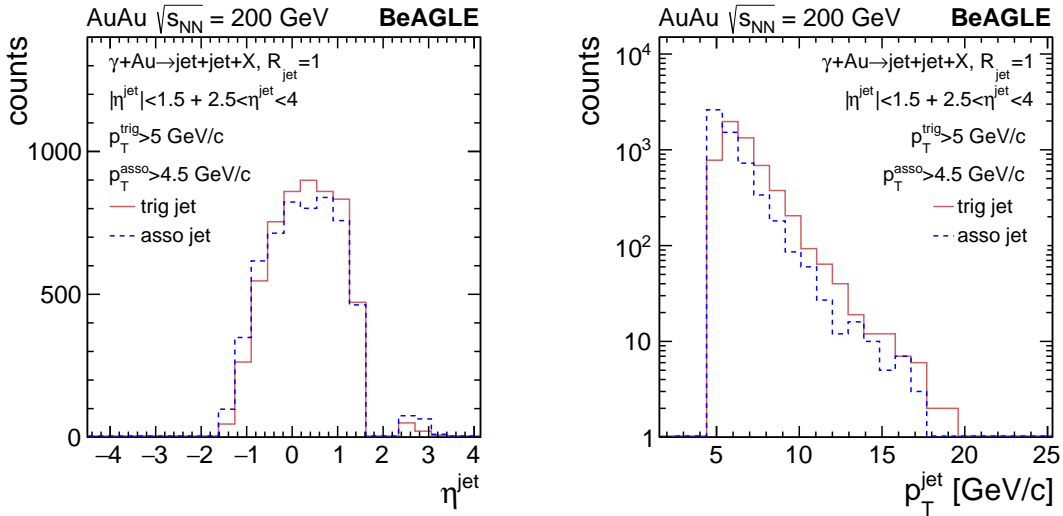


Figure 66: In $\text{Au}+\text{Au}$ UPC collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV, the dijet events are selected with $|\eta^{\text{jet}}| < 1.5 + 2.5 < \eta^{\text{jet}} < 4.0$. For the trigger jet: $p_T^{\text{trig}} > 5$ GeV/c, associate jet: $p_T^{\text{asso}} > 4.5$ GeV/c. Left: the pseudorapidity distributions of the trigger and associated jets; right: the p_T distributions of the trigger and associated jets.

1904 In BeAGLE, depending on the wave function components for the incoming virtual pho-
 1905 ton, the major hard processes are divided into three classes: the direct processes, the soft
 1906 VMD processes and the resolved processes (hard VMD and anomalous). The direct pho-
 1907 ton interacts as a point-like particle with the partons of the nucleon, major subprocesses
 1908 in direct category: LO DIS, Photon-Gluon Fusion (PGF) and QCD Compton (QCDC).
 1909 While the VMD and anomalous components interact through their hadronic structure. Re-
 1910 solved photon processes play a significant part in the production of hard high- p_T processes
 1911 at $Q^2 \approx 0$. The following hard subprocesses are grouped in the resolved processes category:

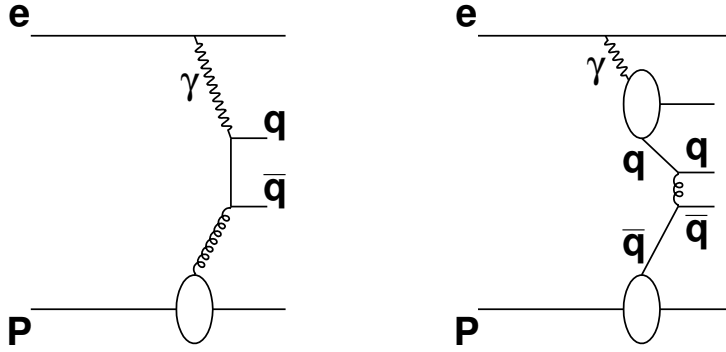


Figure 67: Examples of diagrams for direct (left) and resolved (right) processes in electron-proton scattering. In UPC, the photon emitter is replaced with the Au nucleus.

1912 $qq \rightarrow qq, q\bar{q} \rightarrow q\bar{q}, q\bar{q} \rightarrow gg, qq \rightarrow qq, gg \rightarrow q\bar{q}, gg \rightarrow gg$. The examples of Feynman
 1913 diagrams of resolved and direct processes are shown in FIG. 67.

1914 The momentum fraction of the parton from the exchanged photon (x_γ) and the momen-
 1915 tum fraction of the parton from the gold beam (x_{Au}) can be reconstructed knowing the
 1916 momentum and angles of dijets as

$$x_\gamma = \frac{1}{2E_\gamma} (p_T^{\text{trig}} e^{-\eta_{\text{trig}}} + p_T^{\text{asso}} e^{-\eta_{\text{asso}}}) \quad (4)$$

1917

$$x_{\text{Au}} = \frac{1}{2E_{\text{Au}}} (p_T^{\text{trig}} e^{\eta_{\text{trig}}} + p_T^{\text{asso}} e^{\eta_{\text{asso}}}) \quad (5)$$

1918 where E_γ is the photon energy which can be determined from the hadronic final-state, see
 1919 later for details. Eq. 4 and Eq. 5 are valid in the lab frame in LO.

1920 The reconstructed x_γ and x_{Au} in AuAu UPC dijet events can be seen from Fig. 68. The
 1921 reconstructed x_γ covers a wide range from 0.2 to 0.9 in resolved process, and dominates at
 1922 high x in direct process. The reconstructed x_{Au} distributions contain two peaks as there
 1923 are two pseudorapidity regions. The forward pseudorapidity ($2.5 < \eta^{\text{jct}} < 4.0$) leads to the
 1924 peak at high $x_{\text{Au}} \sim 0.5$, while middle rapidity jets ($|\eta^{\text{jct}}| < 1.5$) contribute the peak at x_{Au}
 1925 $< \sim 0.2$. With the Run 2023 and 2025 data of Au+Au and Run 2024 p^\uparrow +Au at STAR, this
 1926 will become the first measurement at this kinematic region at RHIC with good statistical
 1927 precision.

1928 Taking one step further, the exclusive or diffractive dijets can also be measured in $p^\uparrow p^\uparrow$,
 1929 p^\uparrow +Au, and Au+Au at $\sqrt{s_{\text{NN}}} = 200$ GeV. The process is diffractive such that there are only
 1930 two jets in the event, where the target nucleon or nucleus stay intact. Similar to exclusive
 1931 Vector-Meson production discussed earlier, the exclusive dijets can provide a large impact
 1932 in understanding the nucleon and nuclear structure over a wide range of kinematics. In
 1933 addition, with the unique target polarization at RHIC, the exclusive dijets can be sensitive
 1934 to Generalized Parton Distributions and p_T Dependent PDFs. This process is expected to

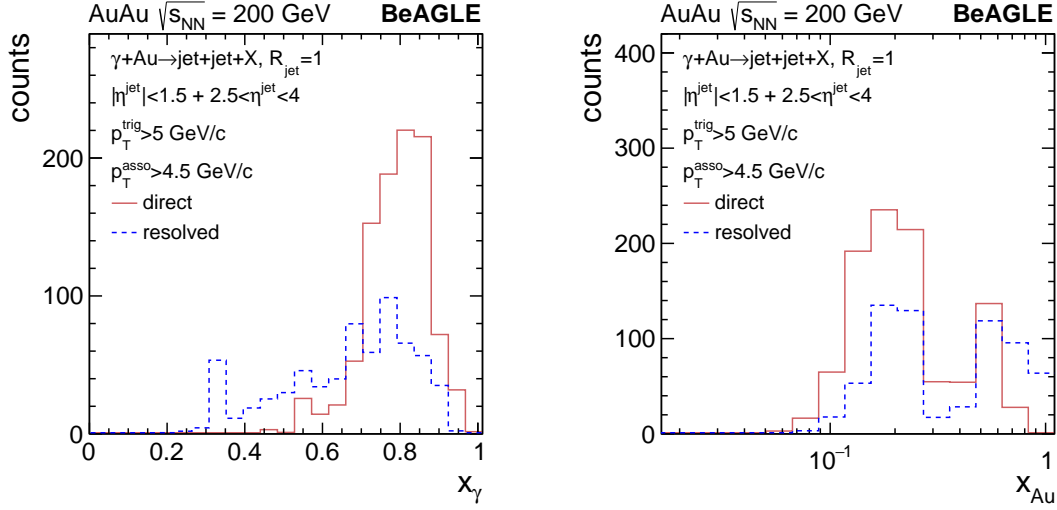


Figure 68: In AuAu UPC collisions at $\sqrt{s_{NN}} = 200$ GeV, the dijet events are selected with $|\eta^{\text{jet}}| < 1.5 + 2.5 < \eta^{\text{jet}} < 4.0$. For the trigger jet: $p_T^{\text{trig}} > 5$ GeV/c, associate jet: $p_T^{\text{asso}} > 4.5$ GeV/c. Left: the x_γ distributions in resolved and direct processes; right: the x_{Au} distributions in resolved and direct processes.

1935 be complementary to the process discussed in Sec. 3.1. In Fig. 69, the diffractive dijets
 1936 photoproduction in $p+\text{Au}$ UPCs are shown, with the transverse energy (E_T) on the left
 1937 panel and the dijet η separation distribution on the right panel. For a first look, the STAR
 1938 Upcoming run 2024 would have enough luminosity to achieve reasonable statistics of this
 1939 measurement; the same measurement can be done in $p+p$ and Au+Au collisions.

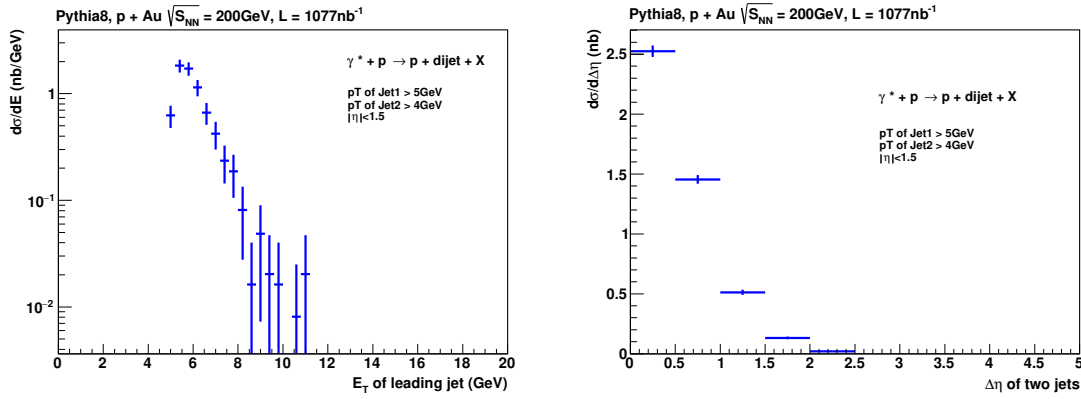


Figure 69: In $p+\text{Au}$ UPC collisions at $\sqrt{s_{NN}} = 200$ GeV, the diffractive dijet events are selected with the trigger jet: $p_T^{\text{trig}} > 5$ GeV/c and associate jet: $p_T^{\text{asso}} > 4.0$ GeV/c. The E_T distributions of the leading jet (left) and $\Delta\eta$ of the dijets distributions (right) are shown with $\sim 1 \mu\text{b}^{-1}$ integrated luminosity.

1940 Additional opportunities are available for STAR Run 2023-2025 based on UPC jets, e.g.,
 1941 measurement of diffractive dijets off polarized proton target, and azimuthal correlation of
 1942 the dijets, which will be sensitive to nPDFs, diffractive nPDFs, QCD factorisation breaking,
 1943 and spin structure of the proton. Here we do not elaborate them in details but defer the
 1944 readers to Refs. [151, 152, 162–165] for both UPCs and at the EIC.

1945 Photoproduction of inclusive charged particles and cross sections

1946
 1947 Inclusive photoproduction processes in high-energy ep collisions have been extensively
 1948 studied at HERA, e.g., charged particle productions, inclusive cross section, heavy-flavor
 1949 production, etc. Recently, there have been efforts re-analyzing the HERA data in photo-
 1950 production and deep inelastic scattering to look for collectivity in terms of azimuthal cor-
 1951 relations [166], inspired by the outstanding flow phenomena in heavy-ion collisions. At the
 1952 LHC, experiments have just begun using the UPCs to look at collisions between photons and
 1953 heavy nuclei in photoproduction, primarily to search for the collective phenomena. However,
 1954 inclusive photoproduction processes in nuclei at high energy remains largely unexplored.

1955 Inclusive photoproduction process is generally challenging for the UPC in heavy-ion ex-
 1956 periments. At HERA, photoproduction in ep scattering can be unambiguously identified
 1957 by the small angle electron taggers, where event kinematics can be reconstructed. How-
 1958 ever, in heavy-ion UPCs, the photon emitting nucleus is invisible to the experiment, leaving
 1959 the kinematics, e.g., W , largely unconstrained. In a recent study using general-purpose eA
 1960 MC model BeAGLE, it is found that the event kinematic reconstruction in UPC can be
 1961 approached based on the hadronic final-state (HFS).

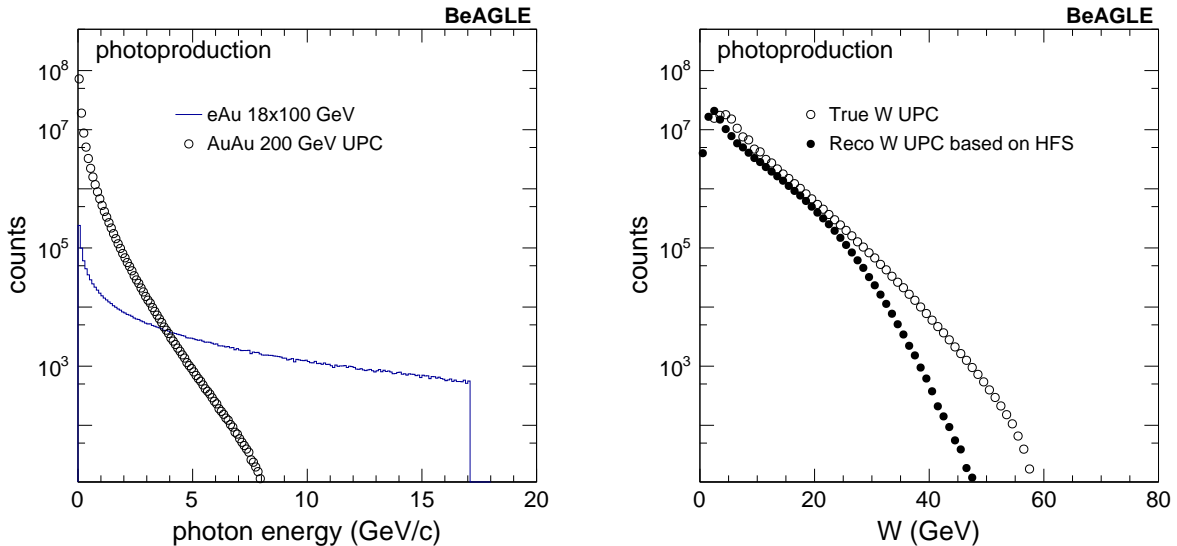


Figure 70: Left: photon energy distribution in eA and Au+Au UPC. Right: The truth level W in Au+Au UPC and the corresponding reconstructed level based on the HFS method.

1962 In Fig. 70 left, it shows the photon energy distribution based on MC simulation of BeA-
 1963 GLE of $e\text{Au } 18 \times 100 \text{ GeV}$. In addition, by using the photon flux generated by the UPC
 1964 at 200 GeV Au+Au collisions, the photon energy spectra is reweighted and shown as the
 1965 open circle. The low photon energy is greatly enhanced due to the large flux generated by
 1966 the heavy nucleus, while the spectra is much steeper than in the $e\text{Au}$ collisions. In Fig. 70
 1967 right, the HFS method has been adopted to reconstruct the kinematic variable W , based
 1968 on the STAR acceptance including the forward upgrade detectors. The smearing from truth
 1969 to reconstructed W is visible and stronger at large W . However, by selecting on the re-
 1970 constructed W , the event kinematics can be better controlled than using the average only.
 1971 Unfolding technique can be used here for correcting the bin migration in W as well. Note
 1972 that there is no detector simulations shown here.

1973 In the upcoming RHIC Run 2023 and Run 2025, the inclusive photoproduction is of
 1974 great interest. The cross section of such events is generally large, while a different trigger is
 1975 required comparing to the standard minimum-bias hadronic collision trigger. The baseline
 1976 trigger has been developed during the Au+Au 200 GeV data taken in 2019, where only a
 1977 ZDC coincidence was required. For Run 2023 and 2025, asymmetry BBC response could be
 1978 added to more efficiently select the inclusive photoproduction process.

1979 Search for Collectivity in Photo-nuclear ($\gamma+\text{Au}$) Processes

1980
 1981 Until the EIC is built, high-energy photoproduction processes (low virtuality limit of
 1982 the deep inelastic scattering) can be studied using ultra-peripheral ion collisions (UPCs)
 1983 that occur when two heavy ions interact at large impact parameters. Such collisions can
 1984 be considered as $\gamma+\text{Au}$ processes but unlike at the EIC, the photons involved in UPCs are
 1985 quasi-real. For UPCs at top RHIC energies one expects the energy of the quasi-real photon
 1986 to be approximately $E_\gamma \approx 3 \text{ GeV}$. The typical range of the center of mass energy of the
 1987 photon-nucleon system will therefore be $W_{\gamma N} \approx 40 \text{ GeV}$. Therefore, Au+Au collisions at
 1988 $\sqrt{s_{NN}} = 200 \text{ GeV}$ will provide access to the $\gamma+\text{Au}$ process at 40 GeV center of mass energy.
 1989 Our specific interest is high activity inclusive $\gamma+\text{Au}$ process to search for collectively and
 1990 improve our understanding of the mechanism of baryon stopping.

1991 A satisfactory microscopic explanation of how collectivity originates from the basic pro-
 1992 cesses of QCD and evolves with collision system size is a topic of broad interest in the
 1993 community of high energy nuclear physics. The formation of a quark-gluon plasma medium
 1994 and its fluid-dynamic expansion explain the origin of collectivity in Au+Au collisions. Re-
 1995 sults from RHIC small system scan indicate fluid-dynamic expansion are essential to drive
 1996 collectivity in $^3\text{He}/\text{d}/\text{p}+\text{Au}$ collisions [167]. A search for collectivity in $\gamma+\text{Au}$ interactions
 1997 at RHIC will be a natural continuation of the recent system size scan [167], extending it at
 1998 the small end to complete the hierarchy: $\text{Au}+\text{Au} > ^3\text{He}+\text{Au} > \text{d}+\text{Au} > \text{p}+\text{Au} > \gamma+\text{Au}$.
 1999 This will help better address how collectivity originates and evolves with system size. If
 2000 collectivity is observed in $\gamma+\text{Au}$ processes it can provide a way to explore the creation of
 2001 a many-body system exhibiting fluid behavior in photon-induced processes [168]. A recent
 2002 calculations in Ref [168] assumes $\gamma+\text{A}$ processes are equivalent to collisions of vector me-

2003 son with ions (ρ +A collisions) and describe first measurements of harmonic coefficients v_n
 2004 in photonuclear processes measured by the ATLAS collaboration [169]. The hypothesis of
 2005 γ +A process as ρ +A collisions and the formation of a fluid-dynamic medium can be tested
 2006 at RHIC in a data-driven way. This can be done by comparing measurements in γ +Au
 2007 processes at $W_{\gamma N} \approx 40$ GeV and in d+Au collisions at $\sqrt{s_{NN}} = 39$ GeV. The former will
 2008 be possible if a high statistics data set is collected for Au+Au collisions at $\sqrt{s_{NN}} = 200$
 2009 during the Run 23 and 25 and the later can be performed with the existing RHIC data
 2010 on tape. It is known from RHIC measurements, argument based on initial geometry and,
 2011 fluid dynamic calculations that elliptic anisotropy coefficient follow a hierarchy of $v_2(\text{d+Au})$
 2012 $> v_2(\text{p+Au})$ at a fixed collision energy and multiplicity [167, 170]. Following a similar ar-
 2013 gument one expects $v_2(\text{d+Au}) > v_2(\text{p+Au})$. In the fluid dynamic picture of Ref [168] the
 2014 elliptic anisotropy coefficient will show the following hierarchy: $v_2(\text{d+Au}) > v_2(\text{p+Au})$. A
 2015 similar test by comparing $v_2(\text{p+Pb})$ and $v_2(\text{p+Au})$ at the LHC is difficult since the center
 2016 of mass energy differs by a factor of six between p+Pb and p+Au collisions.

2017 Photonuclear processes can also be used to study the origin of baryon stopping and
 2018 baryon structure in general. One proposed mechanism for explaining the baryon stopping
 2019 is the baryon junction: a nonperturbative Y-shaped configuration gluons which is attached
 2020 to all three valence quarks. In this picture it is the baryon junction that carries the baryon
 2021 number rather than the valence quarks. The existence of baryon junctions and their inter-
 2022 action with the incoming target or projectile are theorized to be an effective mechanism for
 2023 substantial baryon stopping in pp and AA [171], but this has yet to be confirmed experi-
 2024 mentally. Photonuclear processes allow us to study baryon stopping in the simplest possible
 2025 process. The vast majority of these collisions occur through what is called the resolved pro-
 2026 cess where the quasi-real photon fluctuates into a quark-antiquark pair which then collides
 2027 with the other ion [?]. If the baryon number were carried by the three valence quarks, then
 2028 this quark-antiquark pair would not be able to stop the baryons, but it is possible for the
 2029 quark-antiquark pair to interact with the junction and produce a midrapidity baryon. An
 2030 added benefit is that photonuclear processes are highly asymmetric and baryons only enter
 2031 from one side of the collision. The baryon-junction stopping mechanism is predicted to cause
 2032 an exponential damping of the cross section with rapidity $\sim \exp(-\alpha_0^J(y - Y_{\text{beam}}))$, where
 2033 $\alpha_0^J \simeq 1/2$ is the Regge intercept of the baryon junction [171]. In a symmetric hadronic
 2034 collision, baryons are traveling from either direction so the stopping of both the target
 2035 ($\sim \exp(-\alpha_0^J(y - Y_{\text{beam}}))$) and the projectile ($\sim \exp(\alpha_0^J(y - Y_{\text{beam}}))$) will likely compensate
 2036 for each other, leading to a nearly symmetric distribution. But in an asymmetric system like
 2037 a photonuclear collision, this exponential shape should be visible.

2038 A handful of data sets exist on the disk with the appropriate event trigger selection
 2039 for studying photonuclear processes at RHIC. In Fig.71 we present preliminary results on
 2040 γ +Au-rich interactions using Au+Au 54 GeV data from STAR shown at the Quark Matter
 2041 2022 conference. By identifying the single neutron peak for individual ZDCs, we require the
 2042 cuts equivalent to 1nXn. We apply an asymmetric cut on east and west BBCs to improve
 2043 the purity. We also make sure the position of the primary vertex along collision direction V_z
 2044 from TPC and VPD detectors differs by about 10 cm. After applying such cuts on Au+Au

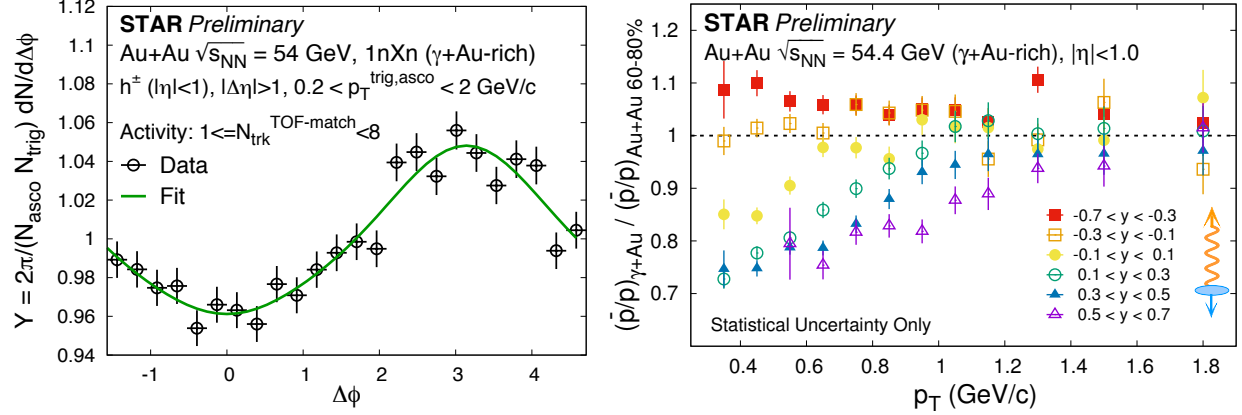


Figure 71: (Left) STAR preliminary data on normalized yield of long range di-hadron correlations in γ +Au-rich events with a relative pseudorapidity gap of $|\Delta\eta| > 1$ between two hadrons. The events are selected by applying asymmetric cuts on the energy deposition of neutrons in ZDCs (1nXn) and on TPC tracks matched with TOF $N_{\text{trk}}^{\text{TOF-match}}$ in the window of $1 \leq N_{\text{trk}}^{\text{TOF-match}} < 8$. The green curve represents a fit to data using a function: $1 + 2 \sum a_n \cos(n\Delta\phi)$. No signatures of collectivity associated with enhancement of correlation near relative azimuthal angle $\Delta\phi \sim 0$ is observed. (Right) The double ratio of antiprotons to protons in γ ++Au-rich events compared to peripheral Au+Au events, indicating significant enhancement of protons at low p_T and at mid-rapidity. The enhancement shows a strong rapidity dependence while going from the photon to ion direction.

2045 54 GeV data we perform measurements in γ +Au+Au rich events.

2046 Fig. 71 (left) shows the normalized yield, differential in relative azimuthal angle of the
 2047 trigger and associated particles $Y(\Delta\phi) = 1/N_{\text{trig}}/N_{\text{asco}} d^2 N^{\text{pair}}/d\Delta\phi$ integrated over a relative
 2048 pseudorapidity window of $|\Delta\eta| > 1$. For this analysis, the p_T of trigger and associated
 2049 particles is chosen to be within $0.2 < p_T^{\text{trig,asco}} < 2$ GeV/c. The distribution $Y(\Delta\phi)$ is shown
 2050 for two different bins of activity characterized by the number of TPC tracks matched with
 2051 the TOF $1 \leq N_{\text{trk}}^{\text{TOF}} < 8$ (low activity). The distribution is fitted using a Fourier function
 2052 of the form $(1 + 2 \sum a_n \cos(n\Delta\phi))$ (green curve). No ridge-like component associated with a
 2053 significant enhancement of $Y(\Delta\phi)$ near $\Delta\phi = 0$ that is related to the signature of collectivity
 2054 is seen.

2055 Fig. 71 (right) shows the measurement of the yield of anti-protons-to-protons (\bar{p}/p) with
 2056 p_T . The quantity plotted is a double ratio of \bar{p}/p for the measurements in γ +Au-rich events
 2057 over the same in 60-80% peripheral Au+Au events. We see a suppression of the \bar{p}/p yield
 2058 in γ +Au events at low $p_T < 0.6$ GeV/c and for the symmetric window of $-0.1 < y < 0.1$
 2059 around mid-rapidity. The suppression of \bar{p}/p yield gets stronger while going from the photon
 2060 to the ion direction, with the double ratio dropping by a factor 0.75 at low p_T . We have
 2061 checked that this trend is not seen for π^-/π^+ , K^-/K^+ and not explained by PYTHIA 6
 2062 model. This important observation provides the necessary impetus for further exploration
 2063 using various available data sets. In particular, we would like to test if this strong rapidity
 2064 dependence of the \bar{p}/p yield is consistent with the picture of baryon junction that predicts an
 2065 exponential dependence of stopping with rapidity of form $\exp(-\alpha(y - Y_{\text{beam}}))$ with $\alpha = 0.5$.

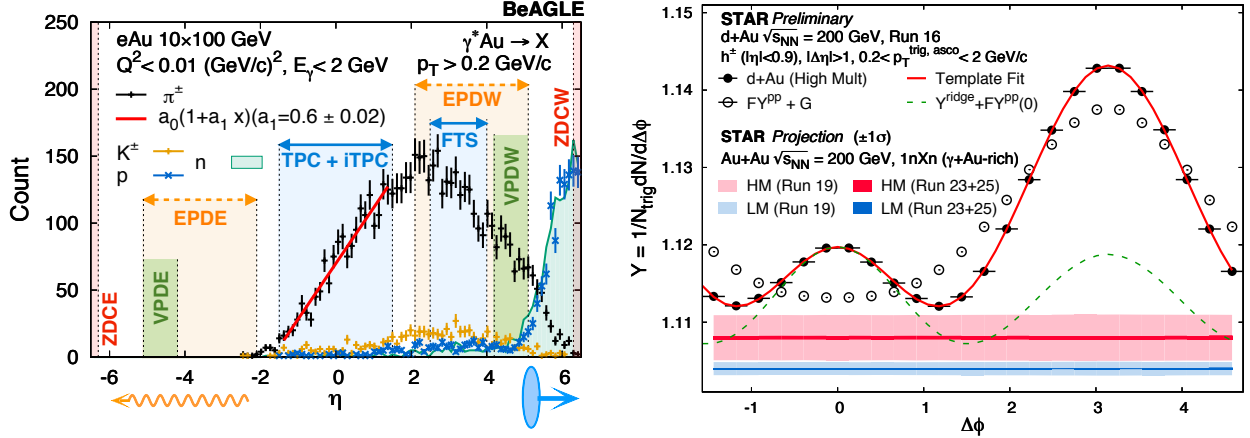


Figure 72: (Left) Pseudorapidity distribution of different particles using the state-of-the-art BeAGLE [172,173] event generator for the EIC in $e+Au$ events. By restricting the virtuality and energy of the photon (γ^*) we try to mimic the kinematics of a $\gamma + Au$ ($Au+Au$ UPC) event. The purpose of this plot is to demonstrate how different STAR detectors will be used to identify such UPC processes. (Right) STAR preliminary data on per-trigger yield estimated using di-hadron correlations in $d+Au$ (hadronic) 200 GeV collisions. The correlation function in pp collisions (open circle) is used as a template to fit the same in relatively high multiplicity $d+Au$ collisions (solid circle) and to extract the long-range ridge-like component. The red and blue band show projections for $\gamma + Au$ enriched events for two different multiplicity bins. The aim is to use the correlation function from the low multiplicity $\gamma + Au$ to perform template fit in the high multiplicity bin.

2066 Our aim will be extend these measurements with high statistics $\gamma+Au$ -rich event samples
 2067 using Run 2023 and 2025 data on $Au+Au$ collisions at $\sqrt{s_{NN}} = 200$ GeV. Fig.72(Left) shows
 2068 the pseudorapidity (η) distribution of identified particles with $p_T > 0.2$ GeV/c in inclusive
 2069 $e+Au$ DIS (γ^*+Au , where γ^* refers to a virtual photon) processes simulated using the EIC
 2070 Monte Carlo BeAGLE event generator [172,173] with electron and ion beam energy of 10 and
 2071 100 GeV, respectively. The virtuality of the exchanged photon is restricted to be $Q^2 < 0.01$
 2072 GeV/c² and photon energy is restricted to be $E_\gamma < 2$ GeV to mimic $\gamma+Au$ interactions in
 2073 $Au+Au$ UPCs at $\sqrt{s_{NN}}=200$ GeV. This figure demonstrates how the combination of the
 2074 inner Time Projection Chamber (iTPC), the new highly granular Event-Plane Detectors
 2075 (EPD) and forward tracking system (FTS) and the Zero-Degree Calorimeters (ZDC) can
 2076 be used to isolate $\gamma+Au$ events from peripheral $Au+Au$ events (symmetric in η with no
 2077 gaps). In terms of triggering the $\gamma+Au$ interactions, the most stringent selection criterion
 2078 is that the ZDCE detector should be restricted to have a single neutron hit (1n), while
 2079 no restriction (Xn) should be placed on the ZDCW to trigger on $\gamma+Au$ candidates with
 2080 east-going photons, and vice versa. We perform a feasibility study using Run 19 data on
 2081 min-bias $Au+Au$ collisions using about 130 M events. Fig.72 shows STAR preliminary data
 2082 on the per-trigger yield in di-hadron correlations in $d+Au$ events where a clear ridge can
 2083 be seen after template fitting. On the same plot we show projections of uncertainties for
 2084 the di-hadron correlations in possible $\gamma+Au$ -rich events using $Au+Au$ 200 GeV data from
 2085 Run 19 (130 M events) and using $Au+Au$ 200 GeV data from anticipated Run 23 and

2086 Run 25 (20 B events). Projections are shown for high activity (HM) and low activity (LM)
 2087 event classes determined by the uncorrected track multiplicity in TPC matched with TOF of
 2088 $15 \leq N_{\text{trk}}^{\text{TOF}} < 25$ and $1 \leq N_{\text{trk}}^{\text{TOF}} < 8$, respectively. Even without any dedicated trigger, 20
 2089 B minbias Au+Au events can already give us enough $\gamma+Au$ candidates to significantly reduce
 2090 the uncertainties shown by the red and blue projection bands in Fig.72. This will enable us
 2091 to perform differential measurements of di-hadron correlations with different combinations
 2092 of triggers and associated p_T and perform a search for collectivity and in addition to testing
 2093 the baryon-junction conjecture.

2094 Other inclusive photoproduction measurements

2095
 2096 Besides the search for collectivity in photon-nucleus collisions, there are many other
 2097 inclusive photoproduction processes are of great interest. In the upcoming Run 2023-2025,
 2098 inclusive photoproduction processes only require a large sample of "minimum-bias" photo-
 2099 nucleus collision events, instead of special triggered events.

2100 For example, one measurement that will have a large impact is the inclusive J/ψ pho-
 2101 toproduction. Note that STAR has results on exclusive J/ψ photoproduction, the comple-
 2102 mentary inclusive measurement (together with exclusive measurements) can be sensitive to
 2103 the saturation or non-linear gluon dynamics. The observable is as follows,

$$\frac{\sigma_{J/\psi}^{\text{exclusive}}/\sigma_{J/\psi}^{\text{inclusive}}|_{\text{Au}}}{\sigma_{J/\psi}^{\text{exclusive}}/\sigma_{J/\psi}^{\text{inclusive}}|_{\text{p}}}. \quad (6)$$

2104 The J/ψ in the inclusive photoproduction provides a hard scale that theoretical calcu-
 2105 lations, e.g., dipole model, can be performed. Qualitatively, the nuclear shadowing model
 2106 (Leading Twist Approximation [146–148]) predicts this double ratio to be below unity, while
 2107 saturation models predict above unity [174]. This is one of the very few observables that
 2108 qualitatively separates these two long standing models. In the upcoming STAR runs of
 2109 Au+Au and p +Au collisions, this measurement will play an important role in understand-
 2110 ing the saturation phenomena before the EIC. For the EIC measurement, see Fig. 73 for
 2111 details. The reason we can do this similar measurement in UPCs is because we can replace
 2112 the inclusive DIS measurement (finite Q^2) with inclusive photoproduction of J/ψ , where the
 2113 charm quark mass provides the hard scale.

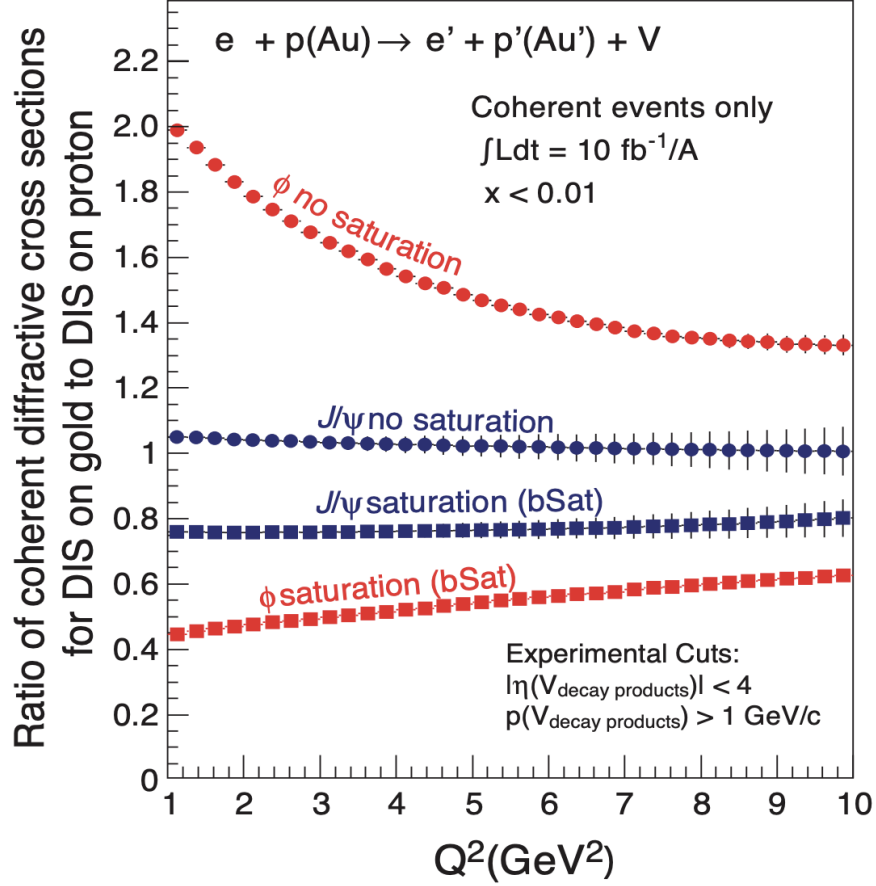


Figure 73: Figure from the EIC White Paper - Fig 1.6 [174]. The ratio of the coherent diffractive cross-section in $e+\text{Au}$ to $e+p$ collisions normalized by $A^{4/3}$ and plotted as a function of Q^2 for both saturation and non-saturation models. The $1/Q$ is effectively the initial size of the quark-antiquark systems (ϕ and J/ψ) produced in the medium.

2114 3 Run-24 Request for Polarized pp and $p+A$ Collisions at 2115 200 GeV

2116 The exploration of the fundamental structure of strongly interacting matter has always
2117 thrived on the complementarity of lepton scattering and purely hadronic probes. As the
2118 community eagerly anticipates the future Electron Ion Collider (EIC), an outstanding sci-
2119 entific opportunity remains to complete “must-do” measurements in $p+p$ and $p+A$ physics
2120 during the final years of RHIC. These measurements will be essential if we are to fully real-
2121 ize the scientific promise of the EIC, by providing a comprehensive set of measurements in
2122 hadronic collisions that, when combined with future data from the EIC, will establish the
2123 validity and limits of factorization and universality. Much of the Run-24 physics program
2124 outlined here is, on the one hand, unique to proton-proton and proton-nucleus collisions and
2125 offers discovery potential on its own. On the other hand, these studies will lay the ground-
2126 work for the EIC, both scientifically and in terms of refining the experimental requirements
2127 of the physics program, and thus are the natural next steps on the path to the EIC. When
2128 combined with data from the EIC these STAR results will provide a broad foundation to a
2129 deeper understanding of fundamental QCD.

2130 The separation between the intrinsic properties of hadrons and interaction-dependent
2131 dynamics, formalized by the concept of factorization, is a cornerstone of QCD and largely
2132 responsible for the predictive power of the theory in many contexts. While this concept
2133 and the associated notion of universality of the quantities that describe hadron structure
2134 have been successfully tested for unpolarized and, to a lesser extent, longitudinally polarized
2135 parton densities, its experimental validation remains an unfinished task for much of what the
2136 EIC is designed to study – the three-dimensional structure of the proton and the physics of
2137 dense partonic systems in heavy nuclei. To establish the validity and limits of factorization
2138 and universality, it is essential to have data from *both* lepton-ion and proton-ion collisions,
2139 with experimental accuracy that makes quantitative comparisons meaningful.

2140 Run-24, with polarized $p+p$ and $p+Au$ collisions at $\sqrt{s_{NN}} = 200$ GeV, will likely be the
2141 last RHIC spin/cold QCD run. This run will provide STAR with the unique opportunity to
2142 investigate these 200 GeV collision systems with the Forward Upgrade providing full track-
2143 ing and calorimetry coverage over the region $2.5 < \eta < 4$ and the iTPC providing enhanced
2144 particle identification and expanded pseudorapidity coverage at mid-rapidity. These power-
2145 ful detection capabilities, when combined with substantially increased sampled luminosity
2146 compared to Run-15, will enable critical measurements to probe universality and factoriza-
2147 tion in transverse spin phenomena and nuclear PDFs and fragmentation functions, as well as
2148 low- x non-linear gluon dynamics characteristic of the onset of saturation. This will provide
2149 unique insights into fundamental QCD questions in the near term, and essential baseline
2150 information for precision universality tests when combined with measurements from the EIC
2151 in the future.

2152 We therefore request at least 11 weeks of polarized $p+p$ data-taking at $\sqrt{s} = 200$ GeV and
2153 11 weeks of polarized $p+Au$ data-taking at $\sqrt{s_{NN}} = 200$ GeV during Run-24. Effectively,
2154 we request approximately equal nucleon-nucleon luminosities for $p+p$ and $p+Au$ which is

2155 essential to optimize several critical measurements that require comparisons of the same
2156 observable in (polarized or unpolarized) $p+p$ and $p+Au$ collisions described in the following
2157 sections.

2158 All of the running will involve vertically polarized protons. Based on recent C-AD guid-
2159 ance, we expect to sample at least 208 pb^{-1} of $p+p$ collisions and 1.2 pb^{-1} of $p+Au$ collisions.
2160 These totals represent 4 times the luminosity that STAR sampled during transversely po-
2161 larized $p+p$ collisions in Run-15 and 2.7 times the luminosity that STAR sampled during
2162 transversely polarized $p+Au$ collisions in Run-15.

2163 The reduction in cyo-weeks from 28 to 24 is projected to have a significant impact on the
2164 sampled luminosity, reducing the statistics quoted above by about a factor of 1.3.

2165 3.1 Spin Physics with Polarized pp and $p+A$ Collisions at 200 GeV

2166 Run-24 will enable STAR to probe the physics questions that can be assessed in the trans-
2167 versely polarized $p+p$ and $p+A$ collisions, including those described in highlights section 1.2
2168 and recent STAR publications [175, 176], but with a far more capable detector and much
2169 larger datasets than were available during Run-15. With the overlapping kinematic cover-
2170 age for both $p+p$ and $p+A$ data, this program is critical to set the stage for related future
2171 measurements at the EIC. Here we give brief descriptions of several of the opportunities
2172 presented by Run-24.

2173 Forward Transverse Spin Asymmetries

2174
2175 The experimental study of spin phenomena in nuclear and particle physics has a long
2176 history of producing important, and often surprising, results. Attempts to understand such
2177 data have pushed the field forward, forcing the development of both new theoretical frame-
2178 works and new experimental techniques. Recent detector upgrades at STAR, at mid- and
2179 forward-rapidity, coupled with the versatility of RHIC, will allow us to gain new insights
2180 into long-standing puzzles, and to probe more deeply the complexities of emergent behavior
2181 in QCD.

2182 Results from PHENIX and STAR have shown that large transverse single-spin asym-
2183 metries (TSSA) for inclusive hadron production, first seen in $p+p$ collisions at fixed-target
2184 energies and modest p_T , extend to the highest RHIC center-of-mass energies, $\sqrt{s} = 510 \text{ GeV}$,
2185 and surprisingly large p_T . Figure 74 summarizes the world data for the inclusive neutral pion
2186 asymmetries A_N as a function of Feynman- x . The asymmetries are seen to be nearly inde-
2187 pendent of \sqrt{s} over the very wide range of roughly 19 to 500 GeV.

2188 To understand the observed TSSAs, one needs to go beyond the conventional leading-
2189 twist (twist-2) collinear parton picture for the hard-scattering processes. Two theoretical
2190 formalisms have been developed to try to explain these sizable asymmetries in the QCD
2191 framework: transverse-momentum-dependent (TMD) parton distribution and fragmentation
2192 functions, such as the Sivers and Collins functions; and transverse-momentum-integrated
2193 (collinear) quark-gluon-quark correlations, which are twist-3 distributions in the initial state

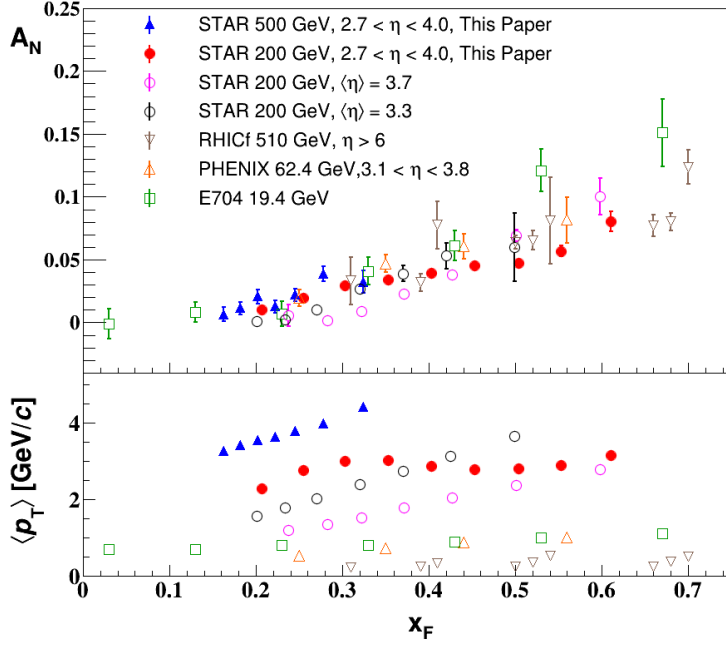


Figure 74: Transverse single-spin asymmetry A_N measurements for neutral pion in $p+p$ collisions at different center-of-mass energies as a function of Feynman- x [175].

2194 proton or in the fragmentation process. For many of the experimentally accessible spin
 2195 asymmetries, several of these functions can contribute, and need to be disentangled in order
 2196 to understand the experimental data in detail, in particular the observed p_T dependence.
 2197 These functions manifest their spin dependence either in the initial state—for example, the
 2198 Siverson distribution and its twist-3 analog, the Efremov-Teryaev-Qiu-Sterman (ETQS) func-
 2199 tion [177]—or in the final state via the fragmentation of polarized quarks, such as in the
 2200 Collins function and related twist-3 function $\hat{H}_{FU}(z, z_z)$.

2201 Incorporating the fragmentation term within the collinear twist-3 approach demonstrated
 2202 the ability of this formalism to describe the large values of A_N for π^0 production observed at
 2203 RHIC [178]. In this work, the relevant (non-pole) 3-parton collinear fragmentation function
 2204 $\hat{H}_{FU}(z, z_z)$ was fit to the RHIC data. The so-called soft-gluon pole term, involving the ETQS
 2205 function $T_{q,F}(x_1, x_2)$, was also included by fixing $T_{q,F}$ through its well-known relation to the
 2206 TMD Siverson function f_{1T}^\perp . The authors obtained a very good description of the data due to
 2207 the inclusion of the non-pole fragmentation function and based on this work they were able
 2208 to make predictions for π^+ and π^- production asymmetries A_N at the forward rapidities
 2209 covered by the STAR upgrades, $2.5 < \eta < 4$. The results are shown in Fig. 75 for $\sqrt{s} = 200$
 2210 and 500 GeV for two rapidity ranges, $2 < \eta < 3$ and $3 < \eta < 4$.

2211 STAR recently published in a pair of papers discussing forward transverse spin asymme-
 2212 tries in $p+p$, $p+Al$, and $p+Au$ collisions measured with the Forward Meson Spectrometer
 2213 (FMS). One paper focuses on the dynamics that underlie the large asymmetries that have
 2214 been seen to date [175]. The data show that A_N for forward π^0 production in $p+p$ collisions

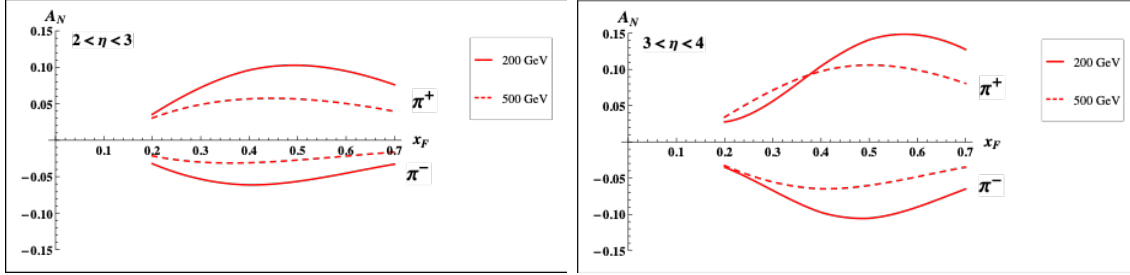


Figure 75: Predictions for A_N for π^+ and π^- production over the ranges $2 < \eta < 3$ (left) and $3 < \eta < 4$ (right) at $\sqrt{s} = 200$ GeV (solid lines) and 500 GeV (dashed lines).

2215 at 200 and 500 GeV is substantially larger when the π^0 is isolated than when it is accom-
 2216 panied by additional nearby photons. The same analysis also shows that A_N for inclusive
 2217 electromagnetic jets (EM-jets) in 200 and 500 GeV collisions is substantially larger than that
 2218 for EM-jets that contain three or more photons and that the Collins asymmetry for π^0 in
 2219 EM-jets is very small. The other paper focuses on the nuclear dependence of A_N for π^0 in
 2220 $\sqrt{s_{NN}} = 200$ GeV collisions [176]. It presents a detailed mapping of A_N as functions of x_F
 2221 and p_T for all three collision systems. It is shown that the observed nuclear dependence is
 2222 very weak. The same analysis shows that isolated *vs.* non-isolated π^0 behave similarly in
 2223 p +Al and p +Au collisions as they do in p + p collisions.

2224 These two papers provide a wealth of new data to inform the ongoing discussion regarding
 2225 the origin of the large inclusive hadron transverse spin asymmetries that have been seen in
 2226 p + p collisions at forward rapidity over a very broad range of collision energies. Nonetheless,
 2227 the STAR Forward Upgrade will be a game changer for such investigations. It will enable
 2228 measurements of A_N for $h^{+/-}$, in addition to π^0 . It will enable isolation criteria to be applied
 2229 to the $h^{+/-}$ and π^0 that account for nearby charged, as well as neutral, fragments. It will
 2230 enable full jet asymmetry and Collins effect measurements, again for $h^{+/-}$ in addition to
 2231 π^0 , rather than just EM-jet measurements. It will permit all of these measurements to be
 2232 performed at both 510 GeV (measured during Run 22), and at 200 GeV (to be measured in
 2233 Run 24).

2234 In addition, all of these observables can be tagged by requiring rapidity gaps to iden-
 2235 tify the diffractive component of the observed transverse spin asymmetries. For p + p there
 2236 will be considerable overlap between the kinematics at the two energies, but the 510 GeV
 2237 measurements will access higher p_T , while the 200 GeV measurements will access higher x_F .
 2238 Moreover, at 200 GeV we will also perform the full suite of measurements in p +Au to identify
 2239 any nuclear effects. Furthermore, it is important to stress that the 200 GeV running with
 2240 the Forward Upgrade will give the unique opportunity for jet reconstruction studies at the
 2241 exact same rapidity that is critical for the future EIC. The data will provide an extraor-
 2242 dinary possibility to exercise new reconstruction techniques incorporating AI/ML methods
 2243 and train the next generation of scientists.

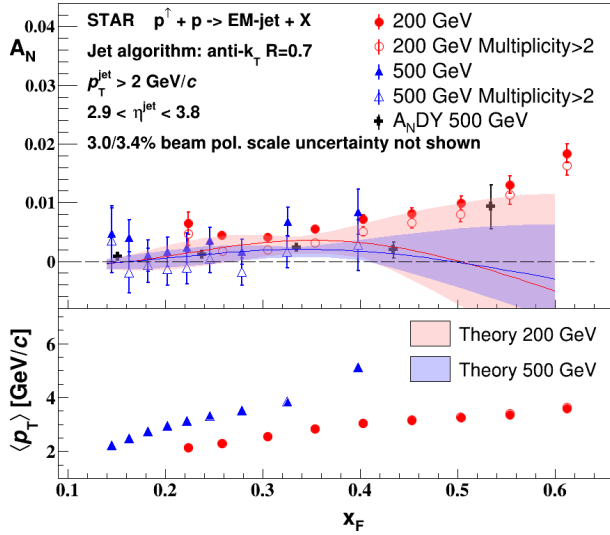


Figure 76: Recent STAR results on inclusive electromagnetic jets TSSA in pp collisions at both 200 and 500 GeV [175]. The results that require more than two photons observed inside a jet are shown as open symbols. Theory curves [182] for TSSA of full jets at rapidity $\langle y \rangle = 3.25$ for 200 GeV (red) and $\langle y \rangle = 3.57$ for 500 GeV (blue) are also shown. The average p_T of the jet for each x_F bin is shown in the lower panel.

Sivers and Efremov-Teryaev-Qiu-Sterman Functions

There is great theoretical interest in testing the relation between the ETQS correlation functions and the Sivers function. As discussed above, both the Sivers and the ETQS functions encapsulate partonic spin correlations within the proton, but they are formally defined in different frameworks. While the Sivers function is a TMD quantity that depends explicitly on spin-dependent transverse partonic motion k_T , the ETQS function is a twist-3 collinear distribution, in which SSAs are generated through soft collinear gluon radiation.

Measurements of forward jet production from the ANDY collaboration [179] indicated rather small asymmetries. This was argued to be consistent with the idea that the twist-3 parton correlation functions for up and down valence quarks should cancel, because their behavior reflects the Sivers functions extracted from fits to the SIDIS data that demonstrate opposite sign, but equal magnitude, up and down quark Sivers functions. STAR results on charge-tagged dijets at mid-rapidity [180] (see Fig. 79) support this interpretation, with the caveat that the measured observable (a spin-dependent $\langle k_T \rangle$) is defined in the TMD, and not the twist-3, framework. Moreover, recently published STAR results for forward inclusive electromagnetic jets [175] also show small TSSA as seen in Fig. 76. The results have been analyzed with the generalized parton model approach [181], and when incorporated in the reweighing procedure of the quark Sivers functions extracted from SIDIS data they significantly improved its uncertainty at larger momentum fraction x (see Fig. 77).

To better test quantitatively the relation between the two regimes, one can measure spin asymmetries for jets which are *intentionally* biased towards up or down quark jets via detection of a high- z charged hadron within the jet. Higher-twist calculations of jet asymmetries based on the Sivers function predict sizeable effects for these flavor-enhanced jets. With the suite of new forward detectors installed at STAR, full jet reconstruction, along with identification of a high- z hadron of known charge sign is possible at high pseudorapidity. Using

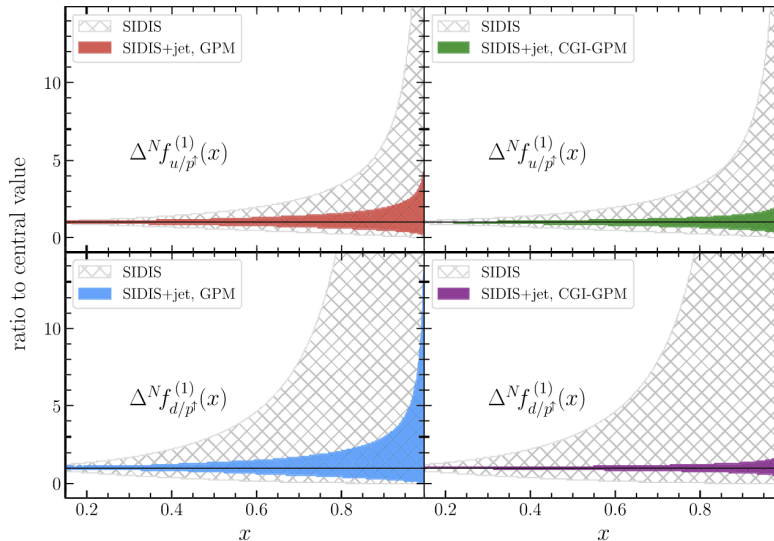


Figure 77: Comparison between the Siverson function first moments normalized to the corresponding central value from SIDIS data and their reweighted counterparts that incorporate new STAR results on electromagnetic jets [175] extracted in [181] in the generalized parton model (left panels) and color gauge invariant generalized parton model (right panels) framework. In both plots, results for u (upper panels) and d (lower panels) quarks are shown.

2270 realistic jet smearing in a forward calorimeter and tracking system, and requiring a charged
 2271 hadron with $z > 0.5$, the asymmetries can be separated and compared to the predictions for
 2272 the Siverson function based on current SIDIS data. The expected uncertainties, plotted at the
 2273 predicted values, can be seen in Fig. 78. Dilutions by underlying event and beam remnants
 2274 were taken into account. The simulations have assumed only an integrated luminosity of
 2275 100 pb^{-1} at $\sqrt{s} = 200 \text{ GeV}$, which is significantly lower than what is currently expected for
 2276 the Run-24 200 GeV polarized $p+p$ run.

2277 In a TMD framework, the Siverson effect manifests itself as a correlation (a triple product)
 2278 between the transverse momentum of a parton (\vec{k}_T) with momentum fraction x , and the
 2279 transverse spin (\vec{S}) of a polarized proton moving in the longitudinal (\vec{p}) direction. Thus,
 2280 for transversely polarized protons, the Siverson effect probes whether the k_T of the constituent
 2281 quarks is preferentially oriented in a direction perpendicular to both the proton momentum
 2282 and its spin. Momentum conservation then implies that the two jets in the final state will
 2283 not emerge back-to-back on average, but instead will ‘tilt’ in the direction of the summed
 2284 k_T of the initial state partons. Moreover, the (average) tilt of interest will reverse direction
 2285 under a ‘flip’ of the proton spin; a spin-dependent $\langle k_T \rangle$ can then be extracted by associating
 2286 the azimuthal opening angle of the jet pair with this tilt.

2287 STAR carried out an earlier measurement of this transverse single-spin asymmetry using
 2288 a dijet dataset with $\sim 1 \text{ pb}^{-1}$ of integrated luminosity [183], and found it to be consistent
 2289 with zero within 2σ . Figure 79 shows the first ever observation of the Siverson effect in dijet

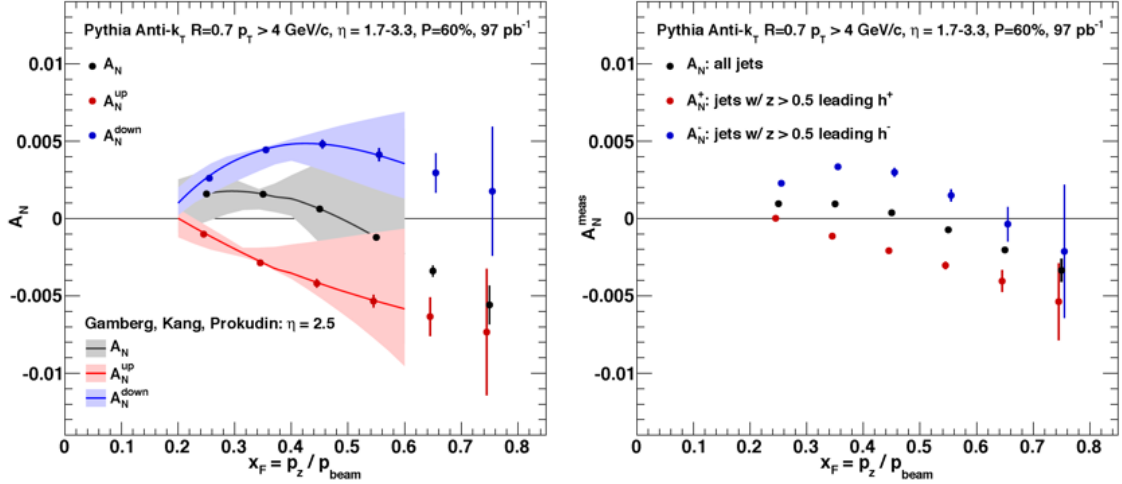


Figure 78: Left: up quark (red points), down quark (blue points) and all jet (black points) single spin asymmetries as a function of x_F as calculated by the ETQS based on the SIDIS Sivvers functions. Right: Expected experimental sensitivities for jet asymmetries tagging in addition a positive hadron with z above 0.5 (red points), a negative hadron with z above 0.5 (blue points) or all jets (black) as a function of x_F . Note: these figures are for 200 GeV center-of-mass energy proton collisions.

2290 production, which just entered GPC for publication. The jets were sorted according to their
 2291 net charge Q , calculated by summing the signed momentum of all particle tracks with $p >$
 2292 0.8 GeV, to minimize underlying event contributions, yielding jet samples with enhanced
 2293 contributions from u quarks (positive Q) and d quarks (negative Q), with a large set near
 2294 $Q = 0$ dominated by gluons. Simple kinematics allow for conversion from the spin-dependent
 2295 ‘tilt’ of the dijet pair to a value of k_T on an event-by-event basis; these are then sorted by the
 2296 Q of the jet and binned by the summed pseudorapidities of the outgoing jets, $\eta^{\text{total}} \equiv \eta_3 + \eta_4$.
 2297 Because the contributions of different partons (u , d , all else) to $\langle k_T \rangle$ vary with both Q and
 2298 also η^{total} , in a way that can be estimated robustly using simulation, the data can be inverted
 2299 to yield values of $\langle k_T \rangle$ for the individual partons, though with coarser binning in η^{total} .

2300 Such measurements are crucial to explore questions regarding factorization of the Sivvers
 2301 function in dijet hadroproduction [184–187]. Those results were derived from 200 GeV trans-
 2302 verse spin data that STAR recorded in Run-12 and Run-15 (total sampled luminosity ~ 75
 2303 pb^{-1} for the two years combined). Nonetheless, the uncertainties remain large, as can be seen
 2304 in Fig. 79. Run-24 data will reduce the uncertainties for $|\eta_3 + \eta_4| < 1$ by about a factor of
 2305 two. The increased acceptance from the iTPC will reduce the uncertainties at $|\eta_3 + \eta_4| \approx 2.5$
 2306 by a much larger factor, while the Forward Upgrade will enable the measurements to be
 2307 extended to even larger values of $|\eta_3 + \eta_4|$. When combined with the 510 GeV data from
 2308 Run-17 and Run-22, the results will provide a detailed mapping *vs. x* for comparison to
 2309 results for Sivvers functions extracted from SIDIS, Drell-Yan, and vector boson production.

2310 Transversity and Related Quantities

2311

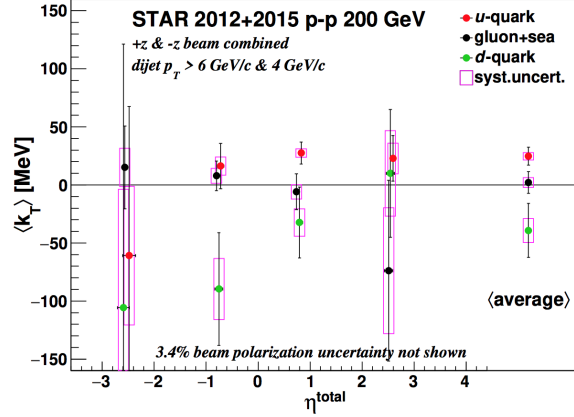


Figure 79: The $\langle k_T \rangle$ for individual partons, inverted using parton fractions from simulation and tagged $\langle k_T \rangle$ in data, is plotted as a function of $\eta_{\text{total}} \sim \log(x_1/x_2)$. The rightmost points represent the average of all the η_{total} bins. The systematic uncertainty in η_{total} is set to be non-zero to improve the visibility of the error bars.

2312 A complete picture of nucleon spin structure at leading twist must include contribu-
 2313 tions from the unpolarized and helicity distributions, as well as those involving transverse
 2314 polarization, such as the transversity distribution [188–190]. The transversity distribution
 2315 can be interpreted as the net transverse polarization of quarks within a transversely polar-
 2316 ized proton. The difference between the helicity and transversity distributions for quarks
 2317 and antiquarks provides a direct, x -dependent connection to nonzero orbital angular mo-
 2318 mentum components in the wave function of the proton [191]. Recently, the first lattice
 2319 QCD calculation of the transversity distribution has been performed [192]. In addition,
 2320 the measurement of transversity has received substantial interest as a means to access the
 2321 tensor charge of the nucleon, defined as the integral over the valence quark transversity:
 2322 $\delta q^a = \int_0^1 [\delta q^a(x) - \delta \bar{q}^a(x)] dx$ [189, 193]. Measuring the tensor charge is very important for
 2323 several reasons. First, it is an essential and fundamental quantity to our understanding of
 2324 the spin structure of the nucleon. Also, the tensor charge can be calculated on the lattice
 2325 with comparatively high precision, due to the valence nature of transversity, and hence is
 2326 one of the few quantities that allow us to compare experimental results on the spin structure
 2327 of the nucleon directly to *ab initio* QCD calculations. Finally, the tensor charge describes
 2328 the sensitivity of observables in low-energy hadronic reactions to beyond the standard model
 2329 physics processes with tensor couplings to hadrons. Examples are experiments with ultra-
 2330 cold neutrons and nuclei.

2331 Transversity is difficult to access due to its chiral-odd nature, requiring the coupling of
 2332 this distribution to another chiral-odd distribution. Semi-inclusive deep-inelastic scattering
 2333 (SIDIS) experiments have successfully probed transversity through two channels: asym-
 2334 metric distributions of single pions, convoluting the TMD transversity distribution with
 2335 the TMD Collins fragmentation function, and azimuthally asymmetric distributions of di-
 2336 hadrons, coupling transversity to the so-called “interference fragmentation function” (IFF)

2337 in the framework of collinear factorization. Yet in spite of a wealth of lepton-scattering
2338 data, the kinematic reach of existing SIDIS experiments limits the precision with which the
2339 proton's transversity can be extracted, as the range of Bjorken- x values that can be accessed
2340 does not extend above $x \sim 0.3$.

2341 In hadronic collisions, the k_T integrated quark transversity distribution may be accessed
2342 mainly via two channels. The first is the single spin asymmetry of the azimuthal distribution
2343 of hadrons in high energy jets [194]. In the jet+hadron channel, the collinear transversity
2344 distribution couples to the TMD Collins function [194, 195]. This makes $p+p$ collisions a
2345 more direct probe of the Collins fragmentation function than Collins asymmetries in SIDIS
2346 [194], where a convolution with the TMD transversity distribution enters. This also makes
2347 the Collins asymmetry in $p+p$ collisions an ideal tool to explore the fundamental QCD
2348 questions of TMD factorization, universality, and evolution. The second channel is the
2349 single spin asymmetry of pion pairs, where transversity couples to the collinear interference
2350 fragmentation function [196]. STAR mid-rapidity IFF data [197] have been included in the
2351 first extraction of transversity from SIDIS and proton-proton IFF asymmetries [198]. In
2352 addition, transverse spin transfer, D_{TT} , of Λ hyperons in $p+p$ collisions is also expected
2353 to be able to provide sensitivity for the strange quark transversity through the polarized
2354 fragmentation functions. The strange quark transversity is not constrained at all currently.
2355 The first D_{TT} measurement of Λ and $\bar{\Lambda}$ hyperons at $\sqrt{s} = 200$ GeV has been performed
2356 with the Run-12 $p+p$ dataset [199] and preliminary results based on Run-15 have been
2357 released [200]. Current results didn't indicate a sizable spin transfer yet. The iTPC upgrade
2358 will help to reach near-forward pseudo-rapidity $\eta < 1.5$ for the spin transfer measurements.

2359 The universality of TMD PDFs and fragmentation functions in $p+p$ collisions has been an
2360 open question. General arguments [184, 185] have shown that factorization can be violated
2361 in hadron-hadron collisions for TMD PDFs like the Sivers function, though very recent
2362 calculations indicate the violations might be quite small [186, 187]. In contrast, while there
2363 is no general proof that the Collins effect in $p+p$ collisions is universal to all orders, explicit
2364 calculations [194, 195, 201, 202] have shown that diagrams like those that violate factorization
2365 of the Sivers function make no contribution to the Collins effect at the one- or two-gluon
2366 exchange level, thereby preserving its universality at least to that level.

2367 Comparisons of the transversity distributions extracted from the Collins and IFF channels
2368 will allow STAR to study the size and nature of any factorization breaking effects for TMD
2369 observables in hadronic collisions. Likewise, comparisons with the transversity, Collins and
2370 IFF distributions extracted from SIDIS collisions will shed light on universality and constrain
2371 evolution effects. The measurement of evolution effects in TMD distributions is particularly
2372 important because, unlike the collinear case, TMD evolution contains a non-perturbative
2373 component that cannot be calculated directly.

2374 Data from 200 GeV $p+p$ collisions will play an essential role toward answering these
2375 questions. Figure 80 shows that 200 GeV $p+p$ collisions interpolate between the coverage
2376 that we will achieve with collected Run-22 data at high- x with the Forward Upgrade and at
2377 low- x with the STAR mid-rapidity detectors. They will also provide a significant overlapping
2378 region of x coverage, but at Q^2 values that differ by a factor of 6. This will provide valuable

2379 information about evolution effects, as well as cross-checks between the two measurements.
 2380 Furthermore, for most of the overlapping x region, 200 GeV $p+p$ collisions will also provide
 2381 the greatest statistical precision (see for example Fig. 81), thereby establishing the most
 2382 precise benchmark for future comparisons to ep data from the EIC. It is important to also
 2383 recognize that the hadron-in-jet measurements with the STAR Forward Upgrade will provide
 2384 very valuable experience detecting jets close to beam rapidity that will inform the planning
 2385 for future jet measurements in similar kinematics at the EIC.

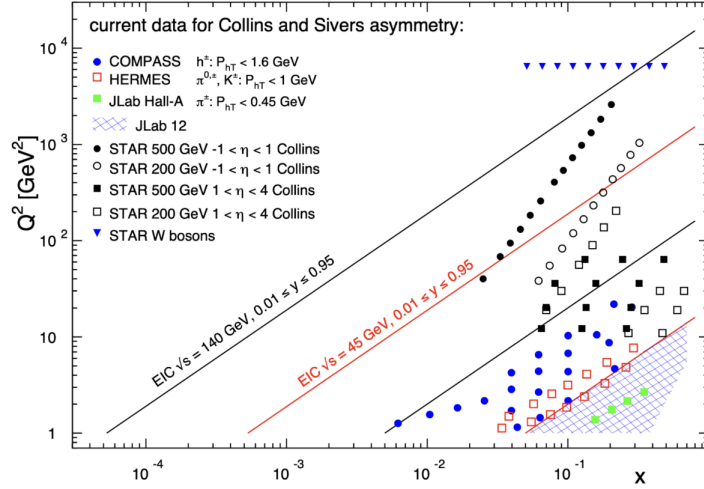


Figure 80: $x - Q^2$ coverage of RHIC measurements compared to existing Collins and Sivers effect measurements in SIDIS and the future coverage of the EIC.

2386 The high statistical precision of the Run-24 data will enable detailed multi-dimensional
 2387 binning for the Collins asymmetry results. This is particularly valuable because, as empha-
 2388 sized in [194, 195], hadron-in-jet measurements in $p+p$ collisions provide a direct probe of
 2389 the Collins fragmentation function since they combine it with the *collinear* transversity dis-

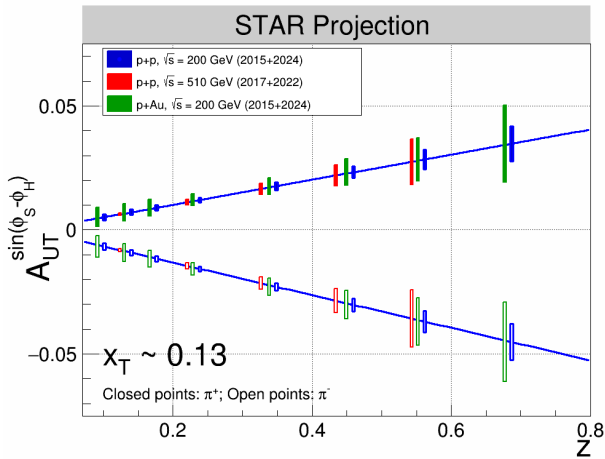


Figure 81: Projected statistical uncertainties for STAR Collins asymmetry measurements at $0 < \eta < 0.9$ in $p+p$ at $\sqrt{s} = 200$ and 510 GeV and p -Au at $\sqrt{s_{NN}} = 200$ GeV. The points have arbitrarily been drawn on the solid lines, which represent simple linear fits to the STAR preliminary 200 GeV $p+p$ Collins asymmetry measurements from 2015. (Note that only one bin is shown spanning $0.1 < z < 0.2$ for 510 GeV $p+p$ whereas three bins are shown covering the same z range for the 200 GeV measurements).

2390 tribution. In general, the observed asymmetries are functions of jet (p_T, η) , hadron (z, j_T) ,
 2391 and Q^2 . However, the physics interpretations associated with these variables separate, with
 2392 p_T and η primarily coupling to the incident quark x and the polarization transfer in the
 2393 hard scattering, while z and j_T characterize the fragmentation kinematics. Thus, A_{UT} vs.
 2394 p_T provides information about the transversity distribution, while the (z, j_T) dependence
 2395 provides a detailed look at the Collins fragmentation function. Recently finalized results
 2396 based on Run-12 and Run-15 datasets, discussed in Sec 1.2, finds the maximum value of
 2397 A_{UT} shift to higher j_T as p_T increases (see Fig. 34) which is not seen in the current theory
 2398 evaluations [203]. The statistical uncertainties in Fig. 34 will be reduced by a factor of about
 2399 2.5 when Run-12, Run-15 and Run-24 data are combined together.

2400 The Run-15 Collins analysis has also, for the first time, measured the Collins effect
 2401 for charged kaons and protons/anti-protons in $p+p$ collisions, as shown in Fig. 35. The
 2402 asymmetries for K^+ , which like π^+ have a contribution from favored fragmentation of u
 2403 quarks, are similar in magnitude to the π^+ asymmetries, while those for K^- , which can only
 2404 come from unfavored fragmentation, are consistent with zero at the 1-sigma level. These
 2405 trends are similar to those found in SIDIS by HERMES [204] and COMPASS [205], and
 2406 provide additional insight into the Collins fragmentation function. This same analysis with
 2407 Run-24 data will yield statistical uncertainties about a factor of 3 smaller than those in
 2408 Fig. 35. This is a much greater improvement than would be expected from the increase
 2409 in sampled luminosity thanks to the improved dE/dx resolution provided by the iTPC. In
 2410 addition, the iTPC will enable the measurements in Figs. 34 and 35 to be extended to an
 2411 additional higher η bin ($0.9 < \eta < 1.3$).

2412 RHIC has the unique opportunity to extend the Collins effect measurements to nuclei.
 2413 This will provide an alternative look at the universality of the Collins effect in hadron-
 2414 production by dramatically increasing the color flow options of the sort that have been
 2415 predicted to break factorization for TMD PDFs like the Sivers effect [184, 185]. This will
 2416 also explore the spin dependence of the hadronization process in cold nuclear matter. STAR
 2417 collected a proof-of-principle dataset during the 2015 $p+Au$ run that is currently under
 2418 analysis. Those data will provide a first estimate of medium-induced effects. However, the
 2419 small nuclear effects seen by STAR for forward inclusive $\pi^0 A_N$ [176] indicate that greater
 2420 precision will likely be needed. Figure 81 shows the projected Run-15 and Run-24 statistical
 2421 uncertainties for the $p+Au$ Collins asymmetry measurement at $\sqrt{s_{NN}} = 200$ GeV, compared
 2422 to those for the $p+p$ at the same energy.

2423 Ultra-peripheral Collisions

2424 The formalism of generalized parton distributions (GPDs) provides a theoretical framework
 2425 which addresses some of the above questions [206–209]. Constraints on GPDs have mainly
 2426 been provided by exclusive reactions in DIS, e.g. deeply virtual Compton scattering. RHIC,
 2427 with its unique capability to collide transversely polarized protons at high energies, has
 2428 the opportunity to measure A_N for exclusive J/ψ production in ultra-peripheral collisions
 2429 (UPCs) [210]. In such a UPC process, a photon emitted by the opposing beam particle (p or
 2430 A) collides with the polarized proton. The measurement is at a fixed $Q^2 \sim M_{J/\psi}^2 \approx 10$ GeV²

2431 and $10^{-4} < x < 10^{-1}$. A nonzero asymmetry would be the first signature of a nonzero GPD
 2432 E_g for gluons, which is sensitive to spin-orbit correlations and is intimately connected with
 2433 the orbital angular momentum carried by partons in the nucleon and thus with the proton
 2434 spin puzzle.

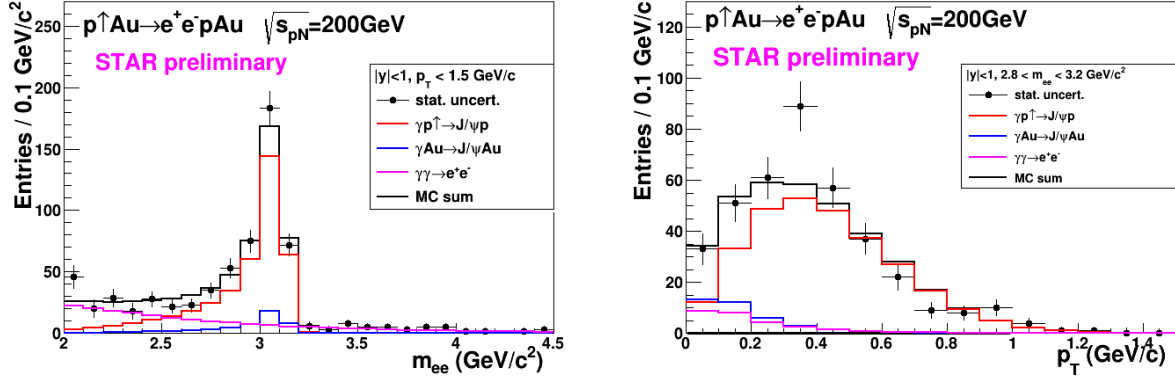


Figure 82: Mass distribution of selected e^+e^- pairs (left), and p_T distribution of the J/ψ mass peak (right). The colored histograms are the indicated processes modelled by STARlight and the sum fit to the data.

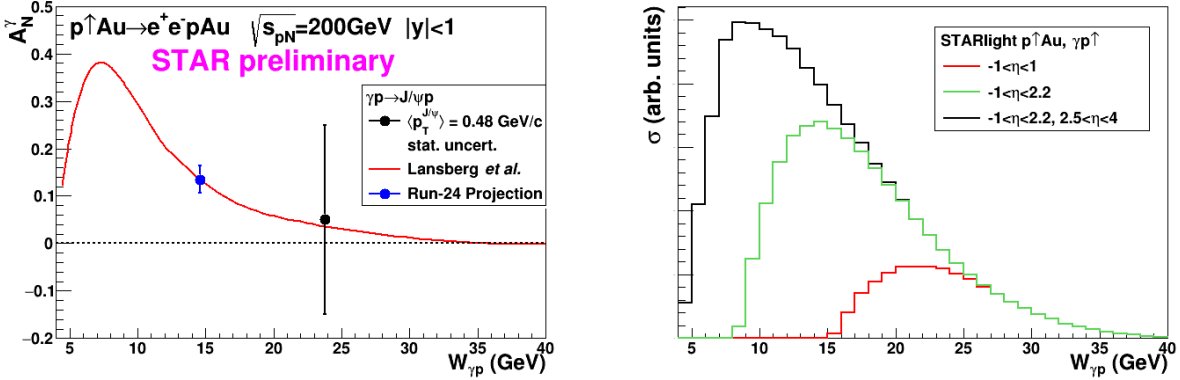


Figure 83: Left: The measured J/ψ transverse asymmetry A_N^γ and a prediction based on a parameterization of E_g . Right: The accepted cross section for $\gamma+p^\dagger \rightarrow J/\psi$ for various detector pseudorapidity η ranges; the black curve shows the result for the full STAR detector with the Forward Upgrade and the iTPC.

2435 The Run-15 p^\dagger +Au data allowed a proof-of-principle of such a measurement. A trigger
 2436 requiring back-to-back energy deposits in the Barrel Electromagnetic Calorimeter selected
 2437 J/ψ candidates. The e^+e^- mass distribution after selection cuts is shown in the left of
 2438 Fig. 82, and the pair p_T distribution of the J/ψ mass peak is shown on the right of that
 2439 figure. The data are well described by the STARlight model [211] (colored histograms in
 2440 the figure), including the dominant $\gamma+p^\dagger \rightarrow J/\psi$ signal process and the $\gamma+Au \rightarrow J/\psi$ and

2441 $\gamma+\gamma\rightarrow e^+e^-$ background processes. The left of Fig. 83 shows the STAR preliminary mea-
 2442 surement (solid circle marker) of the transverse asymmetry A_N^γ for the J/ψ signal, which has
 2443 a mean photon-proton center-of-mass energy $W_{\gamma p} \approx 24$ GeV. The result is consistent with
 2444 zero. Also shown is a prediction based on a parameterization of E_g [212]; the present data
 2445 provide no discrimination of this prediction.

2446 This measurement can be greatly improved with a high statistics transversely polarized
 2447 $p^\dagger+\text{Au}$ Run-24. The integrated luminosity for the Run-15 measurement was 140 nb^{-1} ; the
 2448 Run-24 will provide about 1.2 pb^{-1} , allowing a sizeable reduction of statistical uncertainty in
 2449 the same $W_{\gamma p}$ range. However, the Forward Upgrade and iTPC will also provide a significant
 2450 extension of the $W_{\gamma p}$ range of the measurement. The right panel of Fig. 83 shows the accepted
 2451 cross section for $\gamma+p^\dagger \rightarrow J/\psi$ for various detector pseudorapidity ranges. With the full
 2452 detector, the sensitive cross section is a factor of five times the central barrel alone and the
 2453 expected asymmetry is substantially larger. The projected statistical uncertainty on A_N^γ as
 2454 shown in the left of Fig. 83 (blue square marker) offering a powerful test of a non-vanishing
 2455 E_g . Also, the accepted region has a lower mean $W_{\gamma p} \approx 14$ GeV. Predictions based on E_g
 2456 parameterizations such as shown in the figure have a larger asymmetry at lower $W_{\gamma p}$, with
 2457 increased possibility of a nonzero result. Alternatively, the increased statistics will allow a
 2458 measurement of A_N^γ in bins of $W_{\gamma p}$.

2459 The UPC cross section scales with Z^2 of the the nucleus emitting the photon; for protons
 2460 this is $1/79^2$ relative to Au nuclei, which makes analogous measurements in $p+p$ collisions
 2461 extremely luminosity-hungry. Therefore, the $p+\text{Au}$ run is important for this measurement.

2462 In addition to the J/ψ measurements, the exclusive dijet studies, described in Sec. 2.2,
 2463 can be also sensitive to Generalized Parton Distributions.

2464 3.2 Physics Opportunities with Unpolarized proton-Nucleus Colli- 2465 sions

2466 Our quest to understand QCD processes in Cold Nuclear Matter (CNM) centers on the
 2467 following fundamental questions:

- 2468 • Can we experimentally find evidence of a novel universal regime of non-linear QCD
 2469 dynamics in nuclei?
- 2470 • What is the role of saturated strong gluon fields, and what are the degrees of freedom
 2471 in this high gluon density regime?
- 2472 • What is the fundamental quark-gluon structure of light and heavy nuclei?
- 2473 • Can a nucleus, serving as a color filter, provide novel insight into the propagation,
 2474 attenuation and hadronization of colored quarks and gluons?

2475 Various aspects of these questions have been addressed by numerous experiments and
 2476 facilities around the world, most of them at significantly lower center-of-mass energies and
 2477 kinematic reach than RHIC. Deep inelastic scattering on nuclei addresses some of these

2478 questions with results from, for instance, HERMES at DESY [213–215], CLAS at JLab [216],
2479 and in the future from the JLab 12 GeV. This program is complemented by hadron-nucleus
2480 reactions in fixed target $p+A$ at Fermilab (E772, E886, and E906) [217] and at the CERN-
2481 SPS.

2482 In the following we propose a measurement program unique to RHIC to constrain the
2483 initial state effects in strong interactions in the nuclear environment. We also highlight the
2484 complementarity to the LHC $p+Pb$ program and stress why RHIC data are essential and
2485 unique in the quest to further our understanding of nuclei. The uniqueness of the RHIC
2486 program is based on the flexibility of the RHIC accelerator to run collisions of different
2487 particle species at very different center-of-mass energies. This in combination with the
2488 enhanced STAR detector capabilities in Run-24 allows to disentangle nuclear effects in the
2489 initial and final state as well as leading twist shadowing from saturation effects in a kinematic
2490 regime where all these effects are predicted to be large. Most of the discussed measurements
2491 critically rely on the Forward Upgrade.

2492 The Initial State of Nuclear Collisions

2493 **Nuclear parton distribution functions:** A main emphasis of the Run-15 and later
2494 $p+A$ runs is to determine the initial conditions of the heavy ion nucleus before the collision
2495 to support the theoretical understanding of the A–A program both at RHIC and the LHC.
2496 In the following, the current status of nPDFs will be discussed, including where the unique
2497 contributions of RHIC lie, in comparison to the LHC and the future EIC.

2498 Our current understanding of nuclear parton distribution functions (nPDFs) is still very
2499 limited, in particular, when compared with the rather precise knowledge of PDFs for free
2500 protons collected over the past 30 years. Figure 84 shows an extraction of nPDFs from
2501 available data, along with estimates of uncertainties. All results are shown in terms of
2502 the nuclear modification ratios, i.e., scaled by the respective PDF of the free proton. The
2503 kinematic coverage of the data used in the EPPS21 fits [218] are shown in Fig. 85. Clearly,
2504 high precision data at small x and for various different values of Q^2 are needed to better
2505 constrain the magnitude of suppression in the x region where non-linear effects in the scale
2506 evolution are expected. In addition, such data are needed for several different nuclei, as
2507 the A-dependence of nPDFs cannot be predicted from first principles in pQCD and, again,
2508 currently relies on assumptions. The PHENIX midrapidity $\pi^0 R_{dAu}$ data [219], are the only
2509 data which can probe the gluon in the nucleus directly, but these data also suffer from
2510 unknown nuclear effects in the final state (see [220]). Therefore, it is critical to have high
2511 precision data only sensitive to nuclear modification in the initial state over a wide range in
2512 x and intermediate values of Q^2 (away from the saturation regime) to establish the nuclear
2513 modification of gluons in this kinematic range.

2514 It is important to realize that the measurements from RHIC are compelling and essential
2515 even when compared to what can be achieved in $p-Pb$ collisions at the LHC. Due to the
2516 higher center-of-mass system energy most of the LHC data have very high Q^2 , where the
2517 nuclear effects are already reduced significantly by evolution and are therefore very difficult

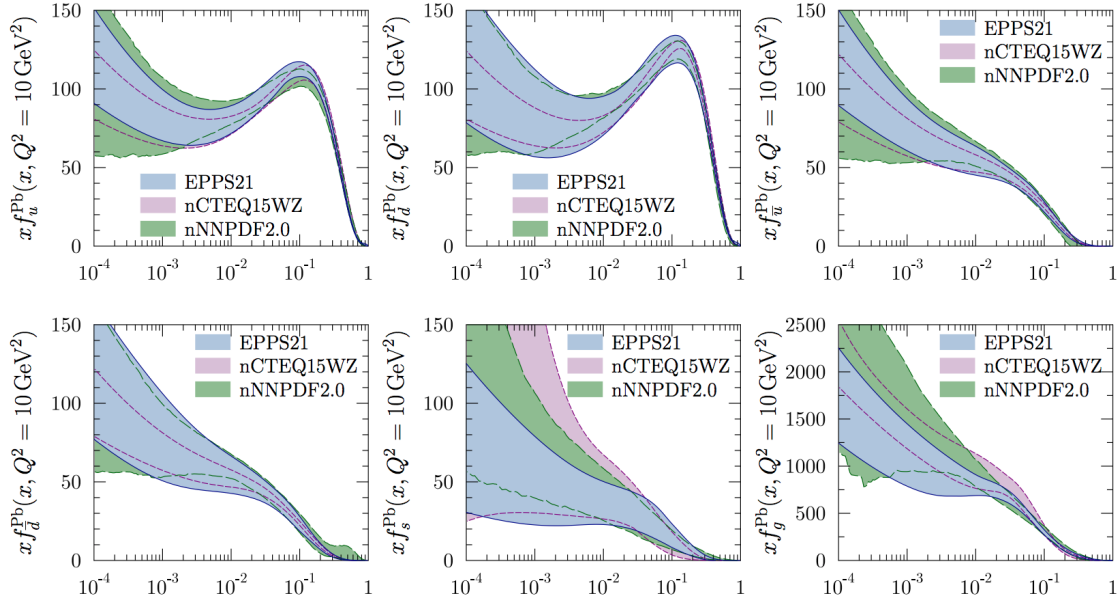


Figure 84: Summary of the most recent sets of nPDFs at 90% confidence-level [218].

2518 to constrain.

2519 RHIC has the *unique* capability to provide data in a kinematic regime (moderate Q^2 and
 2520 medium-to-low x) where the nuclear modification of the sea quark and the gluon is expected
 2521 to be sizable. In addition, and unlike the LHC, RHIC has the potential to vary the nucleus
 2522 in $p+A$ collisions and as such also constrain the A -dependence of nPDFs.

2523 Extraction of this information is less ambiguous if one uses processes in which strong
 2524 (QCD) final-state interactions can be neglected or reduced. Such golden channels would
 2525 include a measurement of R_{pA} for Drell-Yan production at forward pseudo-rapidities with
 2526 respect to the proton direction ($2.5 < \eta < 4$) to constrain the nuclear modifications of sea-
 2527 quarks. Moreover, the R_{pA} for direct photon production in the same kinematic regime will
 2528 help constrain the nuclear gluon distribution. Data for the first measurement of R_{pA}
 2529 for direct photon production have already been taken during the $p+Au$ and $p+Al$ Run-15, with
 2530 recorded luminosities by STAR of $L_{pAu} = 0.45 \text{ pb}^{-1}$ and $L_{pAl} = 1 \text{ pb}^{-1}$, respectively. Like
 2531 all other inclusive probes in $p+p$ and $p+A$ collisions, e.g., jets, no access to the exact parton
 2532 kinematics can be provided event-by-event but global QCD analyses easily account for that.
 2533 After the $p+Au$ Run-24, the statistical precision of the prompt photon data will be sufficient
 2534 to contribute to a stringent test of the universality of nuclear PDFs when combined with the
 2535 expected data from the EIC (see Figure 2.22 and 2.23 in Ref [221]). The Forward Upgrade
 2536 with its tracking at forward rapidities will also provide the possibility to measure R_{pA} for
 2537 positive and negatively charged hadrons. Approximately equal nucleon-nucleon luminosities
 2538 for $p+p$ and $p+Au$ are important for the optimization of R_{pA} measurements as they directly
 2539 compare the same observable—yields—in both collision systems.

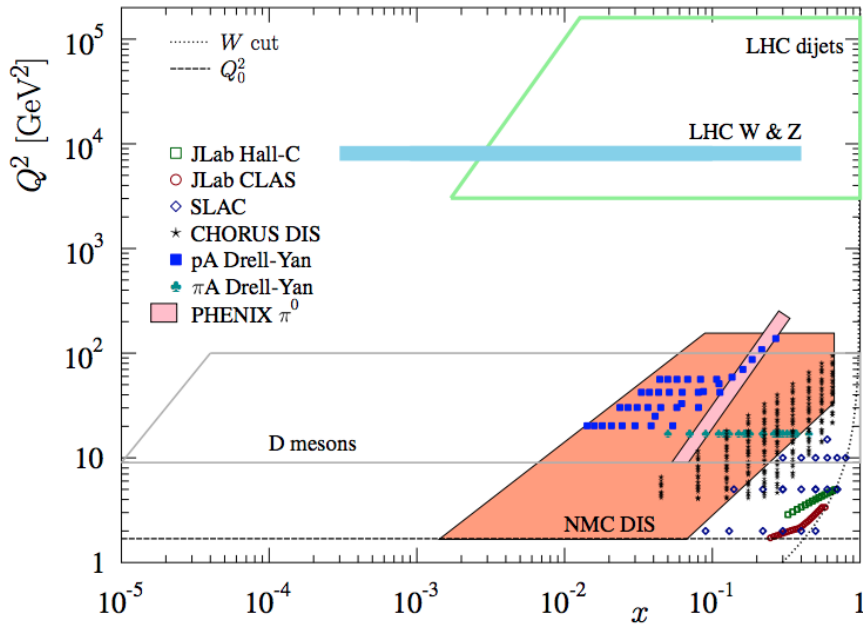


Figure 85: The kinematic x and Q^2 coverage of data used in the EPPS21 nPDF fits [218].

2540 Figure 86 shows the kinematic coverage in x - Q^2 of past, present, and future experiments
 2541 capable of constraining nuclear parton distribution functions. The shown experiments pro-
 2542 vide measurements that access the initial state parton kinematics on an event-by-event basis
 2543 (in a leading order approximation) while remaining insensitive to any nuclear effects in the
 2544 final state. Some of the LHC experiments cover the same x -range as DY at forward pseudo-
 2545 rapidities at RHIC but at a much higher scale Q^2 , where nuclear modifications are already
 2546 significantly reduced [222–224]. At intermediate Q^2 , DY at STAR will extend the low- x
 2547 reach by nearly one decade compared to EIC.

2548 The biggest challenge of a DY measurement is to suppress the overwhelming hadronic
 2549 background: the total DY cross-section is about 10^{-5} to 10^{-6} smaller than the corresponding
 2550 hadron production cross-sections. Therefore, the probability of misidentifying a hadron
 2551 track as a lepton has to be suppressed to the order of 0.1% while maintaining reasonable
 2552 electron detection efficiencies. To that end, we have studied the combined electron/hadron
 2553 discriminating power of the Forward Upgrade. It was found that by applying multivariate
 2554 analysis techniques to the features of EM/hadronic shower development and momentum
 2555 measurements we can achieve hadron rejection powers of 200 to 2000 for hadrons of 15 GeV
 2556 to 50 GeV with 80% electron detection efficiency.

2557 The potential impact of the DY R_{pA} data for the EPPS-19 sets of nPDFs was studied
 2558 through a re-weighting procedure [225]. We expect a significant impact on the uncertainties
 2559 of R_{pA} DY upon including the projected and properly randomized data. Clearly, the DY
 2560 data from RHIC will be instrumental in reducing present uncertainties in nuclear modifica-
 2561 tions of sea quarks. Again, these data will prove to be essential in testing the fundamental

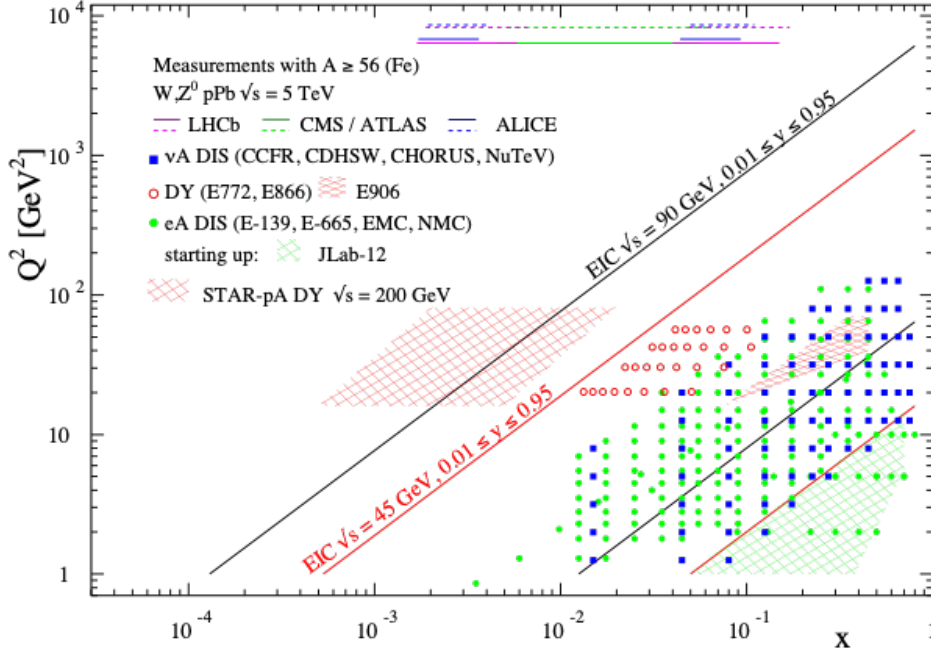


Figure 86: The kinematic coverage in $x - Q^2$ of past, present and future experiments constraining nPDFs with access to the exact parton kinematics event-by-event and no fragmentation in the final state.

2562 universality property of nPDFs in the future when EIC data become available.

2563 STAR’s unique detector capabilities provide data on J/Ψ -production in ultra-peripheral
 2564 collisions. This measurements can provide access to the spatial gluon distribution by mea-
 2565 suring the t -dependence of $d\sigma/dt$. To study the gluon distribution in the gold nucleus, events
 2566 need to be tagged where the photon is emitted from the proton ($\gamma + \text{Au} \rightarrow J/\psi$). However,
 2567 with the signal-to-background ratio in $p + \text{Au}$ collisions (see, e.g., Fig. 82), we expect much
 2568 better sensitivity to the gluon distributions in the Au+Au program. In addition to J/Ψ
 2569 photoproduction in UPC for exclusive reaction, photoproduction back-to-back jets is also
 2570 sensitive the PDFs (nPDFs in Au+Au UPC). This measurement has never been performed
 2571 at RHIC experiments, where the kinematic coverage can go to moderate to high- x . The
 2572 anti-shadowing region in nuclei, for example, is of great interest by comparing to this mea-
 2573 surement in proton. Furthermore, we can possibly extend the measurement from inclusive
 2574 photoproduction dijets to diffractive dijets in $p + p$ and $p + \text{Au}$ collisions, which will be sen-
 2575 sitive to the QCD factorisation breaking [151]. For details, see Sec. 2.2 for discussion in
 2576 UPCs.

2577 **Gluon Saturation:** Our understanding of the proton structure and of the nuclear
 2578 interactions at high energy would be advanced significantly with the definitive discovery
 2579 of the saturation regime [226–232]. Saturation physics would provide an infrared cutoff for
 2580 perturbative calculations, the saturation scale Q_s , which grows with the atomic number of the
 2581 nucleus A and with decreasing value of x . If Q_s is large it makes the strong coupling constant
 2582 small, $\alpha_s(Q_s^2) \ll 1$ allowing for perturbative QCD calculations to be under theoretical
 2583 control.

2584 It is well known that PDFs grow at small- x . If one imagines how such a high number of

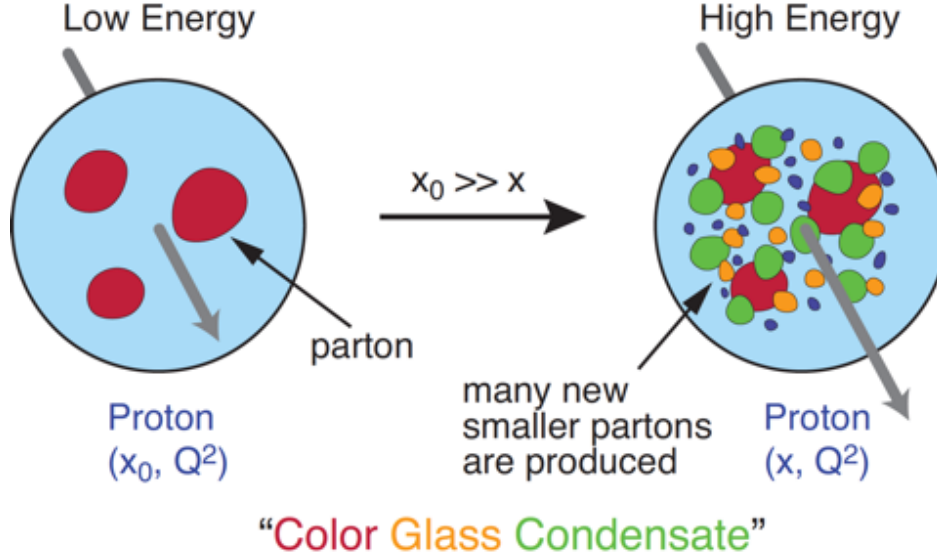


Figure 87: Proton wave function evolution towards small- x .

2585 small- x partons would fit in the (almost) unchanged proton radius, one arrives at the picture
 2586 presented in Fig. 87: the gluons and quarks are packed very tightly in the transverse plane.
 2587 The typical distance between the partons decreases as the number of partons increases, and
 2588 can get small at low- x (or for a large nucleus instead of the proton). One can define the
 2589 saturation scale as the inverse of this typical transverse inter-parton distance. Hence Q_s
 2590 indeed grows with A and decreasing x .

2591 The actual calculations in saturation physics start with the classical gluon fields (as gluons
 2592 dominate quarks at small- x) [233–239], which are then evolved using the nonlinear small- x
 2593 BK/JIMWLK evolution equations [240, 241, 241–249]. The saturation region can be well-
 2594 approximated by the following formula: $Q_s^2 \sim (A/x)^{1/3}$. Note again that at small enough
 2595 x the saturation scale provides an IR cutoff, justifying the use of perturbative calculations.
 2596 This is important beyond saturation physics, and may help us better understand small- x
 2597 evolution of the TMDs.

2598 While the evidence in favor of saturation physics has been gleaned from the data col-
 2599 lected at HERA, RHIC and the LHC, the case for saturation is not sealed and alternative
 2600 explanations of these data exist. The EIC is slated to provide more definitive evidence for
 2601 saturation physics [174]. To help the EIC complete the case for saturation, it is mandatory to
 2602 generate higher-precision measurements in p +Au collisions at RHIC. These higher-precision
 2603 measurements would significantly enhance the discovery potential of the EIC as they would
 2604 enable a stringent test of universality of the CGC. We stress again that a lot of theoretical
 2605 predictions and results in the earlier Sections of this document would greatly benefit from
 2606 saturation physics: the small- x evolution of TMDs in a longitudinally or transversely polar-
 2607 ized proton, or in an unpolarized proton, can all be derived in the saturation framework [250]
 2608 in a theoretically better-controlled way due to the presence of Q_s . Hence saturation physics
 2609 may help us understand both the quark and gluon helicity PDFs as well as the Sivers and

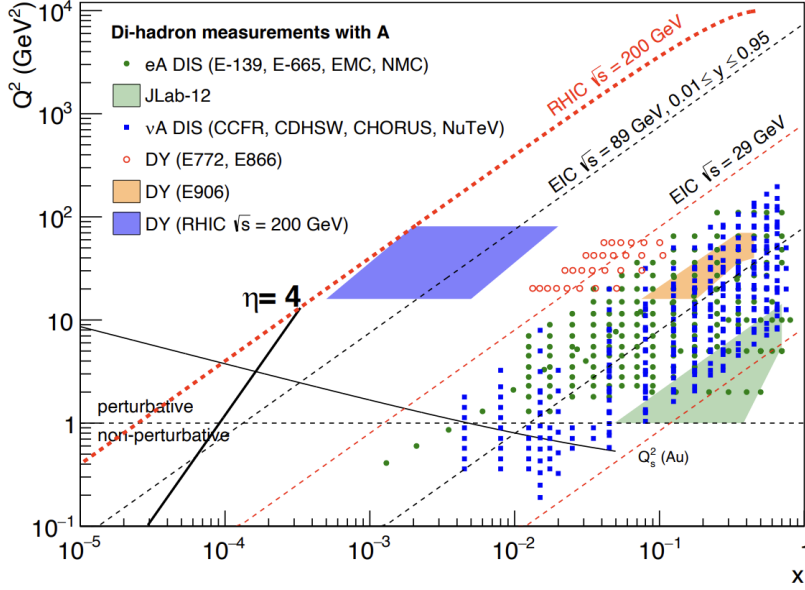


Figure 88: Kinematic coverage in the $x - Q^2$ plane for $p+A$ collisions at RHIC, along with previous $e+A$ measurements, the kinematic reach of an electron-ion collider, and estimates for the saturation scale Q_s in Au nucleus and the line illustrating the range in x and Q^2 covered with hadrons at rapidity $\eta = 4$.

2610 Boer-Mulders functions.

2611 The saturation momentum is predicted to grow approximately like a power of energy,
 2612 $Q_s^2 \sim E^{\lambda/2}$ with $\lambda \sim 0.2 - 0.3$, as phase space for small- x (quantum) evolution opens up.
 2613 The saturation scale is also expected to grow in proportion to the valence charge density at
 2614 the onset of small- x quantum evolution. Hence, the saturation scale of a large nucleus should
 2615 exceed that of a nucleon by a factor of $A^{1/3} \sim 5$ (on average over impact parameters). RHIC
 2616 is capable of running $p+A$ collisions for different nuclei to check this dependence on the mass
 2617 number. This avoids potential issues with dividing say p -Pb collisions in N_{part} classes [251].
 2618 Figure 88 shows the kinematic coverage in the $x - Q^2$ plane for $p+A$ collisions at RHIC, along
 2619 with previous $e+A$ measurements and the kinematic reach of an EIC. The saturation scale
 2620 for a Au nucleus is also shown. To access at RHIC a kinematic regime sensitive to saturation
 2621 with $Q^2 > 1$ GeV² requires measurements at forward rapidities. For these kinematics the
 2622 saturation scale is moderate, on the order of a few GeV², so measurements sensitive to the
 2623 saturation scale are by necessity limited to semi-hard processes.

2624 Until today the golden channel at RHIC to observe strong hints of saturation has been
 2625 the angular dependence of two-particle correlations, because it is an essential tool for testing
 2626 the underlying QCD dynamics [251]. In forward-forward correlations facing the $p(d)$ beam
 2627 direction one selects a large- x parton in the $p(d)$ interacting with a low- x parton in the
 2628 nucleus. For $x < 0.01$ the low- x parton will be back-scattered in the direction of the large-
 2629 x parton. Due to the abundance of gluons at small x , the backwards-scattered partons
 2630 are dominantly gluons, while the large- x partons from the $p(d)$ are dominantly quarks. The
 2631 measurements of di-hadron correlations by STAR and PHENIX [72, 252], have been compared
 2632 with theoretical expectations using the CGC framework based on a fixed saturation scale Q_s
 2633 and considering valence quarks in the deuteron scattering off low- x gluons in the nucleus with
 2634 impact parameter $b = 0$ [71, 253]. Alternative calculations [254] based on both initial and

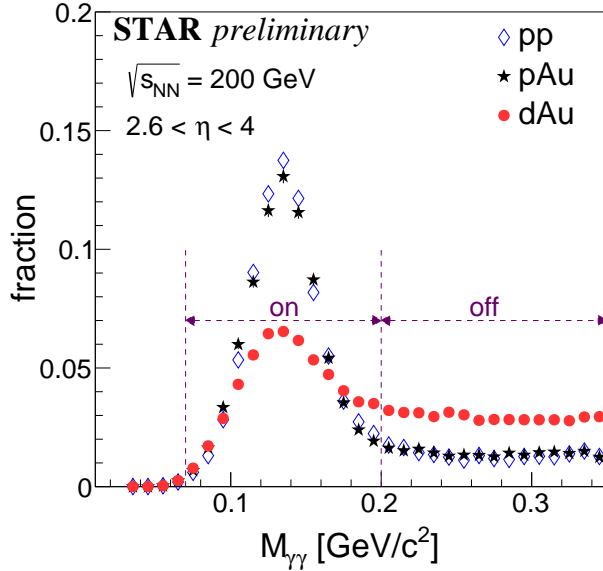


Figure 89: The invariant mass spectra for di-photon in $p+p$, $p+Au$ and $d+Au$. The on mass range is chosen as 0.07-0.2 GeV/c^2 , the off mass range is 0.2-0.35 GeV/c^2 .

2635 final state multiple scattering, which determine the strength of this transverse momentum
 2636 imbalance, in which the suppression of the cross-section in $d+Au$ collisions arises from cold
 2637 nuclear matter energy loss and coherent power corrections have also been very successful to
 2638 describe the data.

2639 The $p+A$ Run-15 at RHIC has provided unique opportunities to study this channel in
 2640 more detail at STAR. The high delivered integrated luminosities allow one to vary the trig-
 2641 ger and associated particle p_T from low to high values and thus crossing the saturation
 2642 boundary as shown in Fig. 88 and reinstate the correlations for central $p+A$ collisions for
 2643 forward-forward π^0 's. Studying di-hadron correlations in $p+A$ collisions instead of $d+Au$
 2644 collisions has a further advantage. In reference [73], the authors point out that the con-
 2645 tributions from double-parton interactions to the cross-sections for $dA \rightarrow \pi^0\pi^0 X$ are not
 2646 negligible. They find that such contributions become important at large forward rapidities,
 2647 and especially in the case of $d+A$ scattering. Figure 33 shows the relative area of back-to-
 2648 back di- π^0 correlations in $p+Al$ and $p+Au$ collisions relative to $p+p$ collisions. The results
 2649 show suppression with increasing A , and an enhanced suppression that scales as $A^{1/3}$. This
 2650 behavior is consistent with different calculations based on the CGC formalism and is a clear
 2651 hint of non-linear effects. A comparison between $p+p$ (Run-15), $p+Au$ (Run-15), and $d+Au$
 2652 (Run-16) collisions can help provide insight into the contributions from multiple parton scat-
 2653 tering [73]. Figure 89 shows the invariant mass spectra for final $p+p$ and $p+Au$ results and
 2654 the preliminary $d+Au$. It is clear from the comparison that there is significantly more back-
 2655 ground in the the $d+Au$ data than the $p+p$ and $p+Au$ data, which makes isolating the signal
 2656 correlation more difficult. The generated combinatoric correlation dominates in $d+Au$ colli-
 2657 sions, which makes it very challenging to identify the signal correlation. The forward di- π^0
 2658 correlation measurement favors for the cleaner $p+A$ collisions rather than $d+A$ collisions.
 2659 Run-24 will be able to measure di-hadron correlations taking advantage of the cleaner $p+Au$
 2660 collisions and the extended pseudorapidity reach of the Forward upgrade detectors.

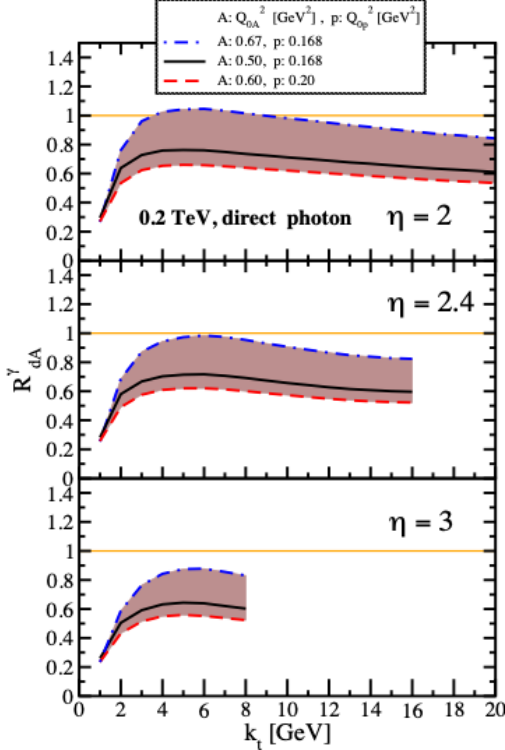


Figure 90: Nuclear modification factor for direct photon production in $p(d)+A$ collisions at various rapidities at RHIC $\sqrt{s} = 200$ GeV. The curves are the results obtained from Eq. (12) in Ref. [255] and the solution to rcBK equation using different initial saturation scales for a proton Q_{op} and a nucleus Q_{oA} . The band shows our theoretical uncertainties arising from allowing a variation of the initial saturation scale of the nucleus in a range consistent with previous studies of DIS structure functions as well as particle production in minimum-bias $p+p$, $p+A$ and $A+A$ collisions in the CGC formalism, see Ref. [255] for details.

2661 It is important to note that for the measurements to date in $p(d)+A$ collisions both initial
 2662 and final states interact strongly, leading to severe complications in the theoretical treatment
 2663 (see [256, 257], and references therein). As described in detail in the Section above in $p+A$
 2664 collisions, these complications can be ameliorated by removing the strong interaction from
 2665 the final state, by using photons and Drell-Yan electrons. The Run-15 $p+A$ run will for the
 2666 first time provide data on R_{pA} for direct photons and therefore allow one to test CGC based
 2667 predictions on this observable as depicted in Fig. 90 (taken from Ref. [255]). The higher
 2668 delivered integrated luminosity for the upcoming $p+Au$ Run-24 together with the Forward
 2669 Upgrade will enable one to study more luminosity hungry processes and/or complementary
 2670 probes to the di- π^0 correlations, i.e. di-hadron correlations for charged hadrons, photon-jet,
 2671 photon-hadron and di-jet correlations, which will allow a rigorous test of the calculation in
 2672 the CGC formalism. It is important to stress that the comparison of these correlation probes
 2673 in $p+p$ and $p+Au$ requires approximately equal nucleon-nucleon luminosities for these two
 2674 collision systems for optimal measurements. It is noted that these results are crucial for
 2675 the equivalent measurements at an EIC, which are planned at close to identical kinematics,
 2676 because only if non-linear effects are seen with different complementary probes, i.e., ep and
 2677 $p+A$ one can claim a discovery of saturation effects and their universality.

2678 We use direct photon plus jet (direct γ +jet) events as an example channel to indicate what
 2679 can be done in Run-24. These events are dominantly produced through the gluon Compton
 2680 scattering process, $g+q \rightarrow \gamma+q$, and are sensitive to the gluon densities of the nucleon and
 2681 nuclei in $p+p$ and $p+A$ collisions. Through measurements of the azimuthal correlations in
 2682 $p+A$ collisions for direct γ +jet production, one can study gluon saturation phenomena at

2683 small- x . Unlike di-jet production that is governed by both the Weizsäcker-Williams and
2684 dipole gluon densities, direct γ +jet production only accesses the dipole gluon density, which
2685 is better understood theoretically [255, 258]. On the other hand, direct γ +jet production
2686 is experimentally more challenging due to its small cross-section and large background con-
2687 tribution from di-jet events in which photons from fragmentation or hadron decay could be
2688 misidentified as direct photons. The feasibility to perform direct γ +jet measurements with
2689 the Forward Upgrade in unpolarized $p+p$ and p +Au collisions at $\sqrt{s_{NN}} = 200$ GeV has been
2690 studied. PYTHIA-8.189 [259] was used to produce direct γ +jet and di-jet events. In order
2691 to suppress the di-jet background, the leading photon and jet are required to be balanced in
2692 transverse momentum, $|\phi^\gamma - \phi^{jet}| > 2\pi/3$ and $0.5 < p_T^\gamma/p_T^{jet} < 2$. Both the photon and jet
2693 have to be in the forward acceptance $1.3 < \eta < 4.0$ with $p_T > 3.2$ GeV/ c in 200 GeV $p+p$
2694 collisions. The photon needs to be isolated from other particle activities by requiring the
2695 fraction of electromagnetic energy deposition in the cone of $\Delta R = 0.1$ around the photon
2696 is more than 95% of that in the cone of $\Delta R = 0.5$. Jets are reconstructed by an anti- k_T
2697 algorithm with $\Delta R = 0.5$. After applying these selection cuts, the signal-to-background
2698 ratio is around 3:1 [260]. The expected number of selected direct γ +jet events is around
2699 0.9M at $\sqrt{s_{NN}} = 200$ GeV in p +Au collisions for the proposed Run-24. We conclude that a
2700 measurement of direct photon-jet correlation from p +Au collisions is feasible, which is sen-
2701 sitive to the gluon density in $0.001 < x < 0.005$ in the Au nucleus where parton saturation
2702 is expected.

2703 **Saturation with Ultra-Peripheral Collisions.** There are other potential opportuni-
2704 ties with the upcoming p +Au and $p + p$ runs for studying the gluon saturation phenomena
2705 using the ultra-peripheral collisions (UPC). For example, one of the most powerful mea-
2706 surements proposed at the EIC for discovery of gluon saturation is to look at double ratio
2707 between heavy nucleus and proton in terms of diffractive processes, see details in Sec. 2.2.
2708 With the STAR Run-2024, the p +Au UPC (also applies to $p + p$ UPC) may provide two
2709 important measurements, e.g., exclusive and inclusive J/ψ production off the proton target.
2710 The same measurement will be performed in Au+Au UPC with Run-2023 and 2025, and
2711 together with different system comparison, the STAR data may provide strong evidences for
2712 saturation.

2713 The Final State

2714 **Fragmentation Functions:** In spite of the remarkable phenomenological successes of
2715 QCD, a quantitative understanding of the hadronization process is still one of the great
2716 challenges for the theory. Hadronization describes the transition of a quark or gluon into a
2717 final state hadron. It is a poorly understood process even in elementary collisions. RHIC's
2718 unique versatility will make it possible to study hadronization in vacuum and in the nuclear
2719 medium, and additionally with polarized beams (see Sect. 3.1 for the latter).

2720 It has long been recognized that the hadron distributions within jets produced in $p+p$
2721 collisions are closely related to the fragmentation functions that have typically been measured
2722 in e^+e^- collisions and SIDIS. The key feature of this type of observable is the possibility to

2723 determine the relevant momentum fraction z experimentally as the ratio of the hadron to the
 2724 jet transverse momentum. Recently [261] a quantitative relationship has been derived in a
 2725 form that enables measurements of identified hadrons in jets in $p+p$ collisions to be included
 2726 in fragmentation function fits on an equal footing with e^+e^- and SIDIS data. Furthermore,
 2727 hadrons in $p+p$ jets provide unique access to the gluon fragmentation function, which is
 2728 poorly determined in current fits [262], in part due to some tension found in the inclusive
 2729 high p_T pion yields measured by the PHENIX and ALICE collaborations. Here, the proposed
 2730 measurements can provide valuable new insight into the nature of this discrepancy.

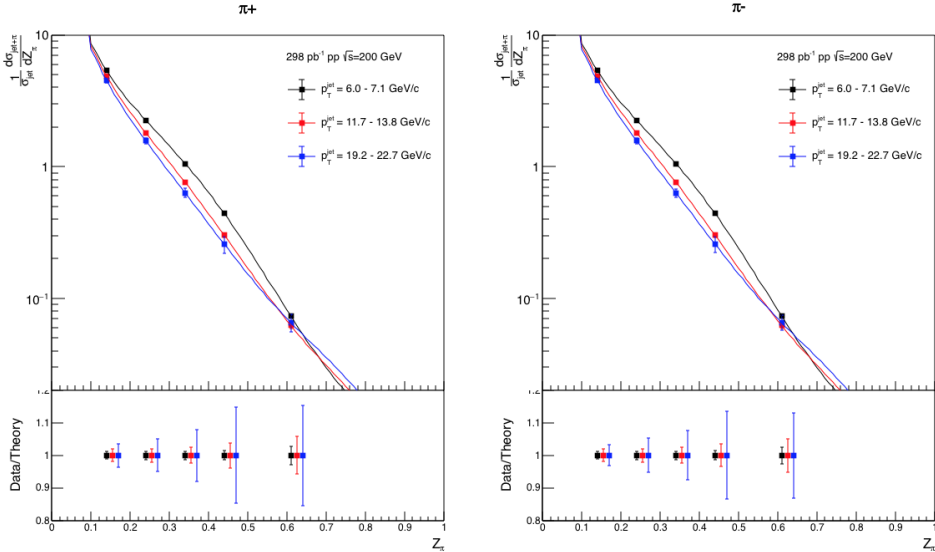


Figure 91: Anticipated precision for identified π^+ (left) and π^- (right) within jets at $|\eta| < 0.4$ in 200 GeV $p+p$ collisions for three representative jet p_T bins. The data points are plotted on theoretical predictions based on the DSSV14 pion fragmentation functions [261, 262]. Kaons and (anti)protons will also be measured, over the range from $z < 0.5$ at low jet p_T to $z < 0.2$ at high jet p_T , with uncertainties a factor of ~ 3 larger than those for pions.

2731 This development motivated STAR to initiate a program of identified particle fragmen-
 2732 tation function measurements using $p+p$ jet data at 200 and 500 GeV from Run-11, Run-12,
 2733 and Run-15. Figure 91 shows the precision that is anticipated for identified π^+ and π^- in 200
 2734 GeV $p+p$ collisions for three representative jet p_T bins after the existing data from Run-12
 2735 and Run-15 are combined with future 200 GeV $p+p$ data from Run-24. Identified kaon and
 2736 (anti)proton yields will also be obtained, with somewhat less precision, over a more limited
 2737 range of hadron z . Once the Run-17 data are fully analyzed, the uncertainties for 510 GeV
 2738 $p+p$ collisions will be comparable to that shown in Fig. 91 at high jet p_T , and a factor of \sim
 2739 2 larger than shown in Fig. 91 at low jet p_T . Identified hadron yields will also be measured
 2740 multi-dimensionally vs. j_T , z , and jet p_T , which will provide important input for unpolarized
 2741 TMD fits.

2742 Data from the HERMES experiment [213, 215, 263] have shown that production rates
 2743 of identified hadrons in semi-inclusive deep inelastic $e-A$ scattering differ from those in ep

2744 scattering. These differences cannot be explained by nuclear PDFs, as nuclear effects of
 2745 strong interactions in the initial state should cancel in this observable. Only the inclusion of
 2746 nuclear effects in the hadronization process allows theory to reproduce all of the dependencies
 2747 (z , x , and Q^2) of R_{eA} seen in SIDIS, as shown in Fig. 92.

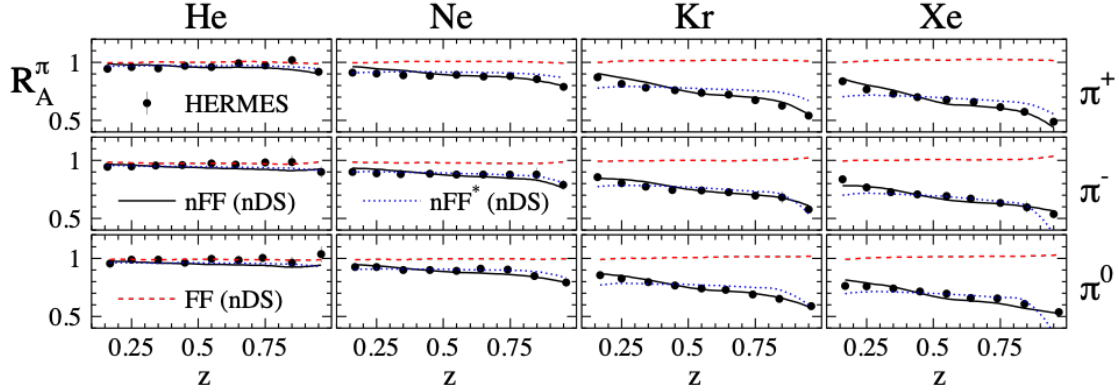


Figure 92: R_{eA} in SIDIS for different nuclei in bins of z as measured by HERMES [213, 215, 263]. The solid lines correspond to the results using effective nuclear FF [220] and the nDS medium modified parton densities [264]. The red dashed lines are estimates assuming the nDS medium modified PDFs but standard DSS vacuum FFs [265, 266] and indicate that nPDFs are insufficient to explain the data

2748 It is critical to see if these hadronization effects in cold nuclear matter persist at the
 2749 higher \sqrt{s} and Q^2 accessed at RHIC and EIC – both to probe the underlying mechanism,
 2750 which is not understood currently, and to explore its possible universality. The combination
 2751 of $p+p$ jet data from RHIC and future SIDIS data from EIC will also provide a much
 2752 clearer picture of modified gluon hadronization than will be possible with EIC data alone.
 2753 Using the Run-15 200 GeV $p+Au$ data, STAR will be able to make a first opportunistic
 2754 measurement of these hadron-jet fragmentation functions in nuclei, but the precision will
 2755 be limited. Additional $p+p$ and $p+Au$ data will be needed in Run-24 in order to provide a
 2756 sensitive test for universality, as shown in Fig. 93.

2757 QGP Droplet Substructure

2758 **Toroidal Vorticity:** In addition to cold QCD effects, a high-statistics measurement
 2759 of $p+Au$ collisions will be highly valuable to explore novel fluid configurations that have
 2760 recently been predicted [267]. In particular, the data is needed to discover vortex rings or
 2761 tubes at midrapidity, included by shear in the asymmetric initial state.

2762 It has been suggested [268] that $p+A$ collisions at RHIC form the "smallest QGP droplets."
 2763 This claim is often based on anisotropic yields, which resemble those from A+A collisions
 2764 that are attributed to hydrodynamic collective flow. Indeed, with well-chosen initial condi-
 2765 tions and tuned parameters, three-dimensional viscous hydro calculations can reproduce the
 2766 measured anisotropies from small, asymmetric collisions [269] at RHIC. However, a claim of

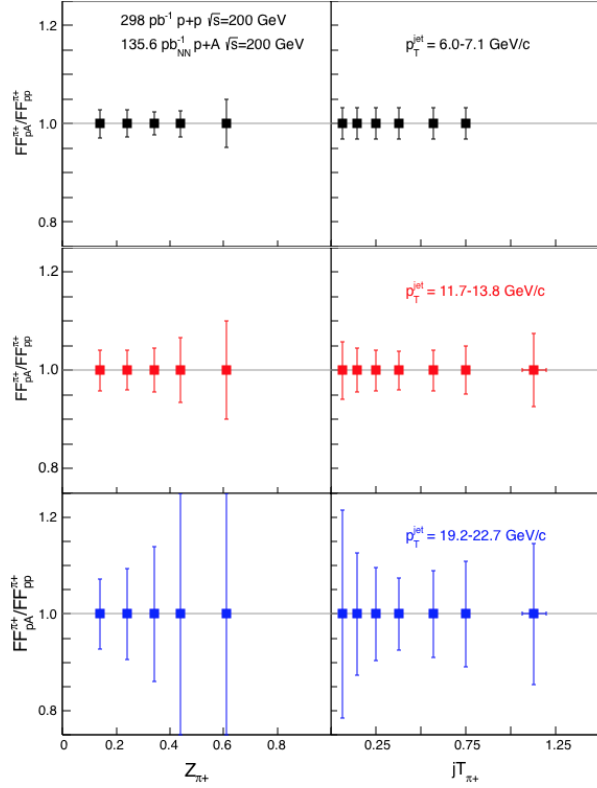


Figure 93: Anticipated precision for measurements of π^+ fragmentation functions in $p+A, p+p$ at $|\eta| < 0.4$ vs. z and j_T in Run-24 for three representative jet p_T bins. Uncertainties for π^- will be similar to those shown here for π^+ , while those for kaons and (anti)protons will be a factor of ~ 3 larger. Note that, to be species independent, the nucleon-nucleon equivalent luminosity is specified for $p+Au$.

2767 QGP formation in such small systems would be much more compelling if it were based on
 2768 more than one observable, especially since other, non-hydrodynamic mechanisms contribute
 2769 to v_n in these systems, e.g. [22].

2770 As Helmholtz observed more than 150 years ago [270], vortex rings are ubiquitous in
 2771 hydrodynamic systems subject to initial conditions characterized by a "push down the mid-
 2772 dle," such as a smoker blowing a ring. Clear observation of this novel phenomenon would
 2773 constitute important evidence that the smallest systems at RHIC truly do form a fluid sys-
 2774 tem.

2775 This signature probes aspects of particular and fundamental importance to the RHIC
 2776 program, as well. The vortex ring structure is sensitive to the degree and timescale of equi-
 2777 libration in these small systems, as well as the extreme shear fields in the initial state [271].
 2778 Fluctuations in the vortical fields probe hydrodynamic structures at the smallest possible
 2779 scales, as they arise directly from rotational derivatives in the "surface" of the flux tube.

2780 The experimental signature of toroidal vortex structure is the so-called "ring parame-
 2781 ter" [267]:

$$\overline{\mathcal{R}}_\Lambda^z \equiv \left\langle \frac{\vec{S}'_\Lambda \cdot (\hat{z} \times \vec{p}'_\Lambda)}{|\hat{z} \times \vec{p}'_\Lambda|} \right\rangle, \quad (7)$$

2782 where $+\hat{z}$ is the direction of the proton beam, and the average is taken over all particles and
 2783 events. This is the average polarization relative to the hyperon production plane. Rings will

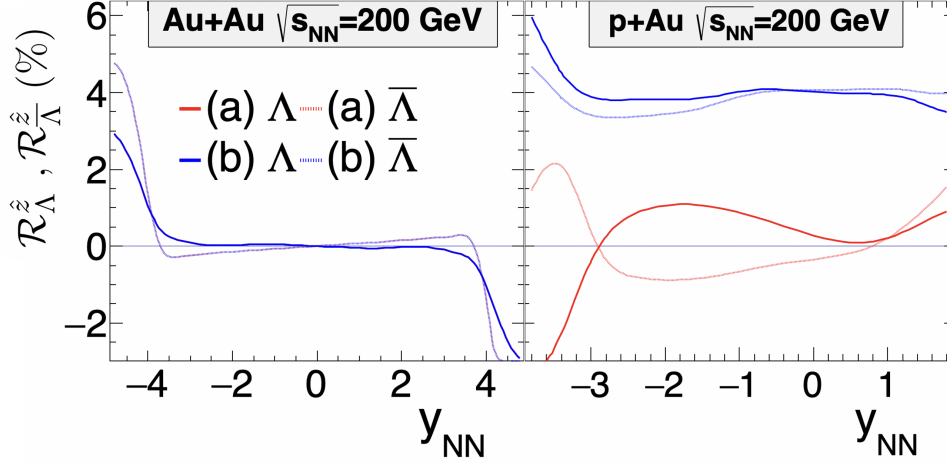


Figure 94: The "ring parameter" $\overline{\mathcal{R}}_{\Lambda}^z$ for $b = 0$ Au+Au and p+Au collisions at top RHIC energy. Blue (red) curves correspond to a scenario in which a toroidal vortex structure is (is not) generated by shear forces in the initial state. Solid (dashed) curves correspond to Λ ($\overline{\Lambda}$); note that baryon current is locally conserved in these collisions, so small differences between Λ and $\overline{\Lambda}$ are expected at finite baryon density. From [267].

2784 be most clear for central collisions, but the detailed centrality dependence of the effect is
 2785 currently under investigation [271]. We focus on 0-10% centrality.

2786 Figure 94 shows $\overline{\mathcal{R}}_{\Lambda}^z$ calculated [267] for completely central Au+Au and p+Au collisions at
 2787 $\sqrt{s_{NN}} = 200$ GeV. Calculations were done with MUSIC [272], a three-dimensional relativistic
 2788 viscous hydrodynamics simulation that locally conserves baryon number, and calculation of
 2789 the thermal vorticity along the freezeout hypersurface.

2790 Initial condition (a) corresponds to the usual Bjorken "boost-invariant" flow profile used
 2791 in most A+A simulations, whereas condition (b) features strong shear fields generated in
 2792 the initial condition, leading to observable vortex toroids. Both initial conditions generate
 2793 identical $dN/d\eta$ distributions, but the latter is argued [267] to be more natural.

2794 The statistical requirement to discover these toroidal vortex structures may be estimated
 2795 by STAR's previous hyperon polarization measurements. The uncertainty on global polar-
 2796 ization measurements $\delta\overline{P}_{\Lambda} \propto N_{\Lambda}^{-1/2} \cdot R_{\text{EP}}^{-1}$, where N_{Λ} is the total number of hyperons in the
 2797 analysis, and R_{EP} is the event plane resolution [7]. Because there is no event plane involved
 2798 in the production plane polarization, on the other hand, the uncertainty on the ring observ-
 2799 able goes as $\delta\overline{\mathcal{R}}_{\Lambda}^z \propto N_{\Lambda}^{-1/2}$. For the same-magnitude signal, then, $\overline{\mathcal{R}}_{\Lambda}^z$ enjoys an effective R_{EP}^{-2}
 2800 "statistical advantage" over \overline{P}_{Λ} . Since STAR measured [102] $\overline{P}_{\Lambda} \approx 1\%$ at $\sqrt{s_{NN}} = 11$ GeV
 2801 with 3.5σ significance, with the same number of hyperons in the analysis, we should be able
 2802 to measure $\overline{\mathcal{R}}_{\Lambda}^z \sim 1\%$ with 7σ significance. The 11-GeV analysis involved 6M Λ s, and we
 2803 estimate 0.02 Λ s per central (0 – 10%) p+Au collision at $\sqrt{s_{NN}} = 200$ GeV. Therefore, the
 2804 7σ measurement will require $6M/0.02 = 300M$ central p+Au collisions.

2805 Also crucial to this measurement is that data must be collected with both polarities
 2806 of STAR's magnetic field. This is because of large and highly nontrivial decay-topology-

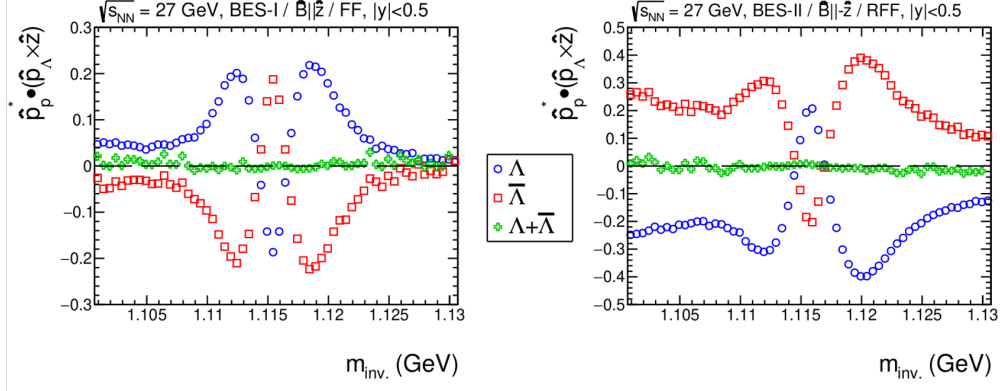


Figure 95: Production-plane polarization (modulo an overall scaling by $\frac{8\pi}{\alpha_\Lambda}$) for Λ (blue) and $\bar{\Lambda}$ (red) candidates, as a function of invariant mass. The data comes from STAR measurements of Au+Au collisions at $\sqrt{s_{NN}}$ in the BES-I (left) and BES-II (right) campaigns. STAR's solenoidal magnetic field was directed to the West and East, respectively, for these two datasets. For the BES-I data, hyperon candidates were identified with "standard" topological cuts, whereas the candidates shown in BES-II were identified using the new KFPARTICLE package.

2807 dependent detector effects, which will give a "false" production plane polarization signal.
 2808 The magnitude of the artifact is an order of magnitude larger than the physical signal of
 2809 interest, and it is highly sensitive to momentum, PID, and topological cuts. We could not
 2810 feel confident applying such large and complex "correction factors" based solely on detector
 2811 simulations, if we claim a completely novel signature with far-reaching physical implications.
 2812 Fortunately, the sign of this artifact flips with the magnetic field polarity.

2813 Figure 95 illustrates these points. Au+Au collisions at $\sqrt{s_{NN}} = 27$ GeV were recorded
 2814 by STAR using opposite polarities of the magnetic field. For Λ s, the quantity $\hat{p}_p \cdot (\hat{p}_\Lambda \times \hat{z})$,
 2815 where \vec{p}_p is the daughter proton momentum, is proportional to $\bar{\mathcal{R}}_\Lambda^z$. For $\bar{\Lambda}$ s, the quantity
 2816 $\hat{p}_{\bar{p}} \cdot (\hat{p}_{\bar{\Lambda}} \times \hat{z})$, where $\vec{p}_{\bar{p}}$ is the daughter proton momentum, is proportional to $-\bar{\mathcal{R}}_{\bar{\Lambda}}^z$.

2817 A rapidity cut symmetric about midrapidity ($|y| < 0.5$) was used; for a symmetric system,
 2818 the physical production plane polarization vanishes by symmetry— any nonvanishing value
 2819 results purely from topologically-sensitive efficiency effects.

2820 Consider first the Λ curve from BES-I, the blue points in the left panel. Clearly, the effect
 2821 has a nontrivial dependence on invariant mass; note even the asymmetry about $m_{inv} = m_\Lambda$.
 2822 Equally clearly, it is large, corresponding to values $\bar{\mathcal{R}}_\Lambda^z = \frac{8}{\pi\alpha_\Lambda} \hat{p}_p \cdot (\hat{p}_\Lambda \times \hat{z}) \approx 50\%$, an order
 2823 of magnitude larger than the predicted value of physical effect of interest.

2824 In terms of topologically-sensitive efficiency effects, substituting $\Lambda \rightarrow \bar{\Lambda}$ is equivalent to
 2825 flipping the sign of the magnetic field. The red datapoints in the left panel are a perfect
 2826 mirror image to the blue points in that panel, as indicated by the vanishing green points,
 2827 which are the sum. Further note that naive interpretation of the data in the left panel would
 2828 suggest that the vortical ring values for the hyperons and antihyperons ($\bar{\mathcal{R}}_\Lambda^z$ and $\bar{\mathcal{R}}_{\bar{\Lambda}}^z$) would
 2829 be identical in magnitude and sign.

2830 The right panel shows the same colliding system, but measured during the BES-II cam-
 2831 paign with the opposite orientation of STAR's magnetic field. As expected from the above

2832 discussion, $\overline{\mathcal{R}}_{\Lambda}^z = -\overline{\mathcal{R}}_{\overline{\Lambda}}^z$. The shape and magnitude of the artifact is different from the BES-I
 2833 case, however, because a different method has been used to identify hyperon candidates.
 2834 This illustrates the cut-dependence of the artifact.

2835 In short, for reliable extraction of the ring vorticity measure, STAR must measure p+Au
 2836 collisions with both field orientations, in order to cancel the complex efficiency-driven arti-
 2837 facts. Finally, we point out that this sort of cancellation is not unique to this observable.
 2838 Indeed, there is an analogous effect for the global polarization, which precludes extracting
 2839 the *first*-order azimuthal dependence of \overline{P}_{Λ} ; there, the artifact is of order 100%, compared
 2840 to the physical and measured value of $\sim 2\%$ [273].

2841 For symmetric collisions (e.g. Au+Au), the quantity $\overline{\mathcal{R}}_{\Lambda}^z$ must be antisymmetric about
 2842 midrapidity. However, at very forward/backward rapidities, circular vorticity has been re-
 2843 ported in hydrodynamic [274–278] and transport [279–285]. This effect, also visible in the
 2844 left panel in figure 94, arises from strong temperature gradients and edge effects in three-
 2845 dimensional space. It is of very different origin than the ring vorticity of interest here.

2846 Finally, production plane polarization at large x_F has been observed (primarily) in p+p
 2847 and (in some) p+A collisions [286–291] at energies up to $\sqrt{s_{NN}} = 41$ GeV. This effect,
 2848 which is believed to be completely hadronic in origin but remains incompletely understood, is
 2849 distinguishable from the hydrodynamically-driven ring vorticity discussed here by its rapidity
 2850 dependence, which is strongly forward-focused, as well as the fact that $\overline{\Lambda}$ s do not display
 2851 production plane polarization at all. Thus, in addition to double-checking topologically-
 2852 dependent efficiency artifacts (discussed above), it is important that STAR will measure
 2853 the effect both for hyperons and antihyperons to distinguish hydrodynamic from hadronic
 2854 phenomena.

2855 4 Computing Resources

2856 In 2019, STAR submitted the computing resource request for years 2021–2025. Recently,
 2857 there was a proposal to upgrade the STAR DAQ system that will allow STAR to take data at
 2858 approximately double the bandwidth in the 2023-2025 runs as compared to the expected 2022
 2859 rates for which the previous resource request was prepared. The increased DAQ bandwidth
 2860 will improve the statistical precision for various observables aimed towards the detail inves-
 2861 tigation of microscopic structure of QGP. These include the net-proton high order cumulant
 2862 ratios C_6/C_2 , thermal dilepton spectra and low p_T J/ψ v_1 , v_2 etc which are unique at STAR
 2863 compared to sPHENIX at the top RHIC energy. Furthermore, STAR will be able to accom-
 2864 modate the triggers reading out forward and mid-rapidity tracking/calorimeter detectors
 2865 together which offers a unique chance to characterize the QGP over a wide pseudorapidity
 2866 coverage.

2867 An updated request on the additional resources due to this upgrade was submitted to
 2868 SDCC in November 2021. The request was discussed with NPP management at the mini-
 2869 retreat on "Nuclear Physics Computing from RHIC to EIC" in January 2022. We would like
 2870 to emphasize that the requested resources are essential for completing the scientific mission of
 2871 the STAR experiment, by producing and finishing the analyses from the requested datasets

2872 taken in 2023–2025 in a timely fashion.

2873 Table 6 and Table 7 list the updated requests on the network capacity needs and the
 2874 storage/CPU resource needs, respectively.

Network and HPSS capability	2022 capacity	2023-2025 needs
DAQ to SDCC network upload	40 Gbps	40 Gbps
SDCC to DAQ local network	28×1 Gbps	48×1 Gbps
Tape Drive Capacity	20 Gbps	40 Gbps

Table 6: Updated request on network capacity needs

Year	Species	Additional HPSS Space Needed (RAW+DST) (PB)	Total Storage Space Needed (Xrootd) (PB)	Total Storage Space Needed (NFS/Central) (PB)	Required CPU Total [kHS06]
2021	BES-II	0.43	3.06	3.504	203
2022	500 GeV $p+p$	11.07	3.63	3.854	295
2023	200 GeV Au+Au	55.4	7.0	4.75	626
2024	200 GeV $p+p/p+Au$	35.5	9.1	4.75	626
2025	200 GeV Au+Au	73.8	13.5	4.75	626

Table 7: Updated request on storage and CPU resources

2875 5 Future Opportunities

2876 Experience from the BES-II has shown us that the excellent performance from RHIC may
 2877 allow us to take short opportunistic datasets that enable unique physics programs with
 2878 minimal extra running time. Below we outline two such opportunistic programs, both are of
 2879 great interest to STAR and the larger nuclear physics community.

2880 5.1 Imaging shape and radial profile of atomic nuclei via collective 2881 flow measurements

2882 The success of the hydrodynamic framework of heavy-ion collisions permits us today to
 2883 perform quantitative extractions of the transport properties of the QGP via the state-of-the-
 2884 art multi-system Bayesian analysis approaches. [92–94] Such extractions rely largely on a
 2885 correct description of the initial condition of the QGP prior to the hydrodynamic expansion.
 2886 Recent experimental data in $^{96}\text{Ru}+^{96}\text{Ru}$ and $^{96}\text{Zr}+^{96}\text{Zr}$ isobar collisions [292], $^{238}\text{U}+^{238}\text{U}$ [29]
 2887 and $^{129}\text{Xe}+^{129}\text{Xe}$ [293–297] collisions, as well as dedicated theoretical studies [30, 33, 36, 37,
 2888 298–303], have indicated the importance of nuclear deformation and the nuclear radial profile,
 2889 i.e. radial distribution of proton and neutrons in the nucleus, on the measured anisotropic
 2890 flow. However, the impact of these collective nuclear structure effects are not yet considered
 2891 in these Bayesian approaches. For a reliable extraction of transport properties and initial-
 2892 state from the collective flow data, we need to ensure that the uncertainty associated with
 2893 the structure of the colliding ions is under control in the hydrodynamic models, especially
 2894 since all species at RHIC and the LHC are expected to present some deformations and some
 2895 uncertainties in the nuclear skin and radius (as indicated in Table 8 for nuclear deformation).
 2896 These uncertainties can be gauged precisely using pairs of isobar collisions, as demonstrated
 2897 by the $^{96}\text{Ru}+^{96}\text{Ru}$ and $^{96}\text{Zr}+^{96}\text{Zr}$ collisions at RHIC, where the ratio of flow observables can
 2898 be determined with $< 0.4\%$ precision [292]. Note that these ratios are made at the same N_{ch}
 2899 in each isobar, and therefore are essentially insensitive to final state effects and are precision
 probes of the initial conditions as we shall discuss below.

	β_2	β_3	β_4
^{238}U	0.286 [304]	0.078 [305]	0.09 [306]
^{208}Pb	0.05 [304]	0.04 [307]	?
^{197}Au	-(0.13-0.16) [306, 308]	?	-0.03 [306]
^{129}Xe	0.16 [306]	?	?
^{96}Ru	0.05-0.16 [304, 306]	?	?
^{96}Zr	0.08 [304]	?	0.06 [306]

Table 8: Some estimates of the deformation values β_2, β_3 , and β_4 for the large nuclei collided at RHIC and the LHC with references given, mostly on global analysis of $B(En)$ transition data over a broad range of nuclei. There are also uncertainties in their values for surface diffuseness a_0 and half radius R_0 which are not listed.

2900

2901 It is straightforward to see why the geometry of heavy-ion collisions is sensitive to nuclear
 2902 deformation and radial profile. We refer to the cartoon in Fig. 96. A nucleus can be modeled
 2903 through a nucleon density of Woods-Saxon form:

$$\rho(r, \theta, \phi) = \frac{\rho_0}{1 + e^{[r - R(\theta, \phi)]/a_0}}, \quad R(\theta, \phi) = R_0 (1 + \beta_2[\cos \gamma Y_{2,0} + \sin \gamma Y_{2,2}] + \beta_3 Y_{3,0} + \beta_4 Y_{4,0}), \quad (8)$$

2904 where the nuclear surface $R(\theta, \phi)$ includes only the most relevant deformation components
 2905 from nuclear structure physics, quadrupole $n = 2$, octupole $n = 3$ and hexadecapole $n = 4$.
 2906 The angle $0 \leq \gamma \leq \pi/3$ controls the triaxiality of the quadrupole deformation or the three
 2907 radii R_a, R_b, R_c of the ellipsoid, with $\gamma = 0$ corresponds to prolate ($R_a = R_b < R_c$), and
 2908 $\gamma = \pi/3$ corresponds to oblate ($R_a < R_b = R_c$). The nuclear radial profile is controlled
 2909 by the surface diffuseness or nuclear skin a_0 and half radius R_0 . In heavy-ion collisions,
 2910 the shape of the deformed ions strongly affects the geometry of overlap. The entire mass
 2911 distribution is probed simultaneously, and one can use multi-particle correlation observables
 2912 to infer information of all these parameters. This way of probing nuclear densities is different
 2913 from the standard techniques of low-energy physics, where β_n, a_0 and R_0 are inferred from
 2914 the orientation-averaged form factor data from e+A and hadron+A scatterings and multipole
 2915 transition probabilities, $B(E_n)$, between low-lying rotational states. Furthermore, the time
 2916 scales involved in high-energy heavy-ion collisions are much shorter ($< 10^{-24}$ s), than the
 2917 typical timescale of the EM transition involved in the rotational bands (typically on the
 2918 order of 10^{-20} s [309]). As we shall also argue below, a remarkable question is whether
 2919 the manifestation of nuclear deformation and nuclear skin – collective features of the nuclear
 2920 many-body system – is the same across energy scales.

2921 The presence of multipoles, β_n , in the colliding ions modifies non-trivially the corre-
 2922 sponding spatial anisotropy, ε_n , of the produced QGP, and consequently the final-state flow
 2923 harmonic, v_n . Similarly, different values of a_0 and R_0 modify the effective size of the over-
 2924 lap region and therefore the “radial” flow or the event-by-event mean transverse momentum
 2925 $[p_T]$. [37] Recent studies show that nuclear skin a_0 also impacts the v_2 , and simple event ac-
 2926 tivity observables such as multiplicity distributions $p(N_{\text{ch}})$ and participants $p(N_{\text{part}})$. [36, 303]
 2927 Predictions for many other observables and their sensitivities to nuclear deformation and nu-
 2928 clear skin have been made, such as p_T fluctuations [32], spectator neutron production [310],
 2929 mixed-flow harmonics [311], and v_n - p_T correlations [297, 300, 312].

2930 Earlier studies of nuclear deformation are mainly focused on the elliptic flow, v_2 in central
 2931 collisions. They have established a simple relation between quadrupole deformation and ε_2
 2932 and v_2 [312, 313],

$$\langle \varepsilon_2^2 \rangle = a' + b' \beta_2^2, \quad \langle v_2^2 \rangle = a + b \beta_2^2, \quad (9)$$

2933 where the a' and a are mean-squared eccentricity and elliptic flow without deformation, while
 2934 the b' and b describe the parametric dependence of the deformation-enhanced component of
 2935 eccentricity and elliptic flow, respectively. The strict quadratic dependence of Eq. 9 leads to a
 2936 very robust equation relating the β_2 between any pair of collision systems. Applied to RHIC
 2937 data, it allows one to derive a constraint on the $\beta_{2,U}$ and $\beta_{2,Au}$, as shown in the right panel of
 2938 Fig. 97. This highlights how, at present, the low-energy nuclear structure model calculation

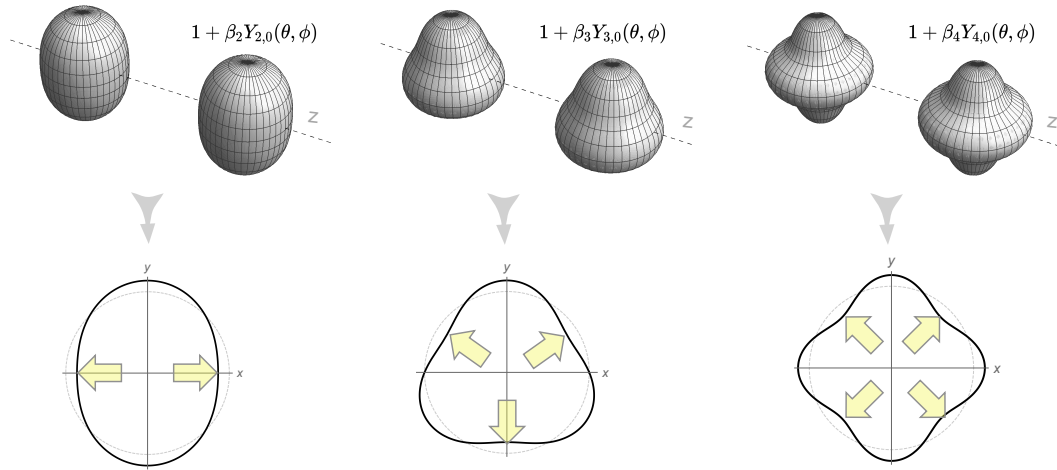


Figure 96: A cartoon of a collision of nuclei with quadrupole (left), octupole (middle) and hexadecapole (right) deformations including only the $Y_{n,0}$ mode and with $\beta_n = 0.25$ (we ignore the large Lorentz contraction in the z -direction). The bottom row shows how the initial condition of the medium formed after the collision looks in the transverse plane. The yellow arrows indicate the direction of maximum pressure gradients along which the medium expands with the largest velocity, leading to final state harmonic flow v_n with n -fold symmetry.

2939 and the flow data from high-energy nuclear collisions are fairly inconsistent. Relations similar
 2940 to Eq. 9 can also be written down for v_3 and v_4 , which can be used to potentially constrain
 octupole and hexadecapole deformations [31].

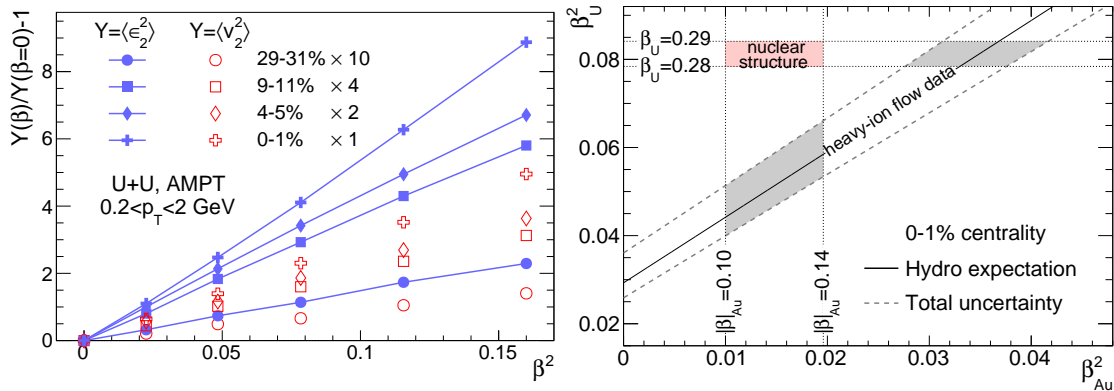


Figure 97: Left panel: $\langle v_2^2(\beta) \rangle / \langle v_2^2(0) \rangle - 1 = b/a \beta_2^2$ (empty symbols) and $\langle \epsilon_2^2(\beta_2) \rangle / \langle \epsilon_2^2(0) \rangle - 1 = b'/a' \beta_2^2$ (full symbols) as a function of β_2^2 in U+U collisions from the AMPT model. Different symbols correspond to different centrality classes. Right panel: $\beta_{2,U}^2$ as a function of $\beta_{2,Au}^2$. The region between the dashed lines is consistent with the hydrodynamic expectation based on Eq. (9) and STAR v_2 data in 0–1% centrality. Figures taken from Ref. [302].

2941

The most precise tool for structure imaging, however, is provided by collision of iso-

baric systems, as demonstrated by recent measurements in $^{96}\text{Ru}+^{96}\text{Ru}$ and $^{96}\text{Zr}+^{96}\text{Zr}$ collisions. [292] The crucial point is that since isobar nuclei have the same mass number, deviations from unity of the ratio of any observable must originate from differences in their structures, which impact the initial state of QGP and its final state observables. Ratios of many observables between $^{96}\text{Ru}+^{96}\text{Ru}$ and $^{96}\text{Zr}+^{96}\text{Zr}$, both published and new preliminary results shown in QM2022 [292, 314], show deviations from unity in an observable- and centrality-dependent manner, which must originate from differences in their structures. Model studies show that the isobar ratio for a given observable \mathcal{O} probes only the nuclear structure parameter differences, i.e. $\Delta\beta_n^2 = \beta_{n\text{Ru}}^2 - \beta_{n\text{Zr}}^2$, $\Delta a_0 = a_{0\text{Ru}} - a_{0\text{Zr}}$ and $\Delta R_0 = R_{0\text{Ru}} - R_{0\text{Zr}}$ [34]:

$$R_{\mathcal{O}} \equiv \frac{\mathcal{O}_{\text{Ru}}}{\mathcal{O}_{\text{Zr}}} \approx 1 + c_1\Delta\beta_2^2 + c_2\Delta\beta_3^2 + c_3\Delta a_0 + c_4\Delta R_0, \quad (10)$$

2942 where the coefficients c_1 – c_4 describes how the heavy-ion initial state is controlled by the
 2943 nuclear structure and are weak functions of system size. Figure 98 highlights some recent
 2944 measurements: ratios of multiplicity distribution $p(N_{\text{ch}})$, v_2 , v_3 , variance of p_{T} fluctuations
 2945 $\langle\delta p_{\text{T}}^2\rangle/\langle p_{\text{T}}\rangle^2$, and $\langle p_{\text{T}}\rangle$ between the isobar systems. All of them show non-monotonic central-
 2946 ity dependence similar in shape to the theoretical predictions that include effects of nuclear
 2947 skin as well as nuclear deformations. [33, 34, 311, 315] In particular, the data imply a larger
 2948 quadrupole deformation β_2 in ^{96}Ru , a larger octupole deformation β_3 in ^{96}Zr , and a larger
 2949 a_0 value consistent with a larger neutron skin in ^{96}Zr , Δr_{np} , defined as the rms radius dif-
 2950 ference between the neutron and proton distributions: $\langle r_n^2\rangle^{1/2} \equiv \langle r_n^2\rangle^{1/2} - \langle r_p^2\rangle^{1/2}$. These
 2951 detailed measurements over-constrain the WS parameters and can be used to test the initial
 2952 conditions used in hydrodynamic models. Note that the neutron skin thickness Δr_{np}
 2953 is directly related to the slope parameter L for the density dependence of the symmetry
 2954 energy, which is particularly important in astrophysics concerning neutron stars. [316] The
 2955 preliminary extraction of L from the measured $\langle p_{\text{T}}\rangle$ ratio in the isobar data seems to prefers
 2956 a value of 47–70 MeV as shown in Fig. 98, quite consistent with low-energy nuclear reaction
 2957 measurements [317] but systematically lower than the PREXII results [318].

An additional observable showing large sensitivity to the nuclear quadrupole deformation is the Pearson correlation coefficient, $\rho(v_2^2, [p_{\text{T}}])$, between v_2 and the mean transverse momentum, $[p_{\text{T}}]$. This observable probes in particular the full quadrupole structure of the colliding ions, i.e., both β_2 and its triaxiality γ in Eq. 8 [32, 312],

$$\rho(v_2^2, [p_{\text{T}}]) \approx a - b \cos(3\gamma)\beta^3, \quad a, b > 0. \quad (11)$$

2958 Therefore prolate deformation in the colliding nuclei is expected to reduce $\rho(v_2^2, [p_{\text{T}}])$, while
 2959 oblate deformation is expected to increase it. This observable has been measured by the
 2960 STAR collaboration in U+U and Au+Au collisions, which established unambiguously the
 2961 large and dominating influence of the nuclear quadrupole deformation, see Fig. 99(a). The
 2962 large prolate deformation of ^{238}U yields a strong negative contribution to the $v_2 - [p_{\text{T}}]$ cor-
 2963 relation, enough to make it change sign. A large impact of $\beta_{2\text{U}}$ has further been observed in
 2964 the fluctuations of $[p_{\text{T}}]$. The same measurement is also performed by the ATLAS and ALICE
 2965 collaborations in $^{129}\text{Xe}+^{129}\text{Xe}$ and $^{208}\text{Pb}+^{208}\text{Pb}$ collisions [296, 319], see Fig. 99(b). A com-
 2966 parison with a Trento model calculation based on input from nuclear structure theory [297]

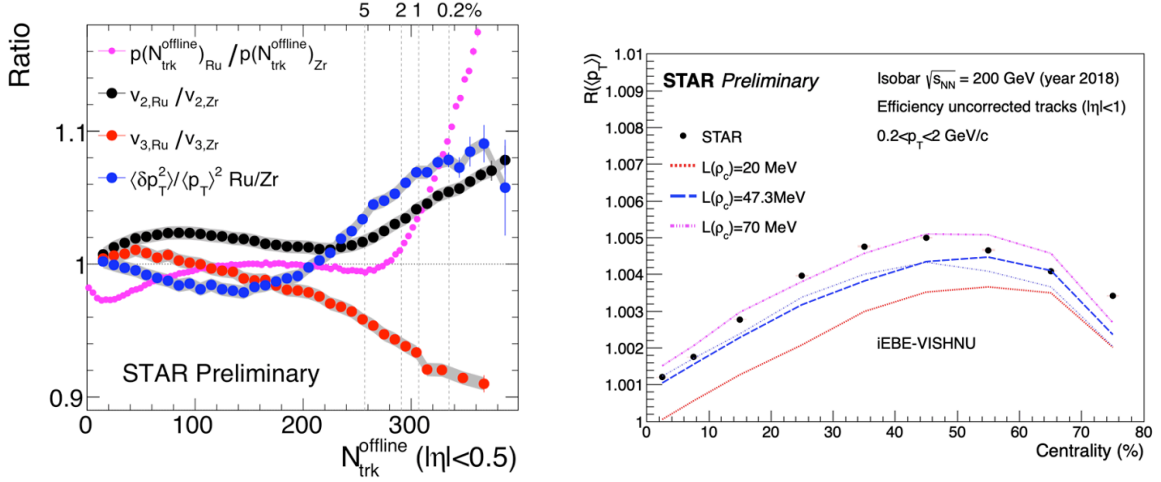


Figure 98: Left panel: STAR preliminary results of isobar ratio of $p(N_{\text{ch}})$, v_2 , v_3 , and variances $\langle \delta p_T^2 \rangle / \langle p_T \rangle^2$ as a function of N_{ch} . Right panel: The centrality dependence of the Ru+Ru/Zr+Zr ratio of $\langle p_T \rangle$, compared with hydrodynamic model calculations [37].

2967 provide strong evidence that ^{129}Xe is a highly-deformed triaxial ellipsoid with an overall
 2968 quadrupole deformation of $\beta_{2\text{Xe}} \sim 0.2$ and triaxiality of $\gamma_{\text{Xe}} \sim \pi/6$. Hydrodynamic models
 2969 based on state-of-the-art initial conditions with deformation values from Table 8 struggle to
 2970 describe quantitatively all these experimental measurements. [25, 320, 321] The reason could
 2971 be that the radial flow response of the system to fluctuations induced by the deformation of
 2972 the colliding ions is not fully captured by the existing models. Collisions of well-deformed
 2973 ions, and their comparisons with the collisions of more spherical species, provide us with a
 new way to test the hydrodynamic description.

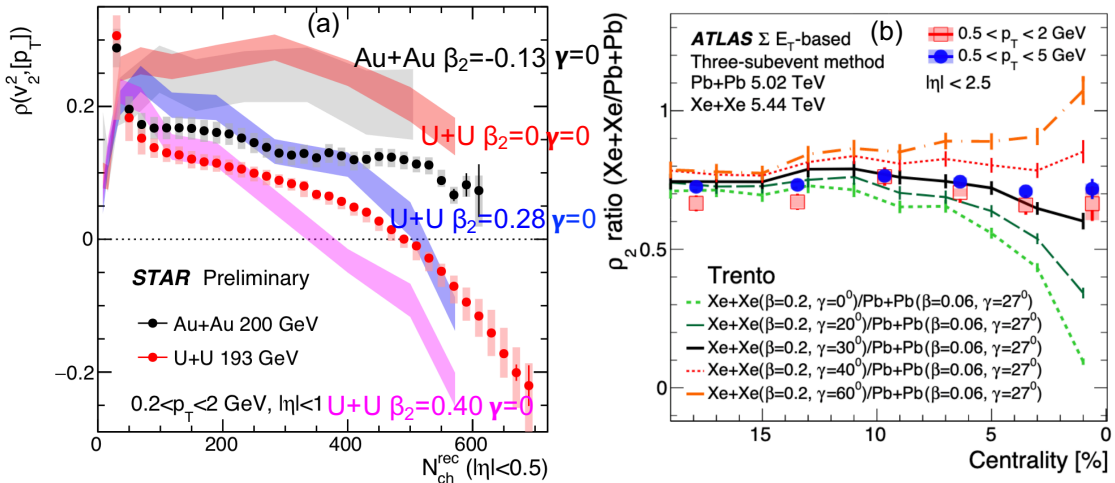


Figure 99: Left panel: STAR preliminary results of Pearson correlation coefficient $\rho(v_2^2, [p_T])$ in U+U and Au+Au collisions, showing a sign-change due to large prolate deformation of ^{238}U . Right panel: ATLAS results of the ratio of $\rho(v_2^2, [p_T])$ between Xe+Xe and Pb+Pb collisions, showing a strong preference for ^{129}Xe being a highly-deformed triaxial ellipsoid.

2974

2975 To summarize, flow measurements in heavy-ion collisions have large potential to provide
2976 detailed information on the shape and radial profile of colliding nuclei. By connecting the
2977 highest and lowest energy scales, they allow us to answer important questions in heavy ion
2978 physics and have broader impact to the larger nuclear physics community. Here are a few of
2979 them:

- 2980 • How distributions of protons and neutrons in atomic nuclei give rise to the complex
2981 initial condition of heavy ion collisions? Can we use nuclear shapes and nuclear radial
2982 profiles as additional handles to understand particle production and generation of ec-
2983 centricities, e.g. by comparing flow observables at the same final state multiplicity in
2984 isobar systems with different nuclear structures?
- 2985 • Can we gauge uncertainties in the extraction of the transport properties of the QGP
2986 due to uncertainties in the initial condition arising from nuclear structure?
- 2987 • Are the nuclear shape and radial profile inferred from hydrodynamic response the
2988 same as those measured in nuclear structure experiments? Can isobar collisions serve
2989 as a precision tool for the extraction of the neutron skin, competitive to the exist-
2990 ing measurements? and what are the energy and longitudinal dependence of nuclear
2991 structures?

2992 To address these and other related questions, several workshops exploring the intersection
2993 between nuclear structure and heavy-ion collisions have been planned, including a month-
2994 long INT program in early 2023. Rapid progress is expected in the next two years.

2995 Thus we propose to collide more species to extract their value of deformation parameters
2996 β_2 , γ , β_3 and β_4 , and a_0 and associated neutron skin from flow measurements, with a twofold
2997 purpose: 1) provide a new handle on the initial state and hydrodynamic response of the
2998 QGP, 2) perform studies of nuclear structure physics at high energy to complement the
2999 information coming from lower energies, and so assess the consistency of nuclear phenomena
3000 across energy scales. The ground state of almost all stable nuclei is deformed (see for example
3001 the interactive chart in Ref. [322]). RHIC, with its flexibility to collide almost any nuclei
3002 from $p+p$ to $U+U$ is a unique facility to perform such studies in the foreseeable future. The
3003 best example to showcase this capability is the run of isobars performed in 2018, where the
3004 two systems, $Zr+Zr$ and $Ru+Ru$, were alternated on a fill-by-fill basis, leading to extremely
3005 small systematic uncertainties on the final observables. [292] This allows one to detect minute
3006 differences in the physics observables such as multiplicity, $[p_T]$ and v_n in the comparison of the
3007 two systems. Consequently, even small differences in the values of β_n and a_0 of the colliding
3008 systems can be precisely mapped. [30] For each species, we need roughly 100 million minimum
3009 bias and 50 million 0–5% central events. Assuming the standard 50% RHIC+STAR up time
3010 and 1.5 KHz DAQ rate, same as $Au+Au$ running, we will be able to collect 130M minbias
3011 events and 64M central events in three days of physics running. This is slightly less than
3012 the existing $U+U$ dataset taken in 2011, but with comparable statistical precision due to
3013 the increased acceptance from the iTPC. Adding two days of setup time, this leads to about
3014 five days of total time for each species.

3015 The system scan we propose can be divided into two steps. Given the tight schedule for
 3016 the next few years, instead of making an explicit proposal on how much running time are
 3017 needed to fully explore these topics, we discuss what can be achieved if we are given certain
 3018 number of days.

3019 • **≈10 days:** In the first step, we would like to scan two nuclei in the vicinity of the
 3020 most studied species at RHIC, ^{197}Au , to improve the modeling of Au+Au collisions,
 3021 information which is crucial for the future precision interpretation of high-statistics
 3022 data expected from Run-23+25. To achieve this, ideal candidates are ^{208}Pb and ^{196}Hg
 3023 (^{198}Hg could be a substitute). Having ^{208}Pb at $\sqrt{s_{\text{NN}}} = 200$ GeV provides a crucial
 3024 bridge with the ^{208}Pb at LHC energies: comparison between ^{208}Pb measurements at
 3025 RHIC and the LHC will constrain any possible energy dependence of the initial state
 3026 effects and pre-equilibrium dynamics. Additionally, ^{208}Pb is nearly spherical, so that
 3027 Pb+Pb collisions at the same energy will allow us to better understand the impact
 3028 of the moderate deformation of ^{197}Au in Au+Au collisions, as well as the impact
 3029 of the difference of a_0 parameter and neutron skin between ^{197}Au and ^{208}Pb . The
 3030 Hg+Hg collisions would then permit us to understand more deeply the nature of the
 3031 deformation of ^{197}Au , which, being an odd-mass nucleus, hasn't been determined in
 3032 low-energy experiments. ^{196}Hg is an oblate nucleus with $|\beta_2| \approx 0.1$, and the observable
 3033 $\rho(v_2^2, [p_{\text{T}}])$ can be used quantify whether ^{197}Au is more or less oblate than ^{196}Hg , an
 3034 information which will gauge more tightly the initial geometry of Au+Au collisions.
 3035 Adding Hg+Hg collisions will also provide an independent cross-check on the initial
 3036 state, for example one can setup three relations like Eq. 9 from Pb+Pb, Hg+Hg and
 3037 Au+Au to triangulate the consistency of the three deformation values. [31]

3038 • **Additional time:** In the second step, our proposal is to use hydrodynamics and
 3039 flow measurements to perform precision cross-checks of low-energy nuclear physics by
 3040 constraining the evolution of the quadrupole deformation and neutron skin along the
 3041 chain of stable samarium isotopes. It would be useful in particular to collide three
 3042 isotopes: ^{144}Sm ($\beta_2 = 0.08$, as spherical as ^{208}Pb), ^{148}Sm ($\beta_2 = 0.14$, triaxial much as
 3043 ^{129}Xe and ^{197}Au), and ^{154}Sm ($\beta_2 = 0.34$ well-deformed like ^{238}U). The evolution of the
 3044 quadrupole deformation can be mapped precisely at RHIC, thus offering a valuable
 3045 test of nuclear structure knowledge. If data on $^{154}\text{Sm}+^{154}\text{Sm}$ collisions is available, it
 3046 would be desirable to also have $^{154}\text{Gd}+^{154}\text{Gd}$ ($\beta_2 = 0.31$) collisions. The comparison
 3047 between the two well-deformed isobaric systems could potentially yield the most precise
 3048 information about the relative deformation and relative neutron skin between two
 3049 ground state nuclei. Theoretical studies further suggest that ground states in the
 3050 region $Z \sim 56/N \sim 88$ [323] (including the samarium isotopes) may display enhanced
 3051 octupole correlations, i.e., β_3 values. These would manifest in high-energy collisions as
 3052 enhanced v_3 , as well as in the correlators $\rho(v_3^2, [p_{\text{T}}])$. Such enhancements are already
 3053 observed in $^{96}\text{Zr}+^{96}\text{Zr}$ relative to $^{96}\text{Ru}+^{96}\text{Ru}$ collisions (Fig. 98 and Ref. [314]), however
 3054 nuclear structure modeling for these medium mass nuclei are quite challenging and it
 3055 is unclear yet whether the observed enhancements are due to octupole correlation or
 3056 static octupole deformation. The heavier species mentioned above would be a more

sensitive choice for identifying static octupole deformation. The study of octupole deformation is also fundamentally interesting because nuclei with large β_3 provides a stringent test of the electric-dipole moment (EDM) [324]. The exact choice of species is still under refinement, presently we have a preference for ^{154}Sm and ^{148}Sm , followed by ^{154}Gd and ^{144}Sm .

Finally, one should note that the STAR DAQ rate for these moderate-sized systems could be significantly larger, possibly reaching 2KHz. This enhanced DAQ rate will compensate partially the smaller number of charged particles expected in these systems compared to larger systems.

5.2 Fixed-target Measurements Using Light Beam and Target Combinations

Although the proposed fixed-target Au+Au energy scan has been completed, if the opportunity exists for further measurements, light beam and target combinations could help to clarify the role and mechanisms of nucleon stopping. Indeed, STAR was recommended to consider installing a beryllium target, that being the lowest Z feasible solid target which could work with the target apparatus. This was not done previously because changing the target requires opening the STAR beampipe and removing the existing target, and that could not be done until the Au+Au energy scan had been completed. Both the collider and STAR have demonstrated that fixed-target runs can be quickly tuned, as the demands on collider operations are modest, and efficiently run, as the collider can control and deliver sufficient intensity to fill the STAR DAQ bandwidth and the experiment can cleanly trigger on these events.

Recently it has come to the attention of the STAR collaboration that fixed-target collisions using light beam and target combinations could also benefit the Space Radiation Protection community. Cosmic rays are a serious concern to astronauts, electronics, and spacecraft. Although 90% of the cosmic ray flux is comprised of energetic protons and another 9% is Helium nuclei, the remaining 1%, which is made up of nuclei from Li to Fe, is not negligible both because the energy loss is proportional to Z^2 and because additional damage is done by the energetic light nuclei (p, d, t, ^3He , and ^4He) produced through the fragmentation of the target and projectile nuclei. The damage done by the light nuclei becomes increasingly important for higher energy cosmic rays. Light ion cross section measurements represent the largest uncertainty in space radiation estimates. The energy spectrum of cosmic rays in the solar system is concentrated at energies below 1 GeV/n. Extensive measurements have been made using the dedicated NSRL facility at the booster, and at other lower energy facilities. However, the Space Radiation Community has recently identified higher energy systems, using beams from 3 to 50 GeV/n on C, Al, and Fe targets as one of the next areas of need. [?] This energy range is dominated by Galactic Cosmic Rays (GCR). The requirements would be to measure the cross section for light nucleus (p, d, t, ^3He , and ^4He) production through fragmentation of the target and projectile. STAR has excellent particle identification for all of these particle species using both dE/dx and

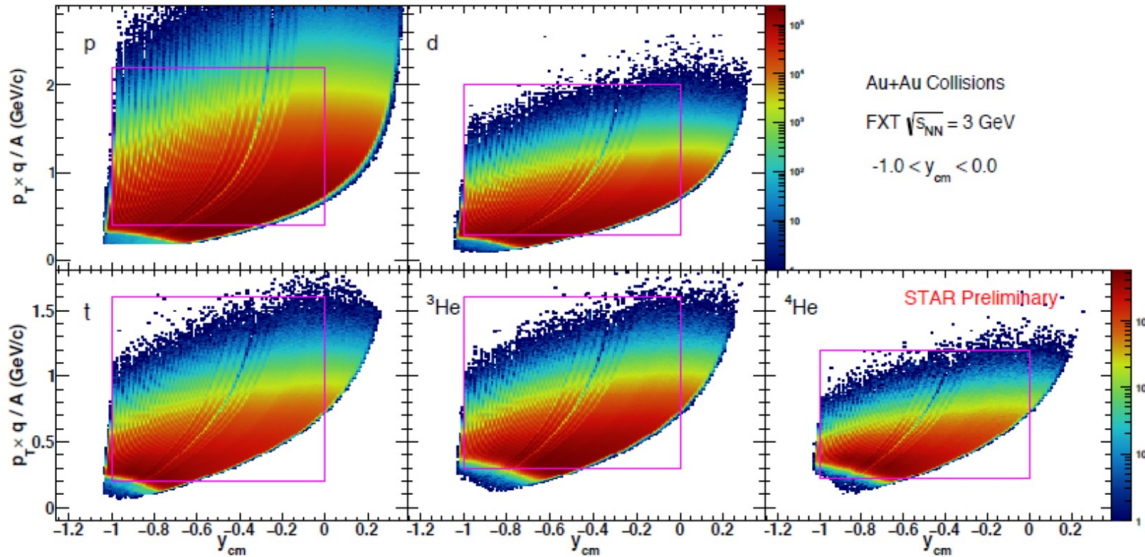


Figure 100: The acceptance for light nuclei (p, d, t, ${}^3\text{He}$, and ${}^4\text{He}$) achieved in the $\sqrt{s_{\text{NN}}}=3\text{ GeV}$ Au+Au system using both dE/dx and ToF.

3097 time-of-flight (capabilities specifically identified as essential in the NASA report [?], however
 3098 the acceptance is only in the target-side of the rapidity distribution (see Fig. 100. For sym-
 3099 metric systems this is not a problem. This can be seen in Fig. 101 which shows the rapidity
 3100 densities (dN/dy) for light nuclei. The results are reflected about midrapidity. The figure
 3101 shows that the light nuclei associated with target fragmentation are seen in the less central
 3102 collisions. The projectile fragmentation can be inferred by reflection. For asymmetric sys-
 3103 tems, for which reflection symmetry is not possible, inference of the projectile fragmentation
 3104 would require both light-on-heavy and heavy-on-light combinations. STAR has reached out
 3105 to determine if the STAR detector has sufficient acceptance in p_T and y to meet the needs
 3106 of the Space Radiation Protection community. An overview of the RHIC/STAR capabil-
 3107 ities was presented at the Workshop for Applied Nuclear Data Activities (WANDA2022)
 3108 conference in February of 2022. In the session summary the opportunity to make these
 3109 measurements at RHIC was characterized as a “unique, time-limited opportunity to obtain
 3110 critical high-energy data”.

3111 NASA had been considering constructing detector systems to make these measurements
 3112 at the FAIR facility at GSI in Darmstadt, Germany. STAR is an existing detector with the
 3113 required capabilities and analysis teams that have proven expertise to measure the light nuclei
 3114 cross sections in fixed-target experiments. The RHIC facility has demonstrated capability
 3115 to efficiently deliver the required beams. In addition, there is significant uncertainty about
 3116 when the SIS-100 accelerator will be available as the construction timeline has been disrupted
 3117 by the war in Ukraine and the cessation of cooperation between Germany and Russia.

3118 As it has been determined that the measurements that could be made at RHIC using the
 3119 STAR detector will meet the needs of the Space Radiation Protection community, STAR

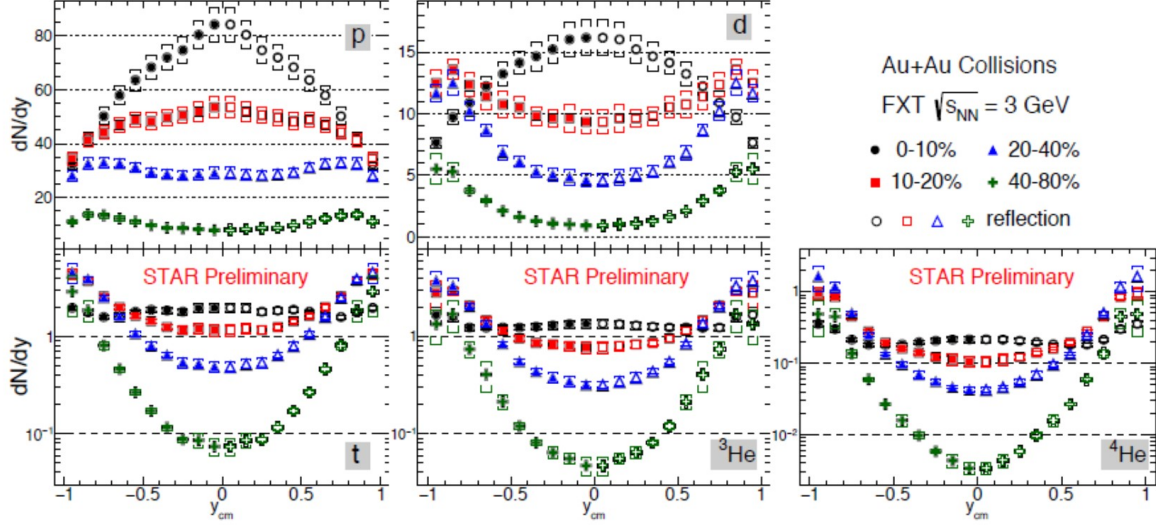


Figure 101: Preliminary dN/dy results for light nuclei (p , d , t , ${}^3\text{He}$, and ${}^4\text{He}$) measured in the $\sqrt{s_{\text{NN}}} = 3$ GeV Au+Au system using both dE/dx and ToF.

3120 is proposing brief energy scans using C, Si, and Fe beams on light targets (C, Al, and Fe).
 3121 We propose three energies for each beam ($E_{\text{Tot}} = 6, 21, \text{ and } 51$ GeV, $E_{\text{Kin}} = 5, 20, \text{ and } 50$,
 3122 or $\sqrt{s_{\text{NN}}} = 3.6, 6.4, \text{ and } 9.8$ GeV respectively). For each beam, the collider would need 12
 3123 hours to develop the beam (this was the amount of time needed to develop the individual
 3124 beams for the Au+Au FXT energy scan). In order to get enough statistics on each of the
 3125 three targets, 36 hours would be needed for each beam-energy combination. Additionally,
 3126 it is likely that the collider would need some time to reconfigure to circulate low energy
 3127 beams (approximately one day). Therefore the request is for three weeks of beam time (one
 3128 week for each of the three beam species). The STAR collaboration considers the full energy
 3129 Au+Au, p +Au, and p + p programs to be the highest priority, and this opportunity would
 3130 only be considered if addition weeks of operations were available.

Beam	Energy	Targets	Time
Machine Setup			1 day
Carbon	5 GeV	C, Al, Fe	2 days
Carbon	20 GeV	C, Al, Fe	2 days
Carbon	50 GeV	C, Al, Fe	2 days
Total			1 week
Machine Setup			1 day
Aluminum	5 GeV	C, Al, Fe	2 days
Aluminum	20 GeV	C, Al, Fe	2 days
Aluminum	50 GeV	C, Al, Fe	2 days
Total			1 week
Machine Setup			1 day
Iron	5 GeV	C, Al, Fe	2 days
Iron	20 GeV	C, Al, Fe	2 days
Iron	50 GeV	C, Al, Fe	2 days
Total			1 week
Grand Total			3 weeks

Table 9: Summary of the FXT beam/target scan request. Assumptions are 12 hours of beam development for each energy and 36 hours of physics running (12 hours for each of the three targets). Additionally one day would be needed to configure RHIC for low energy running.

3131 **6 Charge for the 2022 NPP PAC**

3132 BNL Nuclear Physics PAC 2022 Charge - March 20, 2022

3133 Charge

3134 STAR: Beam Use Requests for Runs 23-25

3135 sPHENIX: Beam Use Requests for Runs 23-25

3136 CeC: Beam Use Requests

3137

3138 The Beam Use Requests should be submitted in written form to PAC by May 6, 2022

3139 The BURs should be based on the following number of cryo-weeks. The first number
3140 is the proposed RHIC run duration for scenario 1 and the second number corresponds to
3141 optimal duration (scenario 2) presented to the DOE-ONP in BNL's FY24 Lab Managers'
3142 Budget Briefing:

3143 • 2023: 24 (28)

3144 • 2024: 24 (28)

3145 • 2025: 24 (28)

3146 Note the eventual running cryo-weeks for each run will depend on the final budget guid-
3147 ance for that year so it can be lower than 24 weeks.

3148 Presentations: STAR: Report on Run 2022, update on BES-II, small systems and spin
3149 physics analyses, and the latest development regarding the Isobar results.

3150 CeC X: Results from Run 2022

3151 PHENIX: Update on ongoing analysis efforts and data archiving efforts

3152 sPHENIX: Installation status and schedule including TPOT status, commissioning, com-
3153 puting plan and readiness for data taking.

3154 Written report from the PAC is expected within two weeks after the meeting.

References

- 3155 [1] H.-j. Xu *et al.*, Chin. Phys. **C42**, 084103 (2018), 1710.07265.
- 3156 [2] S. A. Voloshin, Phys. Rev. **C98**, 054911 (2018), 1805.05300.
- 3157 [3] STAR, M. Abdallah *et al.*, Phys. Rev. Lett. **128**, 092301 (2022), 2106.09243.
- 3158 [4] Y. Feng, J. Zhao, H. Li, H.-j. Xu, and F. Wang, Phys. Rev. C **105**, 024913 (2022),
3159 2106.15595.
- 3160 [5] STAR, N. M. Abdelwahab *et al.*, Phys. Lett. **B745**, 40 (2015), 1409.2043.
- 3161 [6] Y. Feng, Y. Lin, J. Zhao, and F. Wang, Phys. Lett. B **820**, 136549 (2021), 2103.10378.
- 3162 [7] J. Adams *et al.*, Nucl. Instrum. Meth. A **968**, 163970 (2020), 1912.05243.
- 3163 [8] Z.-T. Liang and X.-N. Wang, Phys. Rev. Lett. **94**, 102301 (2005), nucl-th/0410079,
3164 [Erratum: Phys.Rev.Lett. 96, 039901 (2006)].
- 3165 [9] S. A. Voloshin, (2004), nucl-th/0410089.
- 3166 [10] K. Schilling, P. Seyboth, and G. E. Wolf, Nucl. Phys. B **15**, 397 (1970), [Erratum:
3167 Nucl.Phys.B 18, 332 (1970)].
- 3168 [11] X.-L. Sheng, L. Oliva, and Q. Wang, Phys. Rev. D **101**, 096005 (2020), 1910.13684.
- 3169 [12] STAR, J. Adam *et al.*, Phys. Rev. C **98**, 014910 (2018), 1805.04400.
- 3170 [13] A. Bilandzic, C. H. Christensen, K. Gulbrandsen, A. Hansen, and Y. Zhou, Phys. Rev.
3171 **C89**, 064904 (2014), 1312.3572.
- 3172 [14] J. Jia, M. Zhou, and A. Trzupek, Phys. Rev. C **96**, 034906 (2017), 1701.03830.
- 3173 [15] N. Borghini, P. M. Dinh, and J.-Y. Ollitrault, Phys. Rev. C **63**, 054906 (2001), nucl-
3174 th/0007063.
- 3175 [16] ATLAS, M. Aaboud *et al.*, JHEP **01**, 051 (2020), 1904.04808.
- 3176 [17] A. Behera, M. Nie, and J. Jia, Phys. Rev. Res. **2**, 023362 (2020), 2003.04340.
- 3177 [18] B. Schenke and S. Schlichting, Phys. Rev. **C94**, 044907 (2016), 1605.07158.
- 3178 [19] P. Bozek, W. Broniowski, and J. Moreira, Phys. Rev. **C83**, 034911 (2011), 1011.3354.
- 3179 [20] C. Shen and B. Schenke, Phys. Rev. **C97**, 024907 (2018), 1710.00881.
- 3180 [21] L.-G. Pang, G.-Y. Qin, V. Roy, X.-N. Wang, and G.-L. Ma, Phys. Rev. C **91**, 044904
3181 (2015), 1410.8690.
- 3182

- 3183 [22] G. Giacalone, B. Schenke, and C. Shen, Phys. Rev. Lett. **125**, 192301 (2020),
3184 2006.15721.
- 3185 [23] P. Bozek, Phys. Rev. C **93**, 044908 (2016), 1601.04513.
- 3186 [24] P. Bozek and H. Mehrabpour, Phys. Rev. C **101**, 064902 (2020), 2002.08832.
- 3187 [25] B. Schenke, C. Shen, and D. Teaney, Phys. Rev. C **102**, 034905 (2020), 2004.00690.
- 3188 [26] G. Giacalone, F. G. Gardim, J. Noronha-Hostler, and J.-Y. Ollitrault, Phys. Rev. C
3189 **103**, 024909 (2021), 2004.01765.
- 3190 [27] S. H. Lim and J. L. Nagle, (2021), 2103.01348.
- 3191 [28] ALICE, S. Acharya *et al.*, JHEP **07**, 103 (2018), 1804.02944.
- 3192 [29] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **115**, 222301 (2015), 1505.07812.
- 3193 [30] G. Giacalone, J. Jia, and V. Somà, (2021), 2102.08158.
- 3194 [31] J. Jia, Phys. Rev. C **105**, 014905 (2022), 2106.08768.
- 3195 [32] J. Jia, Phys. Rev. C **105**, 044905 (2022), 2109.00604.
- 3196 [33] C. Zhang and J. Jia, Phys. Rev. Lett. **128**, 022301 (2022), 2109.01631.
- 3197 [34] J. Jia and C.-J. Zhang, (2021), 2111.15559.
- 3198 [35] H.-J. Xu *et al.*, Phys. Rev. Lett. **121**, 022301 (2018), 1710.03086.
- 3199 [36] H. Li *et al.*, Phys. Rev. Lett. **125**, 222301 (2020), 1910.06170.
- 3200 [37] H.-j. Xu *et al.*, (2021), 2111.14812.
- 3201 [38] U. Gürsoy, D. Kharzeev, and K. Rajagopal, Phys. Rev. C **89**, 054905 (2014), 1401.3805.
- 3202 [39] STAR, J. Adam *et al.*, Phys. Rev. Lett. **127**, 052302 (2021), 1910.12400.
- 3203 [40] M. A. *et al.* [STAR], (2022), 2204.01625.
- 3204 [41] M. A. *et al.* [STAR], Phys. Rev. Lett. **128** (2022).
- 3205 [42] STAR, M. Abdallah *et al.*, Phys. Lett. B **825**, 136865 (2022), 2110.09666.
- 3206 [43] STAR, M. Abdallah *et al.*, Phys. Rev. D **105**, 032007 (2022), 2109.13191.
- 3207 [44] STAR, L. Adamczyk *et al.*, Phys. Lett. B **771**, 13 (2017), 1607.07517.
- 3208 [45] STAR, J. Adam *et al.*, Phys. Lett. B **797**, 134917 (2019), 1905.13669.

- 3209 [46] PHENIX, A. Adare *et al.*, Phys. Rev. Lett. **101**, 122301 (2008), 0801.0220.
- 3210 [47] S. Marzani, G. Soyez, and M. Spannowsky *Looking inside jets: an introduction to jet*
3211 *substructure and boosted-object phenomenology* Vol. 958 (Springer, 2019), 1901.10342.
- 3212 [48] A. J. Larkoski, S. Marzani, and J. Thaler, Phys. Rev. D **91**, 111501 (2015), 1502.01719.
- 3213 [49] STAR, M. S. Abdallah *et al.*, Phys. Rev. C **105**, 044906 (2022), 2109.09793.
- 3214 [50] STAR, M. Abdallah *et al.*, Phys. Rev. D **104**, 052007 (2021), 2103.13286.
- 3215 [51] STAR, J. Adam *et al.*, Phys. Lett. B **811**, 135846 (2020), 2003.02114.
- 3216 [52] ALICE, S. Acharya *et al.*, Phys. Lett. B **793**, 420 (2019), 1805.05212.
- 3217 [53] CMS, A. M. Sirunyan *et al.*, Phys. Rev. Lett. **127**, 102002 (2021), 2103.14089.
- 3218 [54] S. Wang, W. Dai, B.-W. Zhang, and E. Wang, Eur. Phys. J. C **79**, 789 (2019),
3219 1906.01499.
- 3220 [55] M. A. *et al.* [STAR], Phys. Lett. B **827** (2022).
- 3221 [56] STAR, M. S. Abdallah *et al.*, (2021), 2112.00240.
- 3222 [57] STAR, M. Abdallah *et al.*, (2021), 2110.09513.
- 3223 [58] STAR, M. S. Abdallah *et al.*, (2021), 2108.00924.
- 3224 [59] STAR, M. S. Abdallah *et al.*, Longitudinal double-spin asymmetry for inclusive jet
3225 and dijet production in polarized proton collisions at $\sqrt{s} = 510$ GeV, 2021, 2110.11020.
- 3226 [60] STAR, M. Abdallah *et al.*, (2021), 2111.10396.
- 3227 [61] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **115**, 092002 (2015).
- 3228 [62] STAR, L. Adamczyk *et al.*, Phys. Rev. D **95**, 071103 (2017).
- 3229 [63] STAR, M. S. Abdallah *et al.*, Phys. Rev. D **103**, L091103 (2021).
- 3230 [64] STAR, J. Adam *et al.*, Phys. Rev. D **100**, 052005 (2019).
- 3231 [65] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, Phys. Rev. Lett. **113**,
3232 012001 (2014).
- 3233 [66] NNPDF, E. R. Nocera, R. D. Ball, S. Forte, G. Ridolfi, and J. Rojo, Nucl. Phys. B
3234 **887**, 276 (2014).
- 3235 [67] STAR, J. Kwasizur, Longitudinal Double-Spin Asymmetries for Dijet Production at
3236 Intermediate Pseudorapidity in Polarized Proton-Proton Collisions at $\sqrt{s} = 510$ GeV,
3237 in *APS Division of Nuclear Physics Meeting*, 2020.

- 3238 [68] STAR, J. Kwasizur, Longitudinal Double-Spin Asymmetries for Dijet Production at
3239 Intermediate Pseudorapidity in Polarized Proton-Proton Collisions at $\sqrt{s} = 510$ GeV,
3240 in *24th International Spin Symposium*, 2021.
- 3241 [69] J. L. Albacete, G. Giacalone, C. Marquet, and M. Matas, Phys. Rev. D **99**, 014002
3242 (2019), 1805.05711.
- 3243 [70] D. Kharzeev, E. Levin, and L. McLerran, Nucl. Phys. A **748**, 627 (2005), hep-
3244 ph/0403271.
- 3245 [71] C. Marquet, Nucl. Phys. A **796**, 41 (2007), 0708.0231.
- 3246 [72] PHENIX, A. Adare *et al.*, Phys. Rev. Lett. **107**, 172301 (2011), 1105.5112.
- 3247 [73] M. Strikman and W. Vogelsang, Phys. Rev. D **83**, 034029 (2011), 1009.6123.
- 3248 [74] U. D’Alesio, F. Murgia, and C. Pisano, Physics Letters B **773**, 300 (2017), 1707.00914.
- 3249 [75] Z.-B. Kang, A. Prokudin, F. Ringer, and F. Yuan, Physics Letters B **774**, 635 (2017),
3250 1707.00913.
- 3251 [76] A. Aprahamian *et al.*, Reaching for the horizon: The 2015 long range plan for
3252 nuclear science, 2015, The 2015 Long Range Plan for Nuclear Science “Reach-
3253 ing for the Horizon” [http://science.energy.gov/~media/np/nsac/pdf/2015LRP/
3254 2015_LRPNS_091815.pdf](http://science.energy.gov/~media/np/nsac/pdf/2015LRP/2015_LRPNS_091815.pdf).
- 3255 [77] RHIC luminosity projection [https://www.rhichome.bnl.gov//RHIC/Runs/
3256 RhicProjections.pdf](https://www.rhichome.bnl.gov//RHIC/Runs/RhicProjections.pdf).
- 3257 [78] SN0619 : A Proposal for STAR Inner TPC Sector Upgrade (iTTPC) [https://drupal.
3258 star.bnl.gov/STAR/starnotes/public/sn0619](https://drupal.star.bnl.gov/STAR/starnotes/public/sn0619).
- 3259 [79] STAR, CBM eTOF Group, The CBM Collaboration eTOF Group, (2016), 1609.05102.
- 3260 [80] The STAR Forward Calorimeter System and Forward Tracking System beyond
3261 BES-II [https://drupal.star.bnl.gov/STAR/files/Proposal.ForwardUpgrade.
3262 Nov_.2018.Review.pdf](https://drupal.star.bnl.gov/STAR/files/Proposal.ForwardUpgrade.Nov_.2018.Review.pdf).
- 3263 [81] J. Jia and P. Huo, Phys. Rev. **C90**, 034915 (2014), 1403.6077.
- 3264 [82] L.-G. Pang, H. Petersen, G.-Y. Qin, V. Roy, and X.-N. Wang, Eur. Phys. J. **A52**, 97
3265 (2016), 1511.04131.
- 3266 [83] W. Li, Nucl. Phys. A **967**, 59 (2017), 1704.03576.
- 3267 [84] CMS, V. Khachatryan *et al.*, Phys. Rev. C **92**, 034911 (2015), 1503.01692.
- 3268 [85] ATLAS, M. Aaboud *et al.*, Eur. Phys. J. C **78**, 142 (2018), 1709.02301.

- 3269 [86] STAR Collaboration, L. Adamczyk *et al.*, Manuscript under preparation (2016).
- 3270 [87] S. K. Das *et al.*, Phys. Lett. B **768**, 260 (2017), 1608.02231.
- 3271 [88] S. Chatterjee and P. Bożek, Phys. Rev. Lett. **120**, 192301 (2018), 1712.01189.
- 3272 [89] B. Chen, M. Hu, H. Zhang, and J. Zhao, Phys. Lett. B **802**, 135271 (2020), 1910.08275.
- 3273 [90] STAR Collaboration, J. Adam *et al.*, Phys. Rev. Lett. **121**, 132301 (2018).
- 3274 [91] ATLAS Collaboration, M. Aaboud *et al.*, Phys. Rev. Lett. **121**, 212301 (2018).
- 3275 [92] J. E. Bernhard, J. S. Moreland, S. A. Bass, J. Liu, and U. Heinz, Phys. Rev. C **94**,
3276 024907 (2016), 1605.03954.
- 3277 [93] JETSCAPE, D. Everett *et al.*, Phys. Rev. C **103**, 054904 (2021), 2011.01430.
- 3278 [94] G. Nijs, W. van der Schee, U. Gürsoy, and R. Snellings, (2020), 2010.15130.
- 3279 [95] G. Denicol, A. Monnai, and B. Schenke, Phys. Rev. Lett. **116**, 212301 (2016),
3280 1512.01538.
- 3281 [96] H. Niemi, G. S. Denicol, P. Huovinen, E. Molnar, and D. H. Rischke, Phys. Rev. C **86**,
3282 014909 (2012), 1203.2452.
- 3283 [97] X. Guo, S. Shi, N. Xu, Z. Xu, and P. Zhuang, Phys. Lett. B **751**, 215 (2015),
3284 1502.04407.
- 3285 [98] B. Chen, X. Du, and R. Rapp, Nucl. Part. Phys. Proc. **289-290**, 475 (2017),
3286 1612.02089.
- 3287 [99] ALICE, J. Adam *et al.*, JHEP **05**, 179 (2016), 1506.08804.
- 3288 [100] ATLAS, M. Aaboud *et al.*, Eur. Phys. J. C **78**, 762 (2018), 1805.04077.
- 3289 [101] CMS, A. M. Sirunyan *et al.*, Phys. Rev. Lett. **118**, 162301 (2017), 1611.01438.
- 3290 [102] STAR, L. Adamczyk *et al.*, Nature **548**, 62 (2017), 1701.06657.
- 3291 [103] STAR, J. Adam *et al.*, Phys. Rev. Lett. **126**, 162301 (2021), 2012.13601.
- 3292 [104] STAR, B. Abelev *et al.*, Phys. Rev. C **76**, 064904 (2007), 0706.0472.
- 3293 [105] Z.-T. Liang, J. Song, I. Upsal, Q. Wang, and Z.-B. Xu, (2019), 1912.10223.
- 3294 [106] Y. Xie, D. Wang, and L. P. Csernai, Eur. Phys. J. C **80**, 39 (2020), 1907.00773.
- 3295 [107] X.-L. Sheng, Q. Wang, and X.-N. Wang, (2020), 2007.05106.
- 3296 [108] STAR, M. Abdallah *et al.*, (2022), 2204.02302.

- 3297 [109] V. Voronyuk *et al.*, Phys. Rev. C **83**, 054911 (2011), 1103.4239.
- 3298 [110] W.-T. Deng and X.-G. Huang, Phys.Rev. **C85**, 044907 (2012), 1201.5108.
- 3299 [111] X.-L. Zhao, G.-L. Ma, and Y.-G. Ma, Phys. Rev. C **99**, 034903 (2019), 1901.04151.
- 3300 [112] X.-L. Zhao, Y.-G. Ma, and G.-L. Ma, Phys. Rev. C **97**, 024910 (2018), 1709.05962.
- 3301 [113] D. E. Kharzeev, L. D. McLerran, and H. J. Warringa, Nucl. Phys. **A803**, 227 (2008),
3302 0711.0950.
- 3303 [114] U. Gürsoy, D. Kharzeev, E. Marcus, K. Rajagopal, and C. Shen, Phys. Rev. C **98**,
3304 055201 (2018), 1806.05288.
- 3305 [115] ALICE, S. Acharya *et al.*, Phys. Rev. Lett. **125**, 022301 (2020), 1910.14406.
- 3306 [116] STAR, J. Adam *et al.*, Phys. Rev. Lett. **123**, 162301 (2019), 1905.02052.
- 3307 [117] STAR, F. Seck, Nucl. Phys. A **1005**, 122005 (2021).
- 3308 [118] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **113**, 022301 (2014), 1312.7397, [Adden-
3309 dum: Phys.Rev.Lett. 113, 049903 (2014)].
- 3310 [119] D. Kharzeev and R. D. Pisarski, Phys.Rev. **D61**, 111901 (2000), hep-ph/9906401.
- 3311 [120] D. Kharzeev, Phys.Lett. **B633**, 260 (2006), hep-ph/0406125.
- 3312 [121] R. Milton *et al.*, Phys. Rev. C **104**, 064906 (2021).
- 3313 [122] L. Finch and S. Murray, Phys. Rev. C **96**, 044911 (2017), 1801.06476.
- 3314 [123] F. Becattini, M. Buzzegoli, A. Palermo, and G. Prokhorov, Phys. Lett. B **822**, 136706
3315 (2021), 2009.13449.
- 3316 [124] S. Cao, G.-Y. Qin, and S. A. Bass, Phys. Rev. C **92**, 024907 (2015).
- 3317 [125] JETSCAPE, A. Kumar *et al.*, Nucl. Phys. A **1005**, 122009 (2021), 2002.07124.
- 3318 [126] Y. Mehtar-Tani and K. Tywoniuk, Phys. Rev. D **98**, 051501 (2018), 1707.07361.
- 3319 [127] Y. Mehtar-Tani and K. Tywoniuk, Nucl. Phys. A **979**, 165 (2018), 1706.06047.
- 3320 [128] B. G. Zakharov, Eur. Phys. J. C **81**, 57 (2021), 2003.10182.
- 3321 [129] L. Chen, G.-Y. Qin, S.-Y. Wei, B.-W. Xiao, and H.-Z. Zhang, Phys. Lett. B **773**, 672
3322 (2017), 1607.01932.
- 3323 [130] A. H. Mueller, B. Wu, B.-W. Xiao, and F. Yuan, Phys. Lett. B **763**, 208 (2016),
3324 1604.04250.

- 3325 [131] STAR, J. Adam *et al.*, (2020), 2006.00582.
- 3326 [132] STAR, N. R. Sahoo, Measurement of γ +jet and π^0 +jet in central Au+Au collisions
3327 at $\sqrt{s_{NN}} = 200$ GeV with the STAR experiment, 2020, 2008.08789.
- 3328 [133] F. D’Eramo, K. Rajagopal, and Y. Yin, JHEP **01**, 172 (2019), 1808.03250.
- 3329 [134] ALICE, J. Adam *et al.*, JHEP **09**, 170 (2015), 1506.03984.
- 3330 [135] STAR, L. Adamczyk *et al.*, Phys. Rev. C **96**, 024905 (2017), 1702.01108.
- 3331 [136] A. Bazavov *et al.*, Phys. Rev. D **95**, 054504 (2017), 1701.04325.
- 3332 [137] S. Borsanyi *et al.*, JHEP **10**, 205 (2018), 1805.04445.
- 3333 [138] CMS, V. Khachatryan *et al.*, Phys. Lett. B **772**, 489 (2017), 1605.06966.
- 3334 [139] ALICE, B. Abelev *et al.*, Phys. Lett. B **718**, 1273 (2013), 1209.3715.
- 3335 [140] STAR, L. Adamczyk *et al.*, Phys. Rev. C **96**, 054904 (2017), 1702.07705.
- 3336 [141] ALICE, S. Acharya *et al.*, JHEP **06**, 035 (2020), 2002.10897.
- 3337 [142] ALICE, S. Acharya *et al.*, (2021), 2101.02581.
- 3338 [143] ALICE, S. Acharya *et al.*, (2021), 2101.04623.
- 3339 [144] ALICE, S. Acharya *et al.*, (2021), 2101.04577.
- 3340 [145] LHCb, R. Aaij *et al.*, (2021), 2108.02681.
- 3341 [146] M. Alvioli, L. Frankfurt, V. Guzey, M. Strikman, and M. Zhalov, CERN Proc. **1**, 151
3342 (2018).
- 3343 [147] V. Guzey and M. Zhalov, JHEP **10**, 207 (2013), 1307.4526.
- 3344 [148] V. Guzey, M. Strikman, and M. Zhalov, Phys. Rev. C **99**, 015201 (2019), 1808.00740.
- 3345 [149] B. Sambasivam, T. Toll, and T. Ullrich, Phys. Lett. B **803**, 135277 (2020), 1910.02899.
- 3346 [150] V. Guzey, Phys. Part. Nucl. Lett. **16**, 498 (2019).
- 3347 [151] V. Guzey and M. Klasen, Phys. Rev. D **104**, 114013 (2021), 2012.13277.
- 3348 [152] V. Guzey and M. Klasen, Inclusive and diffractive dijet photoproduction in ultrape-
3349 ripheral heavy ion collisions at the LHC, in *28th International Workshop on Deep*
3350 *Inelastic Scattering and Related Subjects*, 2021, 2106.16084.
- 3351 [153] W. Chang *et al.*, (2022), 2204.11998.

- 3352 [154] V. Guzey, M. Strikman, and M. Zhalov, *Eur. Phys. J. C* **74**, 2942 (2014), 1312.6486.
- 3353 [155] H. Xing, C. Zhang, J. Zhou, and Y.-J. Zhou, The $\cos 2\phi$ azimuthal asymmetry in ρ^0
3354 meson production in ultraperipheral heavy ion collisions, 2020, 2006.06206.
- 3355 [156] STAR, J. the STAR Collaboration, Adam *et al.*, *Phys. Rev. Lett.* **123**, 132302 (2019),
3356 1904.11658.
- 3357 [157] W. Zha *et al.*, *Phys. Rev. C* **97**, 044910 (2018).
- 3358 [158] M. B. G. Ducati and S. Martins, *Phys. Rev. D* **97**, 116013 (2018).
- 3359 [159] W. Shi, W. Zha, and B. Chen, *Phys. Lett. B* **777**, 399 (2018), 1710.00332.
- 3360 [160] M. R. Whalley, D. Bourilkov, and R. C. Group, The Les Houches accord PDFs
3361 (LHAPDF) and LHAGLUE, in *HERA and the LHC: A Workshop on the Implications*
3362 *of HERA and LHC Physics (Startup Meeting, CERN, 26-27 March 2004; Midterm*
3363 *Meeting, CERN, 11-13 October 2004)*, pp. 575–581, 2005, hep-ph/0508110.
- 3364 [161] M. Cacciari and G. P. Salam, *Phys. Lett. B* **641**, 57 (2006), hep-ph/0512210.
- 3365 [162] X. Ji, F. Yuan, and Y. Zhao, *Phys. Rev. Lett.* **118**, 192004 (2017), 1612.02438.
- 3366 [163] Y. Hatta, Y. Nakagawa, F. Yuan, Y. Zhao, and B. Xiao, *Phys. Rev. D* **95**, 114032
3367 (2017), 1612.02445.
- 3368 [164] S. Bhattacharya, R. Boussarie, and Y. Hatta, *Phys. Rev. Lett.* **128**, 182002 (2022),
3369 2201.08709.
- 3370 [165] Y. Hatta, B.-W. Xiao, and F. Yuan, *Phys. Rev. Lett.* **116**, 202301 (2016), 1601.01585.
- 3371 [166] ZEUS, I. Abt *et al.*, *JHEP* **12**, 102 (2021), 2106.12377.
- 3372 [167] PHENIX, C. Aidala *et al.*, *Nature Phys.* **15**, 214 (2019), 1805.02973.
- 3373 [168] W. Zhao, C. Shen, and B. Schenke, (2022), 2203.06094.
- 3374 [169] ATLAS, G. Aad *et al.*, (2021), 2101.10771.
- 3375 [170] STAR, R. A. Lacey, *Nucl. Phys. A* **1005**, 122041 (2021), 2002.11889.
- 3376 [171] D. Kharzeev, *Phys. Lett. B* **378**, 238 (1996), nucl-th/9602027.
- 3377 [172] BeAGLE: a Tool to Refine Detector Requirements for eA Collisions EIC R&D; Project
3378 eRD17 E. Aschenauer, M. Baker, W. Chang, J. Lee, Z. Tu et al. [https://wiki.bnl.](https://wiki.bnl.gov/eic/index.php/BeAGLE)
3379 [gov/eic/index.php/BeAGLE](https://wiki.bnl.gov/eic/index.php/BeAGLE).
- 3380 [173] Z. Tu *et al.*, *Phys. Lett. B* **811**, 135877 (2020), 2005.14706.

- 3381 [174] A. Accardi *et al.*, Eur. Phys. J. A **52**, 268 (2016), 1212.1701.
- 3382 [175] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **103**, 092009 (2021).
- 3383 [176] STAR Collaboration, J. Adam *et al.*, Phys. Rev. D **103**, 072005 (2021).
- 3384 [177] J.-w. Qiu and G. F. Sterman, Phys. Rev. D **59**, 014004 (1999), hep-ph/9806356.
- 3385 [178] K. Kanazawa, Y. Koike, A. Metz, and D. Pitonyak, Phys. Rev. D **89**, 111501 (2014).
- 3386 [179] L. Bland *et al.*, Physics Letters B **750**, 660 (2015).
- 3387 [180] STAR, H. Liu, Measurement of transverse single-spin asymmetries for dijet production
3388 in polarized p+p collisions at $\sqrt{s} = 200$ GeV at STAR, in *Nuclear Physics Seminar*,
3389 *Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8633/>.
- 3390 [181] M. Boglione *et al.*, Phys. Lett. B **815**, 136135 (2021), 2101.03955.
- 3391 [182] L. Gamberg, Z.-B. Kang, and A. Prokudin, Phys. Rev. Lett. **110**, 232301 (2013).
- 3392 [183] STAR, B. Abelev *et al.*, Phys. Rev. Lett. **99**, 142003 (2007), 0705.4629.
- 3393 [184] J. Collins and J.-W. Qiu, Phys. Rev. D **75**, 114014 (2007).
- 3394 [185] T. C. Rogers and P. J. Mulders, Phys. Rev. D **81**, 094006 (2010).
- 3395 [186] X. Liu, F. Ringer, W. Vogelsang, and F. Yuan, Factorization and its breaking in dijet
3396 single transverse spin asymmetries in *pp* collisions, 2020, 2008.03666.
- 3397 [187] Z.-B. Kang, K. Lee, D. Y. Shao, and J. Terry, The sivers asymmetry in hadronic dijet
3398 production, 2020, 2008.05470.
- 3399 [188] J. P. Ralston and D. E. Soper, Nuclear Physics B **152**, 109 (1979).
- 3400 [189] R. Jaffe and X.-D. Ji, Nucl. Phys. B **375**, 527 (1992).
- 3401 [190] P. Mulders and R. Tangerman, Nucl. Phys. B **461**, 197 (1996), hep-ph/9510301,
3402 [Erratum: Nucl.Phys.B 484, 538–540 (1997)].
- 3403 [191] D. Sivers, Nuovo Cim. C **035N2**, 171 (2012), 1109.2521.
- 3404 [192] C. Alexandrou *et al.*, Phys. Rev. D **98**, 091503 (2018).
- 3405 [193] R. Jaffe and X.-D. Ji, Phys. Rev. Lett. **67**, 552 (1991).
- 3406 [194] Z.-B. Kang, X. Liu, F. Ringer, and H. Xing, JHEP **11**, 068 (2017), 1705.08443.
- 3407 [195] Z.-B. Kang, A. Prokudin, F. Ringer, and F. Yuan, Phys. Lett. B **774**, 635 (2017),
3408 1707.00913.

- 3409 [196] J. C. Collins, S. F. Heppelmann, and G. A. Ladinsky, Nucl. Phys. B **420**, 565 (1994),
3410 hep-ph/9305309.
- 3411 [197] L. Adamczyk *et al.*, Phys. Lett. B **780**, 332–339 (2018).
- 3412 [198] M. Radici and A. Bacchetta, Phys. Rev. Lett. **120**, 192001 (2018), 1802.05212.
- 3413 [199] STAR, J. Adam *et al.*, Phys. Rev. D **98**, 091103(R) (2018), 1808.08000.
- 3414 [200] STAR, Y. Xu, Longitudinal and Transverse Spin Transfer of Λ and $\bar{\Lambda}$ Hyperons in
3415 Polarized $p+p$ Collisions at $\sqrt{s} = 200$ GeV at RHIC-STAR, in *XXVIII International*
3416 *Workshop on Deep-Inelastic Scattering and Related Subjects, Stony Brook University,*
3417 *2021*, <https://indico.bnl.gov/event/9726/contributions/46261/>.
- 3418 [201] F. Yuan, Phys. Rev. Lett. **100**, 032003 (2008).
- 3419 [202] F. Yuan, Phys. Rev. D **77**, 074019 (2008).
- 3420 [203] U. D’Alesio, F. Murgia, and C. Pisano, Phys. Lett. B **773**, 300 (2017), 1707.00914.
- 3421 [204] A. Airapetian *et al.*, Phys. Lett. B **693**, 11–16 (2010).
- 3422 [205] C. Adolph *et al.*, Phys. Lett. B **744**, 250–259 (2015).
- 3423 [206] D. Müller, D. Robaschik, B. Geyer, F.-M. Dittes, and J. Hořejši, Fortsch. Phys. **42**,
3424 101 (1994), hep-ph/9812448.
- 3425 [207] X.-D. Ji, Phys. Rev. Lett. **78**, 610 (1997), hep-ph/9603249.
- 3426 [208] A. Radyushkin, Phys. Lett. B **380**, 417 (1996), hep-ph/9604317.
- 3427 [209] M. Burkardt, Phys. Rev. D **62**, 071503 (2000), hep-ph/0005108, [Erratum: Phys.Rev.D
3428 66, 119903 (2002)].
- 3429 [210] S. Klein and J. Nystrand, Photoproduction of J/ψ and Upsilon in pp and anti-p p
3430 collisions, in *5th Workshop on Small x and Diffractive Physics*, 2003, hep-ph/0310223.
- 3431 [211] S. R. Klein, J. Nystrand, J. Seger, Y. Gorbunov, and J. Butterworth, Comput. Phys.
3432 Commun. **212**, 258 (2017), 1607.03838.
- 3433 [212] J. Lansberg, L. Massacrier, L. Szymanowski, and J. Wagner, Phys. Lett. B **793**, 33
3434 (2019), 1812.04553.
- 3435 [213] HERMES, A. Airapetian *et al.*, Phys. Lett. B **577**, 37 (2003), hep-ex/0307023.
- 3436 [214] HERMES, A. Airapetian *et al.*, Nucl. Phys. B **780**, 1 (2007), 0704.3270.
- 3437 [215] HERMES, A. Airapetian *et al.*, Phys. Lett. B **684**, 114 (2010), 0906.2478.

- 3438 [216] W. Brooks and H. Hakobyan, Nucl. Phys. A **830**, 361C (2009), 0907.4606.
- 3439 [217] NuSea, M. Vasilev *et al.*, Phys. Rev. Lett. **83**, 2304 (1999), hep-ex/9906010.
- 3440 [218] K. J. Eskola, P. Paakkinen, H. Paukkunen, and C. A. Salgado, Epps21: A global qcd
3441 analysis of nuclear pdfs, 2021.
- 3442 [219] PHENIX, S. Adler *et al.*, Phys. Rev. Lett. **98**, 172302 (2007), nucl-ex/0610036.
- 3443 [220] R. Sassot, M. Stratmann, and P. Zurita, Phys. Rev. D **81**, 054001 (2010), 0912.1311.
- 3444 [221] E. Aschenauer *et al.*, (2014), 1409.1633.
- 3445 [222] N. Armesto, H. Paukkunen, J. M. Penín, C. A. Salgado, and P. Zurita, Eur. Phys. J.
3446 C **76**, 218 (2016), 1512.01528.
- 3447 [223] H. Paukkunen, K. J. Eskola, and C. Salgado, Nucl. Phys. A **931**, 331 (2014), 1408.4563.
- 3448 [224] K. J. Eskola, H. Paukkunen, and C. A. Salgado, JHEP **10**, 213 (2013), 1308.6733.
- 3449 [225] H. Paukkunen and P. Zurita, JHEP **12**, 100 (2014), 1402.6623.
- 3450 [226] L. Gribov, E. Levin, and M. Ryskin, Phys. Rept. **100**, 1 (1983).
- 3451 [227] E. Iancu and R. Venugopalan, The Color glass condensate and high-energy scattering
3452 in QCD, in *In *Hwa, R.C. (ed.) et al.: Quark gluon plasma* 249-3363*, 2003, hep-
3453 ph/0303204.
- 3454 [228] H. Weigert, Prog. Part. Nucl. Phys. **55**, 461 (2005), hep-ph/0501087.
- 3455 [229] J. Jalilian-Marian and Y. V. Kovchegov, Prog. Part. Nucl. Phys. **56**, 104 (2006),
3456 hep-ph/0505052.
- 3457 [230] F. Gelis, E. Iancu, J. Jalilian-Marian, and R. Venugopalan, Ann. Rev. Nucl. Part. Sci.
3458 **60**, 463 (2010), 1002.0333.
- 3459 [231] G. Giuliani, H. Zheng, and A. Bonasera, Prog. Part. Nucl. Phys. **76**, 116 (2014),
3460 1311.1811.
- 3461 [232] Y. V. Kovchegov and E. Levin *Quantum chromodynamics at high energy* Vol. 33 (Cam-
3462 bridge University Press, 2012).
- 3463 [233] A. H. Mueller and J.-w. Qiu, Nucl. Phys. B **268**, 427 (1986).
- 3464 [234] L. McLerran and R. Venugopalan, Phys. Rev. D **49**, 2233 (1994).
- 3465 [235] L. McLerran and R. Venugopalan, Phys. Rev. D **49**, 3352 (1994).
- 3466 [236] L. McLerran and R. Venugopalan, Phys. Rev. D **50**, 2225 (1994).

- 3467 [237] Y. V. Kovchegov, Phys. Rev. D **54**, 5463 (1996).
- 3468 [238] Y. V. Kovchegov, Phys. Rev. D **55**, 5445 (1997).
- 3469 [239] J. Jalilian-Marian, A. Kovner, L. McLerran, and H. Weigert, Phys. Rev. D **55**, 5414
3470 (1997).
- 3471 [240] A. H. Mueller, Nucl. Phys. **B415**, 373 (1994).
- 3472 [241] A. H. Mueller and B. Patel, Nucl. Phys. B **425**, 471 (1994), hep-ph/9403256.
- 3473 [242] I. Balitsky, Nucl. Phys. **B463**, 99 (1996), hep-ph/9509348.
- 3474 [243] I. Balitsky, Phys. Rev. D **60**, 014020 (1999), hep-ph/9812311.
- 3475 [244] Y. V. Kovchegov, Phys. Rev. **D60**, 034008 (1999), hep-ph/9901281.
- 3476 [245] Y. V. Kovchegov, Phys. Rev. D **61**, 074018 (2000), hep-ph/9905214.
- 3477 [246] J. Jalilian-Marian, A. Kovner, and H. Weigert, Phys. Rev. D **59**, 014015 (1998),
3478 hep-ph/9709432.
- 3479 [247] J. Jalilian-Marian, A. Kovner, A. Leonidov, and H. Weigert, Phys. Rev. D **59**, 014014
3480 (1998), hep-ph/9706377.
- 3481 [248] E. Iancu, A. Leonidov, and L. D. McLerran, Phys. Lett. B **510**, 133 (2001), hep-
3482 ph/0102009.
- 3483 [249] E. Iancu, A. Leonidov, and L. D. McLerran, Nucl. Phys. **A692**, 583 (2001), hep-
3484 ph/0011241.
- 3485 [250] Y. V. Kovchegov and M. D. Sievert, Nucl. Phys. B **903**, 164 (2016), 1505.01176.
- 3486 [251] CMS, S. Chatrchyan *et al.*, Eur. Phys. J. C **74**, 2951 (2014), 1401.4433.
- 3487 [252] STAR, E. Braidot, Nucl. Phys. A **854**, 168 (2011), 1008.3989.
- 3488 [253] J. L. Albacete and C. Marquet, Phys. Rev. Lett. **105**, 162301 (2010), 1005.4065.
- 3489 [254] Z.-B. Kang, I. Vitev, and H. Xing, Phys. Rev. D **85**, 054024 (2012), 1112.6021.
- 3490 [255] J. Jalilian-Marian and A. H. Rezaeian, Phys. Rev. D **86**, 034016 (2012), 1204.1319.
- 3491 [256] J. L. Albacete and C. Marquet, Nucl. Phys. A **854**, 154 (2011), 1009.3215.
- 3492 [257] K. J. Eskola, H. Paukkunen, and C. A. Salgado, JHEP **07**, 102 (2008), 0802.0139.
- 3493 [258] A. H. Rezaeian, Phys. Rev. D **86**, 094016 (2012), 1209.0478.

- 3494 [259] T. Sjostrand, S. Mrenna, and P. Z. Skands, *Comput. Phys. Commun.* **178**, 852 (2008),
3495 0710.3820.
- 3496 [260] Di-jet production from pythia8.189 is scaled down due to its overestimation of inclusive
3497 π_0 yields compared to those reported by BRAHMS in *phys. rev. lett.* 98 (2007) 252001
3498 and STAR in *phys. rev. lett.* 97 (2006) 152302.
- 3499 [261] T. Kaufmann, A. Mukherjee, and W. Vogelsang, *Phys. Rev. D* **92**, 054015 (2015),
3500 1506.01415, [Erratum: *Phys.Rev.D* 101, 079901 (2020)].
- 3501 [262] D. de Florian, R. Sassot, M. Epele, R. J. Hernández-Pinto, and M. Stratmann, *Phys.*
3502 *Rev. D* **91**, 014035 (2015), 1410.6027.
- 3503 [263] A. Khouaja *et al.*, *Nucl. Phys. A* **780**, 1 (2006).
- 3504 [264] D. de Florian and R. Sassot, *Phys. Rev. D* **69**, 074028 (2004), hep-ph/0311227.
- 3505 [265] D. de Florian, R. Sassot, and M. Stratmann, *Phys. Rev. D* **75**, 114010 (2007), hep-
3506 ph/0703242.
- 3507 [266] D. de Florian, R. Sassot, and M. Stratmann, *Phys. Rev. D* **76**, 074033 (2007),
3508 0707.1506.
- 3509 [267] M. A. Lisa *et al.*, (2021), 2101.10872.
- 3510 [268] J. L. Nagle and W. A. Zajc, *Ann. Rev. Nucl. Part. Sci.* **68**, 211 (2018), 1801.03477.
- 3511 [269] PHENIX, C. Aidala *et al.*, *Phys. Rev. C* **95**, 034910 (2017), 1609.02894.
- 3512 [270] H. Helmholtz, *The London, Edinburgh, and Dublin Philosophical Magazine and Jour-*
3513 *nal of Science* **33**, 485 (1867), <https://doi.org/10.1080/14786446708639824>.
- 3514 [271] C. Shen *et al.*, In preparation.
- 3515 [272] B. Schenke, S. Jeon, and C. Gale, *Phys. Rev.* **C82**, 014903 (2010), 1004.1408.
- 3516 [273] I. Upsal, *Global Polarization of the $\Lambda/\bar{\Lambda}$ system in the STAR BES*, Ph.D. thesis, The
3517 Ohio State University, 2018.
- 3518 [274] Y. B. Ivanov and A. Soldatov, *Phys. Rev. C* **95**, 054915 (2017), 1701.01319.
- 3519 [275] Y. B. Ivanov and A. Soldatov, *Phys. Rev. C* **97**, 044915 (2018), 1803.01525.
- 3520 [276] Y. B. Ivanov, V. Toneev, and A. Soldatov, *Phys. Atom. Nucl.* **83**, 179 (2020),
3521 1910.01332.
- 3522 [277] Y. B. Ivanov, V. Toneev, and A. Soldatov, *J. Phys. Conf. Ser.* **1435**, 012012 (2020).
- 3523 [278] B. Fu, K. Xu, X.-G. Huang, and H. Song, (2020), 2011.03740.

- 3524 [279] M. Baznat, K. Gudima, A. Sorin, and O. Teryaev, Phys. Rev. C **88**, 061901 (2013),
3525 1301.7003.
- 3526 [280] O. Teryaev and R. Usubov, Phys. Rev. C **92**, 014906 (2015).
- 3527 [281] M. I. Baznat, K. K. Gudima, A. S. Sorin, and O. Teryaev, Phys. Rev. C **93**, 031902
3528 (2016), 1507.04652.
- 3529 [282] W.-T. Deng and X.-G. Huang, Phys. Rev. C **93**, 064907 (2016), 1603.06117.
- 3530 [283] D.-X. Wei, W.-T. Deng, and X.-G. Huang, Phys. Rev. C **99**, 014905 (2019), 1810.00151.
- 3531 [284] X.-L. Xia, H. Li, Z.-B. Tang, and Q. Wang, Phys. Rev. C **98**, 024905 (2018),
3532 1803.00867.
- 3533 [285] A. Zinchenko, A. Sorin, O. Teryaev, and M. Baznat, J. Phys. Conf. Ser. **1435**, 012030
3534 (2020).
- 3535 [286] G. Bunce *et al.*, Phys. Rev. Lett. **36**, 1113 (1976).
- 3536 [287] COSY-TOF, F. Hauenstein *et al.*, Eur. Phys. J. A **52**, 337 (2016), 1607.06305.
- 3537 [288] F. Abe *et al.*, Phys. Rev. D **34**, 1950 (1986).
- 3538 [289] B. Lundberg *et al.*, Phys. Rev. D **40**, 3557 (1989).
- 3539 [290] HERA-B, I. Abt *et al.*, Phys. Lett. B **638**, 415 (2006), hep-ex/0603047.
- 3540 [291] HADES, G. Agakishiev *et al.*, Eur. Phys. J. A **50**, 81 (2014), 1404.3014.
- 3541 [292] STAR, M. Abdallah *et al.*, Phys. Rev. C **105**, 014901 (2022), 2109.00131.
- 3542 [293] ALICE Collaboration, Phys. Lett. B **784**, 82 (2018), 1805.01832.
- 3543 [294] CMS, A. M. Sirunyan *et al.*, Phys. Rev. C **100**, 044902 (2019), 1901.07997.
- 3544 [295] ATLAS Collaboration, (2019), 1911.04812.
- 3545 [296] ATLAS, (2022), 2205.00039.
- 3546 [297] B. Bally, M. Bender, G. Giacalone, and V. Somà, Phys. Rev. Lett. **128**, 082301 (2022),
3547 2108.09578.
- 3548 [298] U. W. Heinz and A. Kuhlman, Phys.Rev.Lett. **94**, 132301 (2005), nucl-th/0411054.
- 3549 [299] Q. Y. Shou *et al.*, Phys. Lett. B **749**, 215 (2015), 1409.8375.
- 3550 [300] G. Giacalone, Phys. Rev. Lett. **124**, 202301 (2020), 1910.04673.

- 3551 [301] G. Giacalone, *A matter of shape: seeing the deformation of atomic nuclei at high-*
3552 *energy colliders*, PhD thesis, U. Paris-Saclay, 2020, 2101.00168.
- 3553 [302] G. Giacalone, J. Jia, and C. Zhang, (2021), 2105.01638.
- 3554 [303] H.-j. Xu, H. Li, X. Wang, C. Shen, and F. Wang, *Phys. Lett. B* **819**, 136453 (2021),
3555 2103.05595.
- 3556 [304] S. Raman, C. W. G. Nestor, Jr, and P. Tikkanen, *Atom. Data Nucl. Data Tabl.* **78**, 1
3557 (2001).
- 3558 [305] S. E. Agbemava, A. V. Afanasjev, and P. Ring, *Phys. Rev. C* **93**, 044304 (2016),
3559 1603.03414.
- 3560 [306] P. Möller, A. J. Sierk, T. Ichikawa, and H. Sagawa, *Atom. Data Nucl. Data Tabl.*
3561 **109-110**, 1 (2016), 1508.06294.
- 3562 [307] L. M. Robledo and G. F. Bertsch, *Phys. Rev. C* **84**, 054302 (2011), 1107.3581.
- 3563 [308] S. Hilaire and M. Girod, *Eur. Phys. J. A* **33**, 237 (2007).
- 3564 [309] T. Nakatsukasa, K. Matsuyanagi, M. Matsuo, and K. Yabana, *Rev. Mod. Phys.* **88**,
3565 045004 (2016), 1606.04717.
- 3566 [310] L.-M. Liu *et al.*, (2022), 2203.09924.
- 3567 [311] S. Zhao, H.-j. Xu, Y.-X. Liu, and H. Song, (2022), 2204.02387.
- 3568 [312] J. Jia, S. Huang, and C. Zhang, (2021), 2105.05713.
- 3569 [313] G. Giacalone, *Phys. Rev. C* **99**, 024910 (2019), 1811.03959.
- 3570 [314] Haojie Xu talk, Chunjian Zhang poster, (STAR Collaboration), Constraints on neutron
3571 skin thickness and nuclear deformations using relativistic heavy-ion collisions from
3572 STAR, <https://indico.cern.ch/event/895086/contributions/4724887/>,
3573 <https://indico.cern.ch/event/895086/contributions/4749420/>.
- 3574 [315] G. Nijs and W. van der Schee, (2021), 2112.13771.
- 3575 [316] J. M. Lattimer and M. Prakash, *Phys. Rept.* **442**, 109 (2007), astro-ph/0612440.
- 3576 [317] B.-A. Li, B.-J. Cai, W.-J. Xie, and N.-B. Zhang, *Universe* **7**, 182 (2021), 2105.04629.
- 3577 [318] PREX, D. Adhikari *et al.*, *Phys. Rev. Lett.* **126**, 172502 (2021), 2102.10767.
- 3578 [319] ALICE, S. Acharya *et al.*, (2021), 2111.06106.
- 3579 [320] G. Giacalone, *Phys. Rev. C* **102**, 024901 (2020), 2004.14463.

- 3580 [321] G. Giacalone, F. G. Gardim, J. Noronha-Hostler, and J.-Y. Ollitrault, Phys. Rev. C
3581 **103**, 024910 (2021), 2004.09799.
- 3582 [322] Chart of Nuclear Quadrupole Deformations, <https://www.nndc.bnl.gov/nudat2/>.
- 3583 [323] P. A. Butler, J. Phys. G **43**, 073002 (2016).
- 3584 [324] C. J. Lister and J. Butterworth, Nature **497**, 190 (2013).